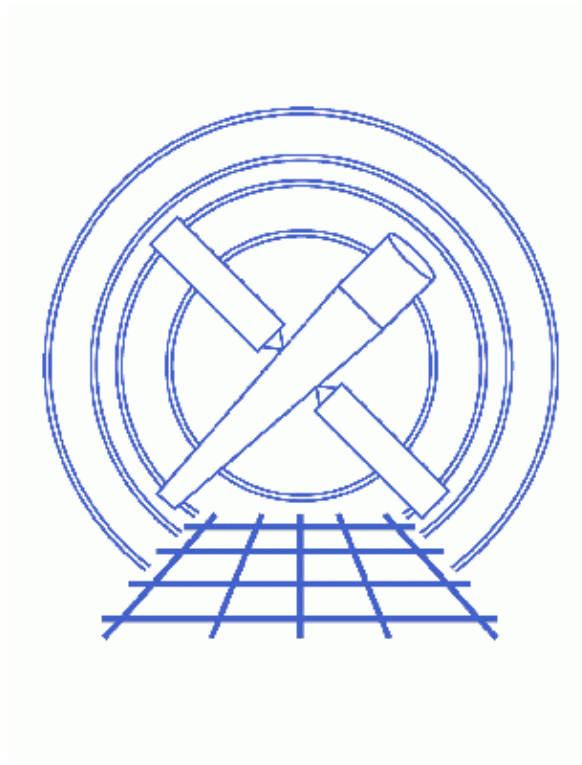


Introduction to the Data Model



CIAO 3.4 Science Threads

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Introduction to the Data Model

CIAO 3.4 Science Threads

Overview

Last Update: 1 Dec 2006 – updated for CIAO 3.4: updates to screen output in [dmgti example](#)

Synopsis:

The CIAO Data Model is a versatile interface used to examine standard format datafiles (FITS, IMH); it also allows powerful filtering of datafiles. The filtered output can either be saved or input directly to an analysis task as a "virtual file." A brief introduction to the Data Model interface and examples on how to use the four core tools are provided here.

Related Links:

- Full details on all the dmtools can be found in their respective [ahelp files](#).

Proceed to the [HTML](#) or [hardcopy \(PDF: A4 / letter\)](#) version of the thread.

Get Started

For illustration, this thread utilizes the ObsID 1843 (ACIS-I, G21.5–0.9) data that was downloaded in the [How to Download Chandra Data from the Archive](#) thread.

If this is your first time using CIAO, please read the [Starting CIAO](#) thread to ensure that your environment is configured properly.

Data Model Tools

The four most frequently used (aka "core") tools are:

- ***dm^list*** – List contents or structure of a file
- ***dm^copy*** – Filter and bin tables and images
- ***dm^extract*** – Make a histogram table file (e.g. PHA or PI file) for spectral fitting or a radial profile for spatial fitting.
- ***dm^gti*** – Define custom good time intervals (GTIs)

In addition, there are a number of more specialized tools:

- ***dm^append*** – Append multiple blocks/extensions to an existing output file

- ***dmarfadd*** – Add multiple ARF files together, weighting by exposure time
- ***dmcontour*** – Make contour regions from a 2–D image
- ***dmcoords*** – Convert between Chandra instrumental coordinate systems
- ***dmdiff*** – Compare values in two FITS files
- ***dmfilth*** – Replace pixel values in source regions of an image with interpolated values
- ***dmgroup*** – Group a specified column in a table
- ***dmgroupreg*** – Translates DS9 regions and groups to CIAO format
- ***dmhedit*** – Edit Data Model file headers
- ***dmhistory*** – Extracts processing history from file header
- ***dmimg2jpg*** – Make a color JPEG image from three image files
- ***dmimgcalc*** – Perform arithmetic on images
- ***dmimghist*** – Make histogram of values in a 2–D image
- ***dmimgpick*** – Maps image values to rows in a table
- ***dmimgthresh*** – Set low exposure regions of an image to zero
- ***dmjoin*** – Join two files whose sampling is different
- ***dmkeypar*** – Retrieve information about a keyword from an input file
- ***dmmakepar*** – Write header keywords to a parameter file
- ***dmmakereg*** – Create a FITS region file from an ASCII region description
- ***dmmerge*** – Merge two or more compatible tables into one
- ***dmpaste*** – Add new columns to a table
- ***dmreadpar*** – Add parameters from a .par file to a file header
- ***dmregrid*** – Rebin a stack of two–dimensional images
- ***dmselect*** – Sort a table block on a given column
- ***dmstat*** – Compute standard statistics for the column in a table or image
- ***dmtcalc*** – Define new table columns as functions of old ones
- ***dmtree2split*** – Create a type 1 file for specified rows of a type 2 file

Running Data Model Tools

Tools are usually executed from the command–line prompt:

```
unix% dmcopy "acisf01843N001_evt2.fits[EVENTS]" acis_events.fits
```

Alternatively, the parameters can be input to the parameter file before the tool is run:

```
unix% punlearn dmcopy
unix% pset dmcopy infile=acisf01843N001_evt2.fits
unix% pset dmcopy outfile=acis_events.fits
unix% dmcopy
Input dataset/block specification (acisf01843N001_evt2.fits):
Output dataset name (acis_events.fits):
```

The parameter file that was used can be examined with plist dmcopy. Note that the Data Model automatically uses the "interesting" block (e.g. EVENTS for an event file, SPECTRUM for a PHA file) when one is not provided, so we chose not to specify it in the second example.

All of the tool capabilities are also available from a GUI (peg) accessed via the Analysis menu of *Prism* and *filtwin*; see the Introduction to the Analysis Menu and the Introduction to peg threads for more information.

Virtual Files

The Data Model offers an easy and powerful means of filtering data. As mentioned, the filtered file can be directly input to an analysis task without writing it to disk first; this is known as a "virtual file." The virtual file, which can also be referred to as a subspace, is simply a means of defining a subset of interest in the dataset.

The basic syntax of a virtual file has four arguments:

```
filename[block][filter][columns/binning][newblock]
```

- `block` – the section of the file to use
- `filter` – the filter to be applied
- `columns/binning` – either the columns from a table to be included in an output table or the binning
- `newblock` (optional) – the name for the new block in the output file, default is the block used from the input file

See [ahelp dmfiltering](#) for more information on filtering syntax.

Examples of virtual files:

- Select the first three columns of the EVENTS block by number:

```
acisf01843N001_evt2.fits[EVENTS][cols #1,#2,#3]
```

or by name:

```
acisf01843N001_evt2.fits[cols time,ccd_id,node_id]
```

- Select certain rows from a FITS file:

```
acisf01843N001_evt2.fits[#row=1:4]
```

- Bin an event file to create a PI spectrum (using [dmextract](#)):

```
acisf01843N001_evt2.fits[EVENTS][bin pi=1:1024:1]
```

or an image (using [dmcopy](#)):

```
acisf01843N001_evt2.fits[EVENTS][bin x=3200:4800:4,y=3200:4800:4]
```

More information on binning data is available from [ahelp dmbinning](#).

The Core Tools

These examples show just a few of the uses for each tool – see the [ahelp](#) files for more options and examples.

1. `dmclist`

[FITS](#) files generally contain several different blocks, each containing multi-dimensional data. The `dmclist` tool allows the user to inspect all or part of a data file by column, row, etc.

- To list the blocks in a file:

```
unix% dmclist acisf01843N001_evt2.fits blocks
```

```
-----  
Dataset: acisf01843N001_evt2.fits  
-----
```


Block	1: PRIMARY	Null		
Block	2: EVENTS	Table	14 cols x 475869	rows
Block	3: GTI7	Table	2 cols x 1	rows
Block	4: GTI0	Table	2 cols x 1	rows
Block	5: GTI1	Table	2 cols x 1	rows
Block	6: GTI2	Table	2 cols x 1	rows
Block	7: GTI3	Table	2 cols x 2	rows
Block	8: GTI6	Table	2 cols x 1	rows

2. dmcoppy

The `dmcoppy` tool can be used to manipulate data. Unlike `dmlist`, which produces text output, this tool produces a new data file in one of the supported formats.

- Copy the events from the central region of a file into a new FITS file:

```
unix% dmcoppy "acisf01843N001_evt2.fits[events][x=3600:4000,y=3800:4200]" \
  acis_center.fits
```

- Generate a blocked image:

A convenient way to display the full field of view of an event file is to bin it into an image. In this example, the full range of sky coordinates is blocked by a factor of 4:

```
unix% dmcoppy "acisf01843N001_evt2.fits[events][bin x>:::4,y>:::4][IMAGE]" \
  acis_img.fits
```

Since the image file is a binned version of the event file, only the selected variable (i.e. sky coordinates) is retained. All other information (photon arrival times, energy, etc.) is lost.

- Filter using a region defined in sky coordinates:

```
unix% dmcoppy "acisf01843N001_evt2.fits[sky=ellipse(1628,4116,92,160,0)]" \
  source.fits
```

Note that the name of the column in the filter must match the name within the file itself; in Chandra data, "sky" is shorthand for "(x, y)."

3. dmextract

The `dmextract` tool is similar to `dmcoppy`. It is used to bin tables into images, but writes the binned data to a table instead of creating an image.

- Make a histogram from a table column:

```
unix% dmextract "acisf01843N001_evt2.fits[bin pha=1:2048:2]" histogram.pha
```

- Extract the PI spectrum of a source in sky coordinates:

```
unix% dmextract "acisf01843N001_evt2.fits[sky=region(ds9.reg)][bin pi]" \
  spectrum.fits
```

where

```
unix% more ds9.reg
# Region file format: CIAO version 1.0
ellipse(1628,4116,92,172,0)
```

More examples of region syntax are available from [ahelp dmregions](#).

- Use a stack input:

This tool can take a `stack` as input, output, or both. To combine both of the previous examples into one command:

```
unix% more infiles.lis
acisf01843N001_evt2.fits[bin pha=1:2048:2]
```



```

acisf01843N001_evt2.fits[sky=ellipse(1628,4116,92,172,0)][bin pi]

unix% more outfiles.lis
histogram_lis.pha
spectrum_lis.fits

unix% dmextract infile=@infiles.lis outfile=@outfiles.lis

```

4. dmgti

Pipeline processing of Chandra data uses `dmgti` to calculate good time intervals (GTIs) based on input MTL files. Although this is what the tool was designed for, `dmgti` will work equally well on any GTI filter as long as the first column of the input file is time.

A look at a lightcurve file (created with `dmextract` and [these parameters](#)) shows background flares where the count rate reaches values much higher than the mean (several rows were omitted for the sake of space):

```

unix% dmlist "bkg_lc.fits[cols time,count_rate,stat_err]" data
-----
Data for Table Block LIGHTCURVE
-----

```

ROW	TIME	COUNT_RATE	STAT_ERR
1	84270903.7887200117		0
2	84270907.0297600031		0
3	84270910.2707999945		0
.			
.			
490	84272488.6572799981	7.8124999849	5.0
491	84272491.8983199894	6.8749999867	4.6904157598
492	84272495.1393600106	10.6249999794	5.8309518948
493	84272498.3804000020	7.8124999849	5.0
.			
.			
742	84273305.3993600011	83.7499998377	16.3707055437
743	84273308.6403999925	121.8749997639	19.7484176581
744	84273311.8814400136	84.0624998371	16.4012194669
745	84273315.1224800050	97.4999998111	17.6635217327
746	84273318.3635199964	65.6249998728	14.4913767462
.			
.			
3005	84280639.8728800118		0
3006	84280643.1139200032		0
3007	84280646.3549599946		0

The tool `dmsort` is used to sort the `count_rate` column in descending order, to show more clearly the times of high background:

```

unix% dmsort bkg_lc.fits sorted_bkg_lc.fits keys=-count_rate copyall=yes

unix% dmlist "sorted_bkg_lc.fits[cols time,count_rate,stat_err]" data | more
-----
Data for Table Block LIGHTCURVE
-----

```

ROW	TIME	COUNT_RATE	STAT_ERR
1	84274261.5061599910	179.6874996518	23.9791576166

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2	84274245.3009600043	177.4999996561	23.8327505756
3	84274248.5419999957	168.4374996736	23.2163735325
4	84274235.5778400004	164.9999996803	22.9782505862
.			
.			
453	84280056.4856800139	77.1874998504	15.7162336455
454	84280046.762560010	77.1874998504	15.7162336455
455	84279926.8440800011	77.1874998504	15.7162336455
.			
.			
2322	84272702.5659199953	8.1249999843	5.0990195136
2323	84272634.5040799975	8.1249999843	5.0990195136
2324	84272997.5005600154	7.8124999849	5.0
.			
.			
3005	84270910.2707999945	0	0
3006	84270907.0297600031	0	0
3007	84270903.7887200117	0	0

If we wanted to exclude the times when the count rate was less than 40.0 and greater than 100.0:

```

unix% dmgti infile=bkg_lc.fits outfile=acis_gti.fits \
  userlimit="(count_rate>40.0)&&(count_rate<100.0)"

unix% dmlist "acis_gti.fits[gti]" data
-----
Data for Table Block GTI
-----

```

ROW	START	STOP
1	84273206.5476399958	84273209.7886800170
2	84273213.0297200084	84273232.4759600163
3	84273242.1990800202	84273307.0198799968
.		
.		
72	84279795.5819600224	84279798.8230000138
73	84279818.2692400217	84279857.1617200077
74	84279860.4027599990	84280443.7899599969

The event file may now be filtered on the newly calculated GTI:

```
unix% dmcopu "acisf01843N001_evt2.fits[@acis_gti.fits]" acis_filtered_evt2.fits
```

Doing a dmlist on each of the files shows how the number of events and the GTI blocks are affected by this filter:

```

unix% dmlist acisf01843N001_evt2.fits blocks
-----
Dataset: acisf01843N001_evt2.fits
-----

```

Block Name	Type	Dimensions
Block 1: PRIMARY	Null	
Block 2: EVENTS	Table	14 cols x 475869 rows
Block 3: GTI7	Table	2 cols x 1 rows
Block 4: GTI0	Table	2 cols x 1 rows
Block 5: GTI1	Table	2 cols x 1 rows
Block 6: GTI2	Table	2 cols x 1 rows
Block 7: GTI3	Table	2 cols x 2 rows
Block 8: GTI6	Table	2 cols x 1 rows

```

unix% dmlist acis_filtered_evt2.fits blocks

```

Dataset: acis_filtered_evt2.fits				
Block Name	Type	Dimensions		
Block 1: PRIMARY	Null			
Block 2: EVENTS	Table	14 cols x 407991	rows	
Block 3: GTI7	Table	2 cols x 74	rows	
Block 4: GTI0	Table	2 cols x 74	rows	
Block 5: GTI1	Table	2 cols x 74	rows	
Block 6: GTI2	Table	2 cols x 74	rows	
Block 7: GTI3	Table	2 cols x 75	rows	
Block 8: GTI6	Table	2 cols x 74	rows	

Parameters for /home/username/cxcds_param/dmcopy.par

```

infile = acisf01843N001_evt2.fits Input dataset/block specification
outfile = acis_events.fits Output dataset name
(kernel = default)      Output file format type
(option = )             Option - force output type
(verbose = 0)           Debug Level
(clobber = no)         Clobber existing file
(mode = ql)

```

Parameters for /home/username/cxcds_param/dmextract.par

```

infile = acisf01843N001_evt2.fits[(x,y)=field()-ellipse(1628,4116,100,140,0)][bin time=:
outfile = bkg_lc.fits      Enter output file name
(bkg = )                  Background region file or fixed background (counts/pixel/s) sub
(error = gaussian)        Method for error determination(gaussian|gehrels|<variance file>
(bkgerror = gaussian)     Method for background error determination(gaussian|gehrels|<var
(bkgnorm = 1.0)           Background normalization
(exp = )                  Exposure map image file
(bkgexp = )               Background exposure map image file
(sys_err = 0)             Fixed systematic error value for SYS_ERR keyword
(opt = ltcl)              Output file type
(defaults = ${ASCDS_CALIB}/cxo.mdb -> /soft/ciao/data/cxo.mdb) Instrument defaults file
(wmap = )                 WMAP filter/binning (e.g. det=8 or default)
(clobber = no)            OK to overwrite existing output file(s)?
(verbose = 0)             Verbosity level
(mode = ql)

```

History

- 23 Dec 2004 reviewed for CIAO 3.2: no changes
- 01 Dec 2005 updated for CIAO 3.3: default value of dmextract error and bkgerror parameters is "gaussian"; updates to screen output in [dmgti example](#)
- 01 Dec 2006 updated for CIAO 3.4: updates to screen output in [dmgti example](#)

