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# Pileup and the ACIS/HRMA Point Spread Function

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Smithsonian Astrophysical Observatory

Chandra Calibration Workshop 2007 Oct 25

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Summar

# Introduction Pileup and the HRMA/ACIS PSF

- Overview of ACIS operation
- Effects of Pileup
- Pileup Observations (isolated pulsar PSR J0437-4715)
- Spectral fits: pileup " $\alpha$ " model
- PSF profile with pileup
  - depression of core
  - broadening of the PSF
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#### **PSF Pileup**

ACIS Operation (cartoon)

- $\circ$  frame exposed:  $au_{\mathsf{frame}}$  (typically  $\sim$ 3 s)
  - ∘ X-ray photons  $\Rightarrow$  e<sup>-</sup> charge clouds
- 3 × 3 detect island:
  - $\circ$  energy  $\propto$  charge in island
  - charge distribution: "grade"; good=0,2,3,4,6, bad=1.3,7
     "good grade" ⇒ likely X-ray
     "bad grade" ⇒ likely not X-ray
- frame rapidly transferred to frame store
  - $\circ$  parallel transfer rate  $\equiv$  40 $\mu$ s per parallel transfer
  - each pixel of column sees short exposure of all emission along the column ⇒ transfer streak or smear.
  - o transfer ≪ dither; "instantaneous smeared snapshot"

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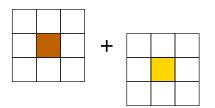
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### HRMA/ACIS PSF Pileup

- pileup: 2+ photons/detect island/frame (overlapping e<sup>-</sup> clouds)
- Pileup has several effects:
  - Spectral distortion + photon loss:
    - 2 photons ⇒ 1 event with higher energy
  - Grade distortion:
    - merging of charge clouds; "good" (likely X-ray) events convert to "bad" (likely non-X-rays)
  - Spatial distortion:
    - o merging of events and grade morphing
      - $\Rightarrow$  loss of events
    - most severe in PSF core, ⇒ PSF shape change





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Pileup PSR JN437-4715

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## **PSF Pileup**

Pileup Observations - PSR J0437-4715

- $\circ$  PSR J0437-4715: isolated pulsar;  $\sim$  0.2 ct/s
- Aim: make data similar to data with different count rate.
- Use different frame times to sample pileup rates:
  - o 0.2s, 0.4s, 1.5s, 3s, 6s, 10s
- o scale dither periods inversely with  $\tau_{\text{frame}}/3 \text{ sec}$ 
  - max dither velocities comparable
  - reduce additional dither "blur"
  - Operational limits (dither rate) for shortest frame times

Upshot: observed  $\sim$  16–24 ks for each frame time

#### Spectral Pileup Model Fits

# **PSF Pileup**

Spectral Pileup Model Fits

Based on Davis (2001); see also The Chandra ABC Guide to Pileup.

- spectral extraction  $r \ge 2''$ : 95+% of point source flux
- $\circ$   $\alpha$ : probability that adding a photon to an event yields a valid X-ray event
- Approach: simultaneous fit of PSR J0437-4715 data
  - Fit "α-model" using XSPEC. Parameters "tied". tframe parameter set according to observation  $\tau_{frame}$ .
  - examined parameters vs. extraction radius, 2–20 pixel radius.
  - plausible parameters:  $\alpha \approx 0.66$ .

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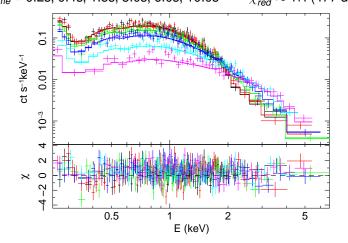
Pileup and E0

### **PSF** Pileup

#### Pileup Model Spectral Fits

Simultaneous fit to pileup observations using pileup model

$$\tau_{frame}$$
 = 0.2s, 0.4s, 1.5s, 3.0s, 6.0s, 10.0s  $\chi_{red}^2 \approx 1.1$  (477 dof)



Pileup and the ACIS/HRMA Point Spread **Function** 

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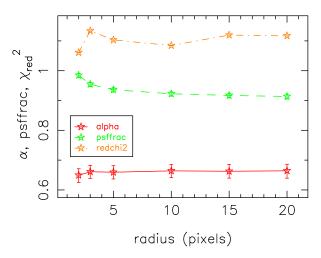
Pileup Model Fit

#### Parameters

### **PSF Pileup**

Pile Model Spectral Fit Parameters

Variation of fit parameters with extraction region radius.

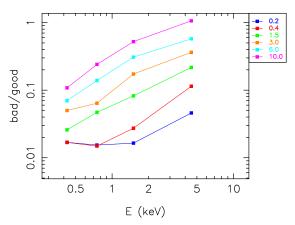


Pileup Grade Ratios

#### **PSF Pileup**

Bad/Good fraction vs.  $\tau_{frame}$ 

- Pileup causes grade morphing
- Examine grade ratios for PSF core
- Bad/Good grade fraction for central 1 pixel radius



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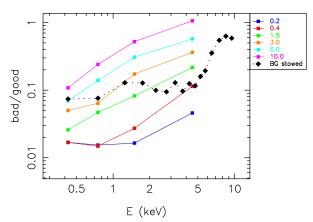
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### **PSF Pileup**

Bad/Good fraction vs.  $\tau_{frame}$ 

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#### **PSF Pileup**

#### Effects of Pileup on PSF Profile

- Loss of events:
  - o merging of events ⇒ event loss + spectral distortion
  - morphing of grades & event loss
- o most severe in core ⇒ depression of core
- Pileup  $\alpha$ -model for spectrum:
  - $\Lambda$  = mean rate/frame (i.e., Poisson mean)
  - $\alpha \sim$  prob. adding photon to event
    - ⇒ still have valid X-ray event
  - fraction of expected count rate lost due to pileup

$$f_r = 1 - \frac{\left[\exp(\alpha\Lambda) - 1\right] \exp(\Lambda)}{\alpha\Lambda} \approx 1 - \left(1 - \frac{\alpha}{2}\right)\Lambda$$

rate loss linear in Λ for small/modest pileup

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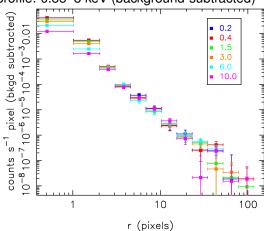
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#### PSF Pileup Profiles

PSF profile: 0.35–8 keV (background-subtracted)



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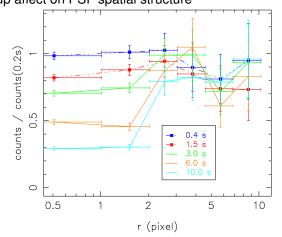
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## **PSF** Pileup

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PSF profile: ratio to  $\tau_{frame} = 0.2s$  (0.35–8 keV) Pileup affect on PSF spatial structure



PSF Relative Profile

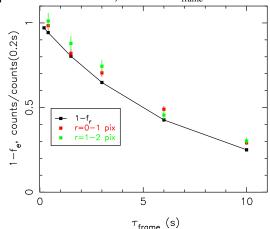
Pileup Losses

## **PSF Pileup**

Pileup Losses

PSF profile: losses in the PSF core

Pileup model:  $\alpha = 0.662$ ,  $\Lambda = 0.22 \tau_{frame}$ 



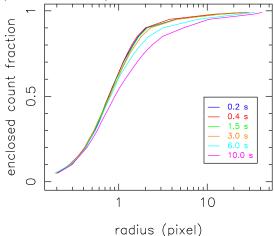
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#### **Enclosed Count Fractions**

**Profile Broadening** 

#### Pileup broadens ECF profile



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#### Summary

- Examined a range of pileup conditions for a single source (PSR J0437-4715)
- o compared to spectral " $\alpha$ -model" for pileup  $\alpha \approx$  0.66
- examined grade migration ("bad grades"/"good grades")
   vs. rate and radius.
- examined PSF profile;
   depression of PSF core due to pileup
  - ⇒ broadening of PSF