# **ACIS Optical Sensitivity**

## Scott J. Wolk SAO/CXC

# **OBF** Design

Optical blocking on both ACIS and HRC is achieved with a polyimide coated on both sides with a thin layer of aluminum

- ACIS-I AI/Polyimide/AI 1200Å 2000Å 400Å
- ACIS-S AI/Polyimide/AI 1000Å 2000Å 300Å
- Calibration was performed via analytical methods.
  - Results of the analysis where compared to non-flight filter.



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# **OBF** Design

Door in open position

Internal Cal & LED Location Spectroscopy Array FP Alignment Mirror

Fie

Frame Store Shield (3 plc)

DH Alignment Mirrors (2plc)

FID Light Mounting Locations (6 plc)

#### Imaging Array

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# **OBF** Design





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# Identification of problem



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# Identification of problem



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## **Quantification of Problem**



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# **Quantification of Problem**

- The problem was traced to an interference effect which enhanced transmission at longer optical wavelengths.
- This conclusion was confirmed by analysis of flight-like OBF spare.
- Stellar magnitude required to produce 1ADU in a 3.3 second ACIS frametime.

Stellar Temperature	BI Chip in S array	FI Chip in I array (V-Magnitude)	
(K)	(V-Magnitude)		
4000	8.1	2.87	
5000	7.93	2.44	
6500	7.73	1.79	
10000	7.66	1.17	
20000	7.6	0.97	

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# Amelioration -I

# ACIS-I -Mars Shorter exposures LETG - Venus

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12000 km



# **Amelioration -II**



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## Jupiter Cal. Observations - AO2



## Jupiter Cal. Observations - AO2



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## Jupiter Cal. Observations - AO2

Detect Threshold= 20 Split Threshold = 13

Event 1

2	19	7
8	35	5
16	3	9

ACLSGRADE=65 ASCA GRADE = 2 PHA=70 Enetgy ~ 329 eV Detect Threshold= 36 Split Threshold = 29

2	19	7
8	35	5
16	3	9

Non-Detection

 3
 42
 0

 +
 50
 0

 28
 2
 +

ACLSGRADE=64 ASCA GRADE = 2 PHA= 92 Energy ~ 410 eV

Event 2



ACLSGRADE=65 ASCA GRADE = 2 PHA= 120 Energy ~ 514 eV

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