Chandra Aspect

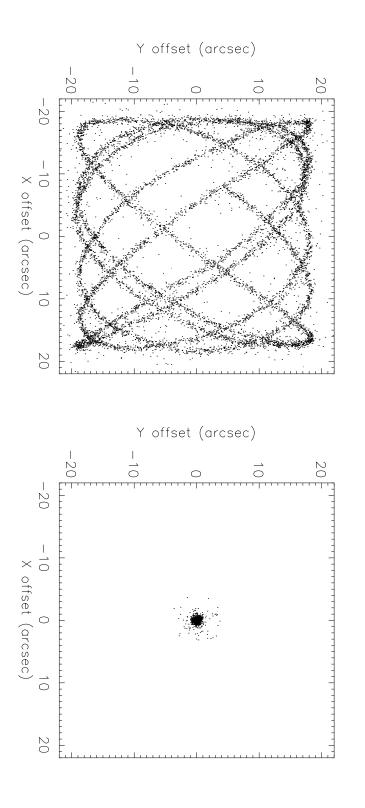
Tom Aldcroft (SAO/CXC)

With contributions from:

Dan Schwartz Mark Cresitello-Dittmar David Morris Rob Cameron Maxim Markevitch Margarita Karovska TRW/FOT Brett Unks

Chandra Aspect

- Aspect solution is a time history of the exact pointing attitude and spacecraft alignment
- sition (RA,Dec), as well as contruction of exposure maps Allows conversion from detector pixel coordinate to sky po-



Resources

Aspect chapter of Proposers Observatory Guide — Description of hardware, aspect processing and products, and operations

Aspect Information page — http://cxc.harvard.edu/cal/ASPECT/ pect performance Contains latest information on caveats, calibration, and as-

Caveats - http://cxc.harvard.edu/cal/ASPECT/aspect_caveats.html

Help desk





CXC Search

Chandra Aspect

Absolute astrometry

Current absolute astrometric accuracy

Thread to detect and fix aspect offsets

High-precision astrometry and image reconstruction with Chandra

Plate scale and relative chip (plate) positions

General aspect help / information

Align event files

ACIS pile-up mimicking bad aspect

Image reconstruction performance

CXC aspect processing caveats

Chandra absolute pointing accuracy

Papers

| II | 1 | Postscript PDF Poster |
|----|---|--------------------------|
| II | Initial performance of the aspect system on the Chandra Observatory: Post-facto aspect reconstruction | Postscript PDF |
| | Initial performance of the attitude control and aspect determination sub-systems on the Chandra Observatory | Postscript PDF |

Technical notes and Talks

CIAO workshop aspect viewgraphs

Chandra Calibration Review: Aspect

CXC internal aspect page

Calibration and Performance

Image reconstruction

```
(http://cxc.harvard.edu/cal/ASPECT/img_recon/report.html)
```

to a gaussian sigma of less than 0.1 arcsec. struction introduces an almost negligible blurring, equivalent Measures the effective blurring of the X-ray PSF due aspect reconstruction. Latest analysis shows aspect reconto

Celestial location

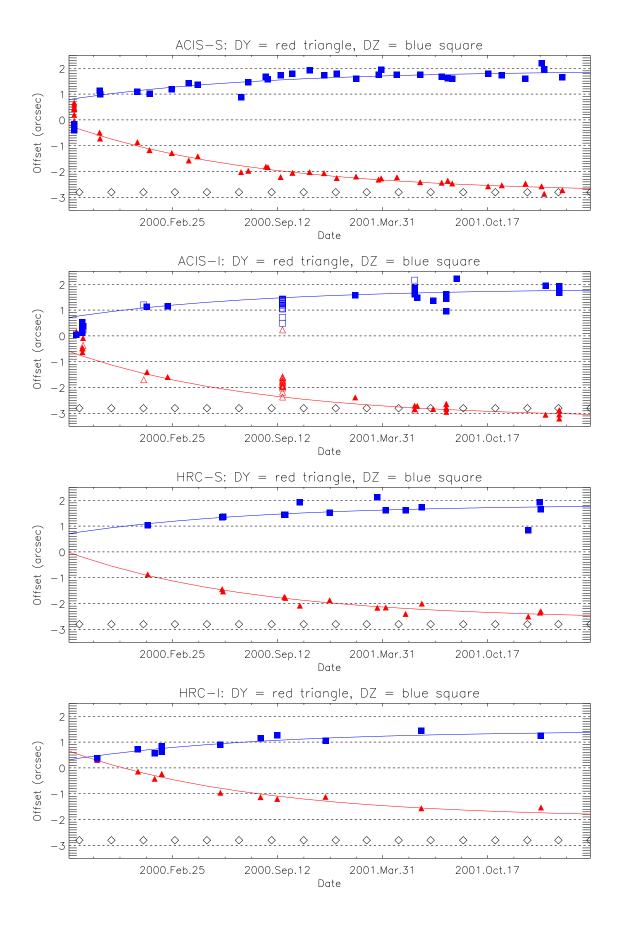
```
(http://cxc.harvard.edu/cal/ASPECT/celmon/)
```

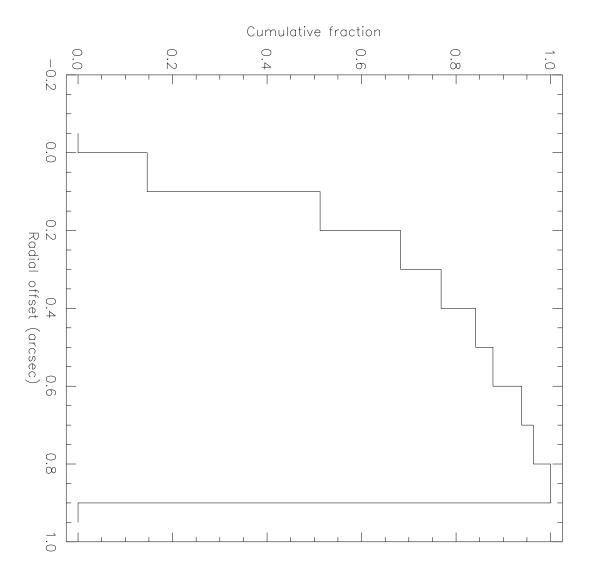
a radius of 0.6 arcsec known coordinates, the source location error circle (90%) has tions. Based on observations of point sources with accurately Measures absolute accuracy of Chandra X-ray source loca-

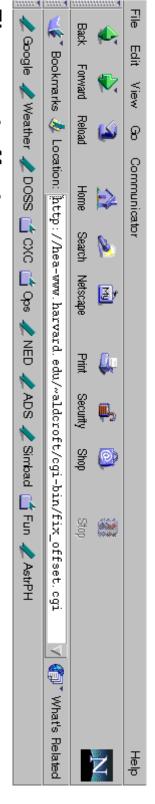
thread: been developed to easily correct these offsets. CAVEAT: Offsets exist in some observations. http://cxc.harvard.edu/ciao/threads/arcsec_correction/ A tool has See CIAO

Celestial location calibration

- A select set of observations were reprocessed with alignment files from early in the mission.
- Moderately bright (but not piled) point sources with accurate optical astrometry.
- Observed at the nominal SIM-Z position
- within 2 arcmin of the HRMA optical axis, with most being observed on-axis
- Your mileage may vary:
- Observations at large off-nominal SIM-Z can suffer additional residual aspect offsets of up to 1 arcsec
- Sources at HRMA off-axis angles greater than ~ 3 arcmin relative chip (plate) positions" page) metry, and plate scale uncertainties (see "Plate scale and may have additional offsets due to PSF blurring and asym-







Fix aspect offset

- Copy your event file (evt_file) to what will be the new aspect-corrected version
- o If the file is unzipped already

 cp evt_file aspcorr_evt2.fits

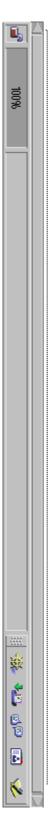
 o If the file is zipped

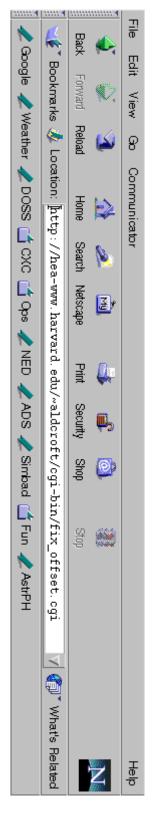
- gunzip --stdout evt_file > aspcorr_evt2.fits
 o If you only want the central 1.0 arcminute
 dmcopy "evt_file[events][sky=circle(4096,4096,1.0')" aspcorr_evt2.fits
 In the text box below, cut and paste below the results of doing:

dmlist aspcorr_evt2.fits header,raw,clean | egrep '(TCTYP|TCRVL|ASCDSVER|OBS_ID)'

```
OBS_ID
                                                                                         ASCDSVER
                                                                     TCTYP11
ICRVL6
         TCTYP6
                   ICRVL5
                             TCTYP5
                                       TCRVL12
                                                 TCTYP12
                                                          TCRVL11
         CPCY
                                                                                        R4CU5UPD5.1
                             CPCX
                                                 DEC--TAN
                                                                    RA---TAN
                                                           52.27564786
                                       31.32790157
                                                                               Observation id
                                                                                         ASCDS version number
                                      [degrees]
                                                          [degrees]
```







Fix aspect offset

Cut and paste the following commands to fix the event file aspcorr_evt2.fits

```
punlearn dmhedit
dmhedit infile=a
                                                 dmhedit
                                                                                                                  dmhedit
                                                                                                                                                                               dmhedit
infile=aspcorr_evt2.fits filelist=none operation=add unit=degrees key=TCRVL11 value=189.211101504764 infile=aspcorr_evt2.fits filelist=none operation=add unit=degrees key=TCRVL12 value=62.22975563905 infile=aspcorr_evt2.fits filelist=none operation=add unit=degrees key=RA_NOM_value=189.211101504764 infile=aspcorr_evt2.fits filelist=none operation=add unit=degrees key=DEC_NOM_value=62.22975563905
```



Improving absolute astrometry

- Improved celestial location precision is possible for some obhigh-precision optical, IR, or radio catalogs servations by cross-correlating detected X-ray sources with
- This technique has been used to achieve absolute astromsec (Orion Nebula cluster). field), \pm /-0.15 arcsec (Hubble Deep Field), and \pm /-0.1 arcetry accurate to $\pm/-0.3$ arcsec (90% confidence, Sgr A*
- Details available in CIAO thread:

http://cxc.harvard.edu/ciao/threads/arcsec_correction/