

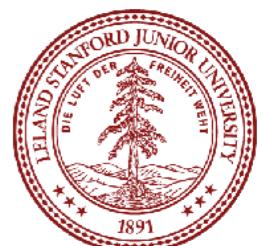


Unveiling the Structure and Evolution of Black Hole Coronae

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Einstein Fellows' Symposium 2017



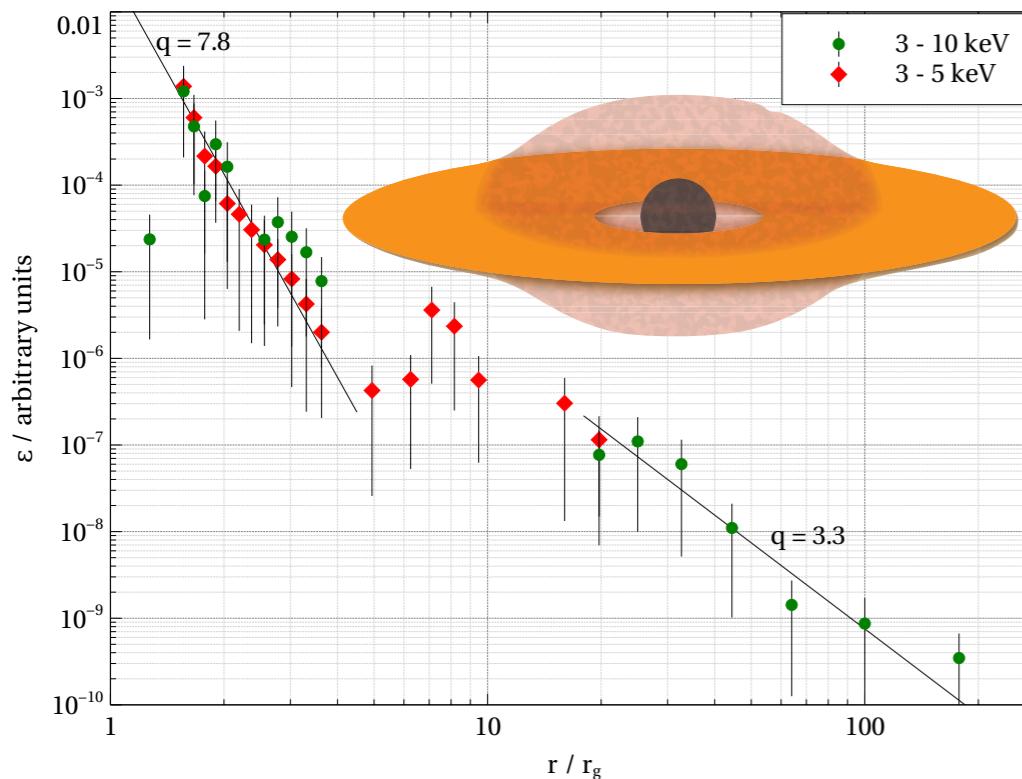
Why do we care about the corona?

- How are some of the brightest objects in the Universe powered?
- How is the corona formed and powered by the accretion flow?
- How does the corona evolve in time? What regulates AGN activity?
- How are jets launched and are they related to the corona?

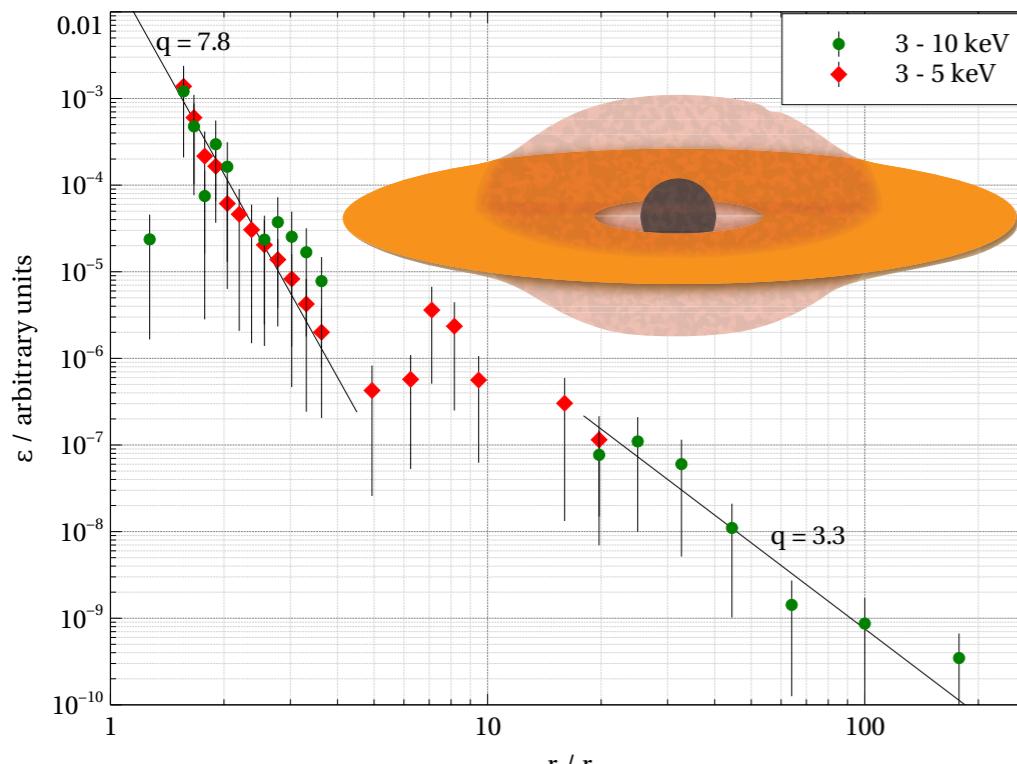
Outline

- What can X-ray reverberation tell us about the structure of the corona?
- Structure in the corona of I Zw 1
- Evolution of the corona during flares

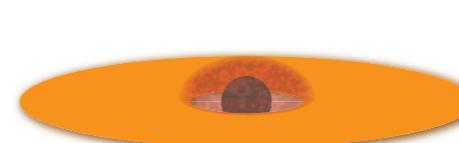
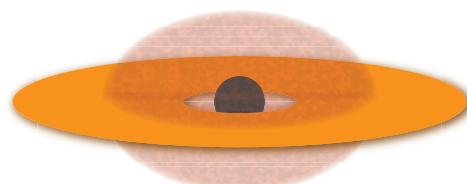
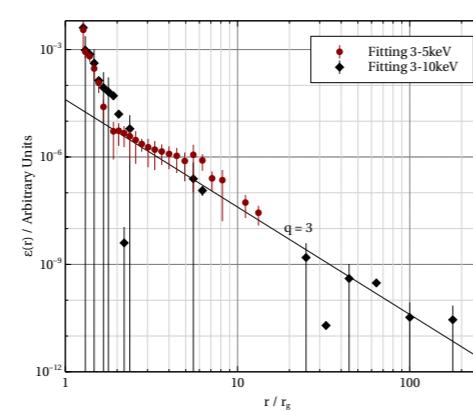
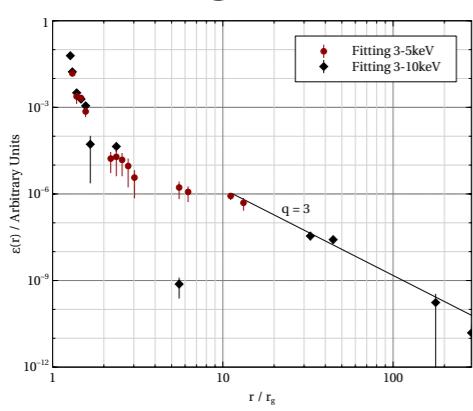
What we know about the corona?



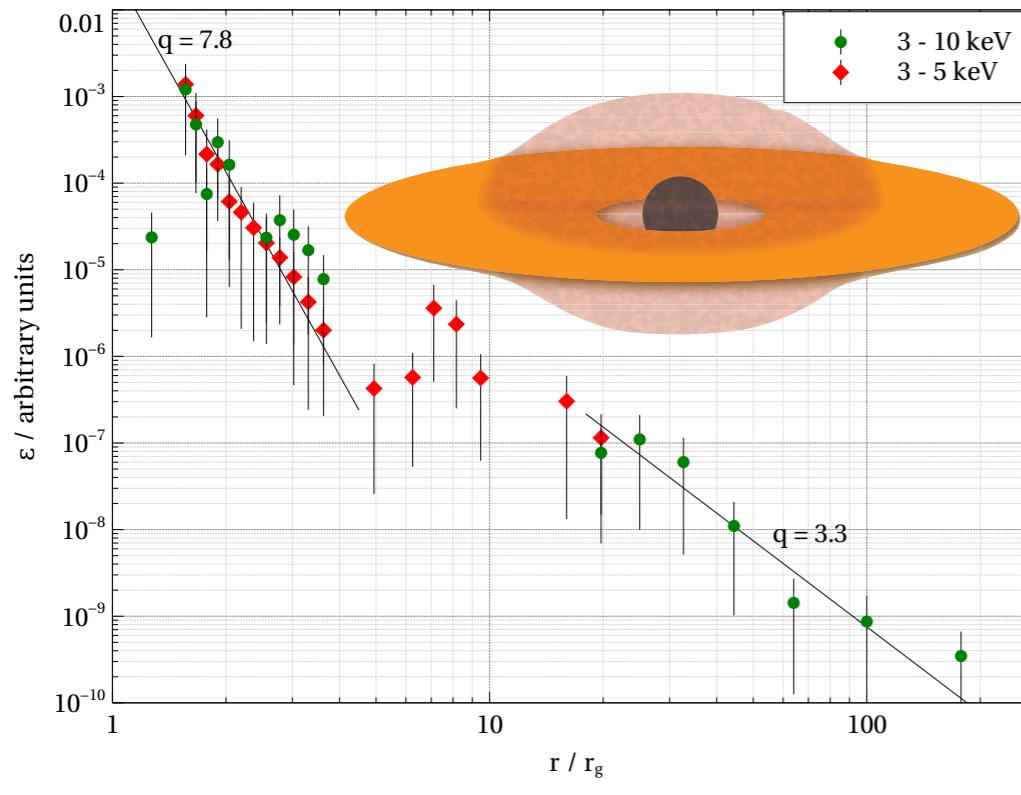
What we know about the corona?



Low Flux

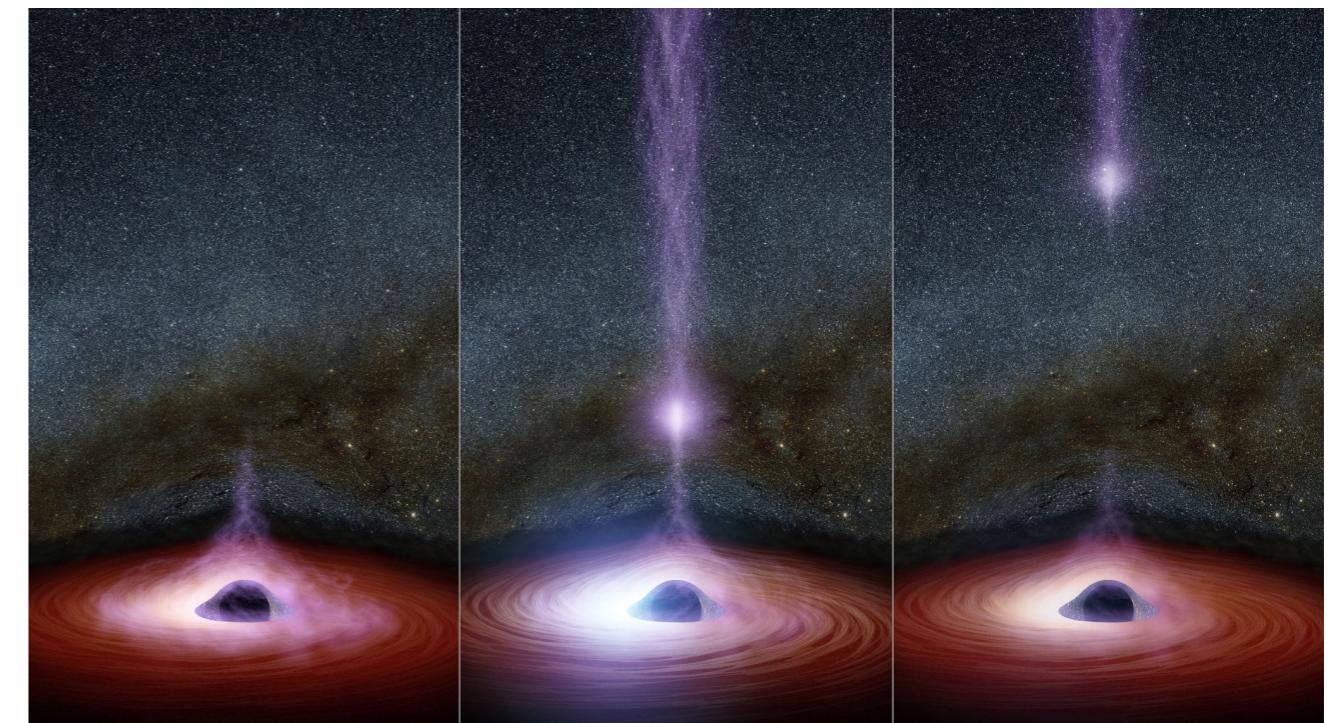
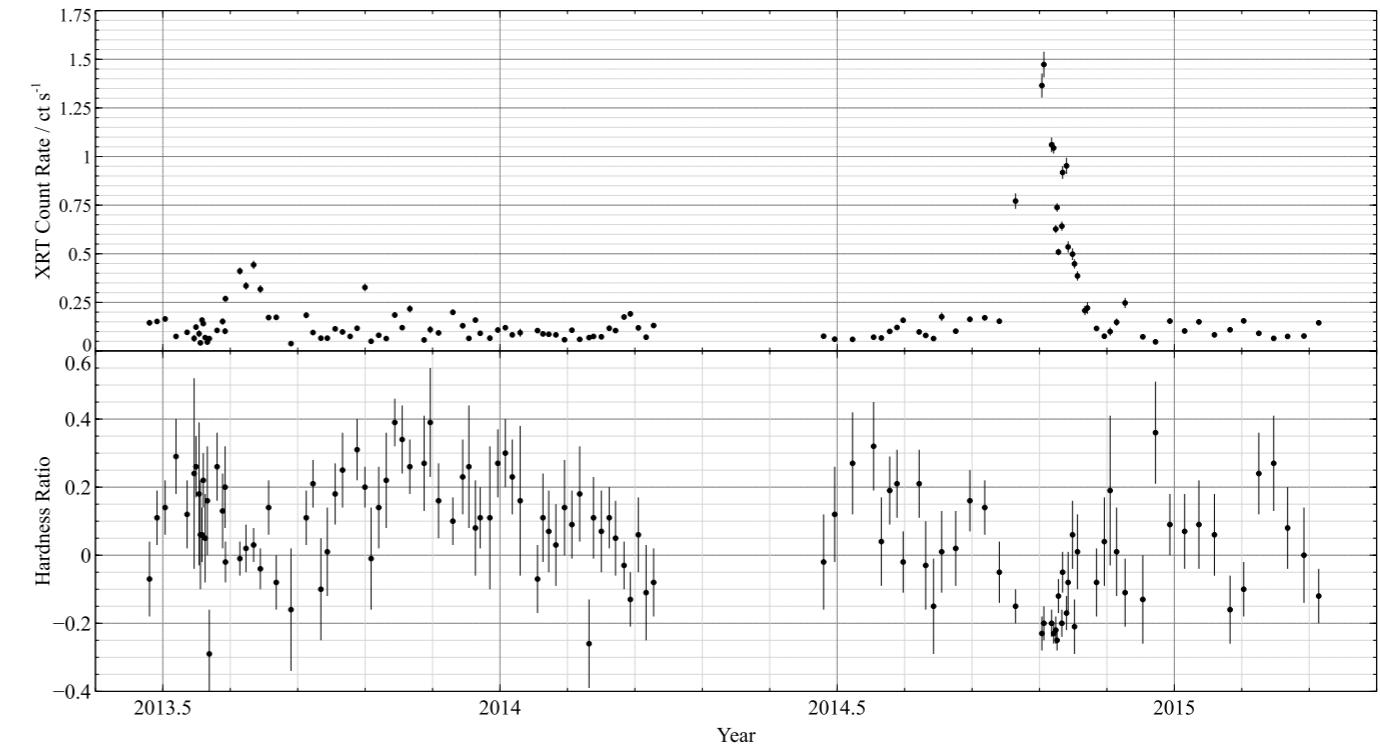
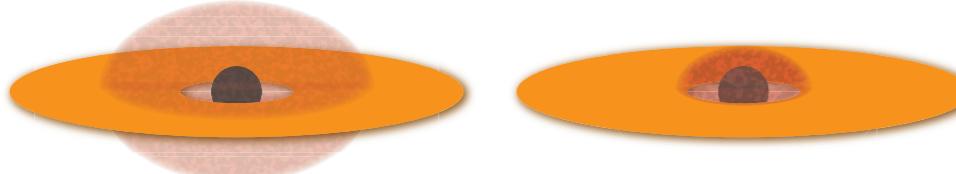
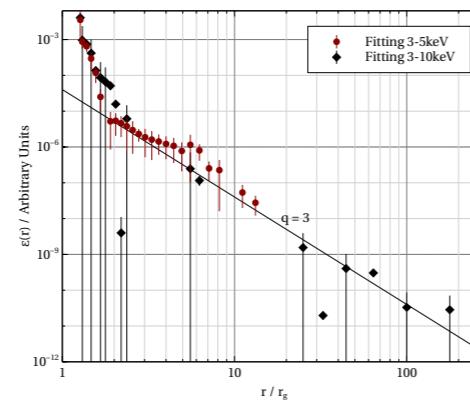
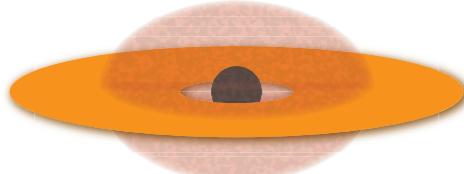
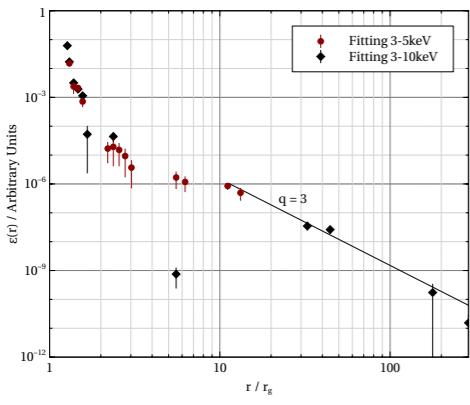


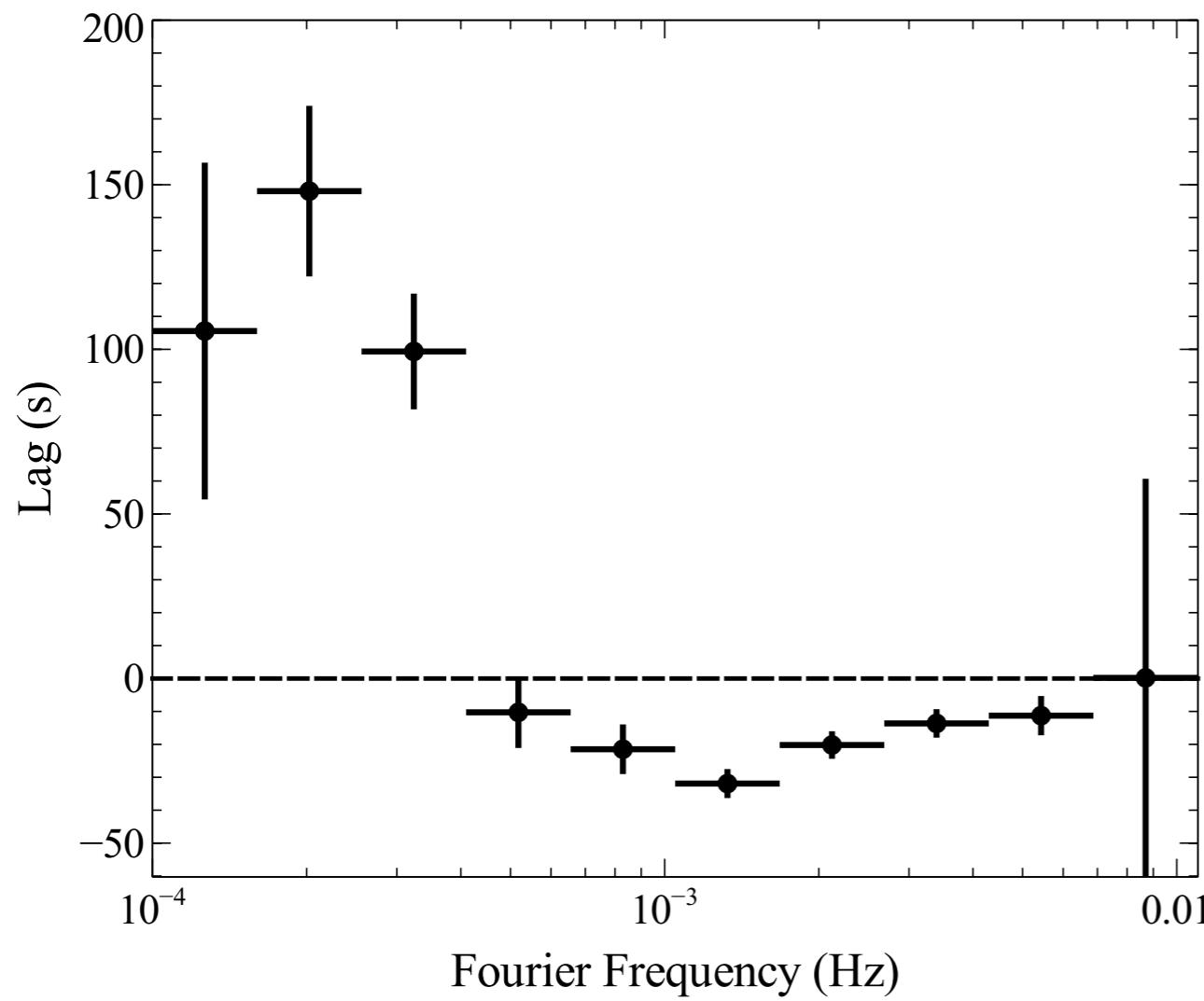
What we know about the corona?



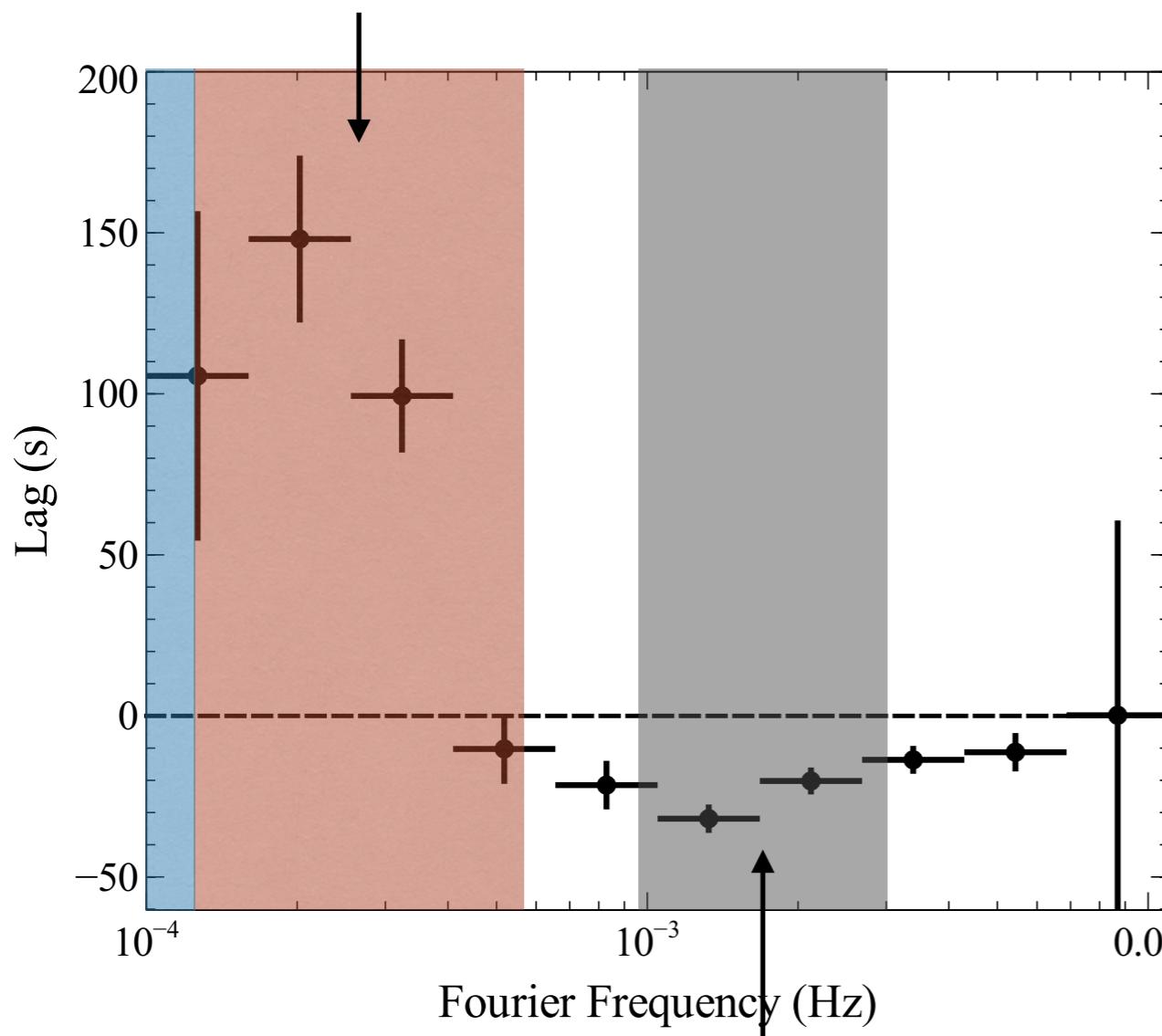
High Flux

Low Flux

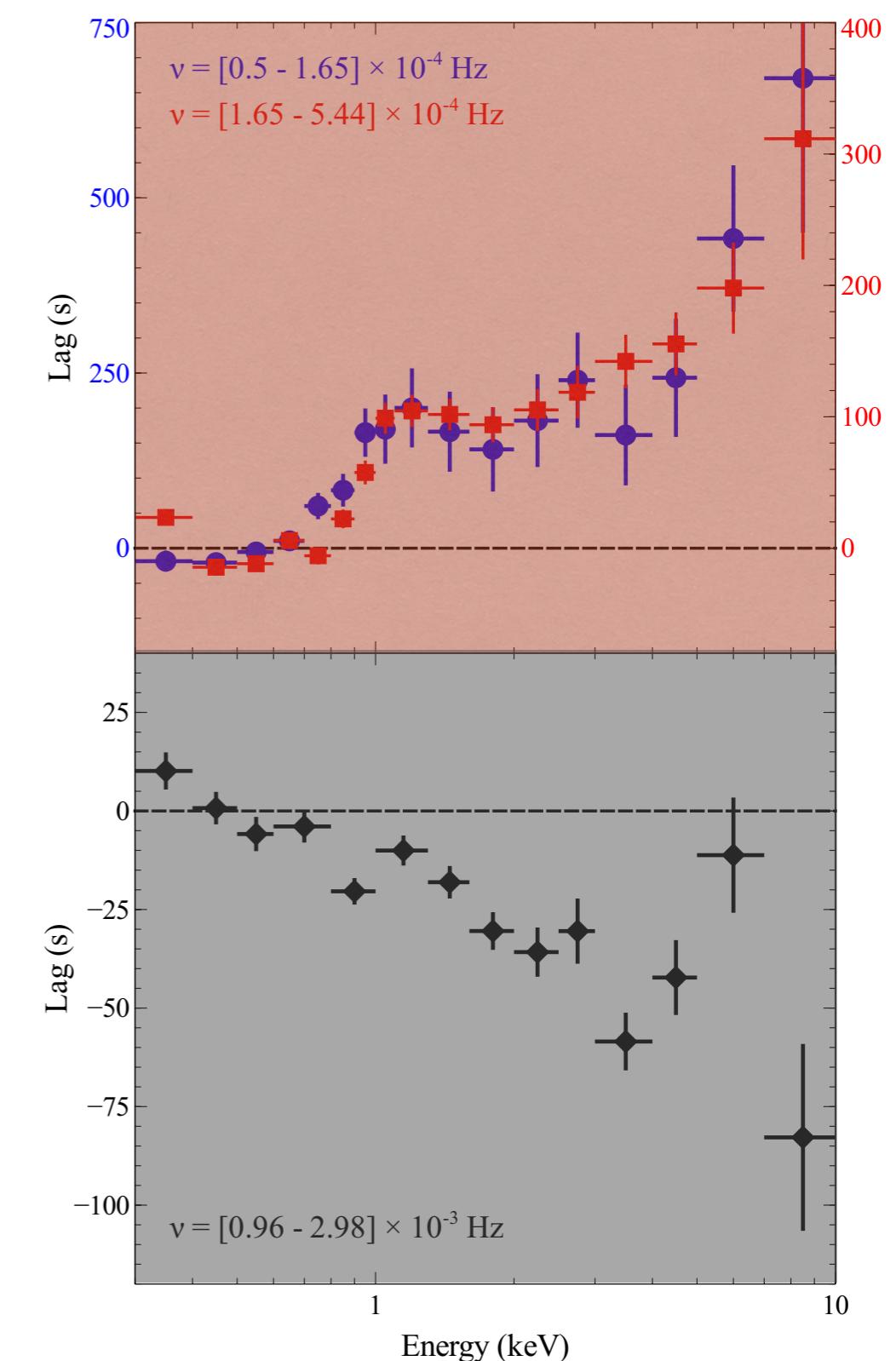




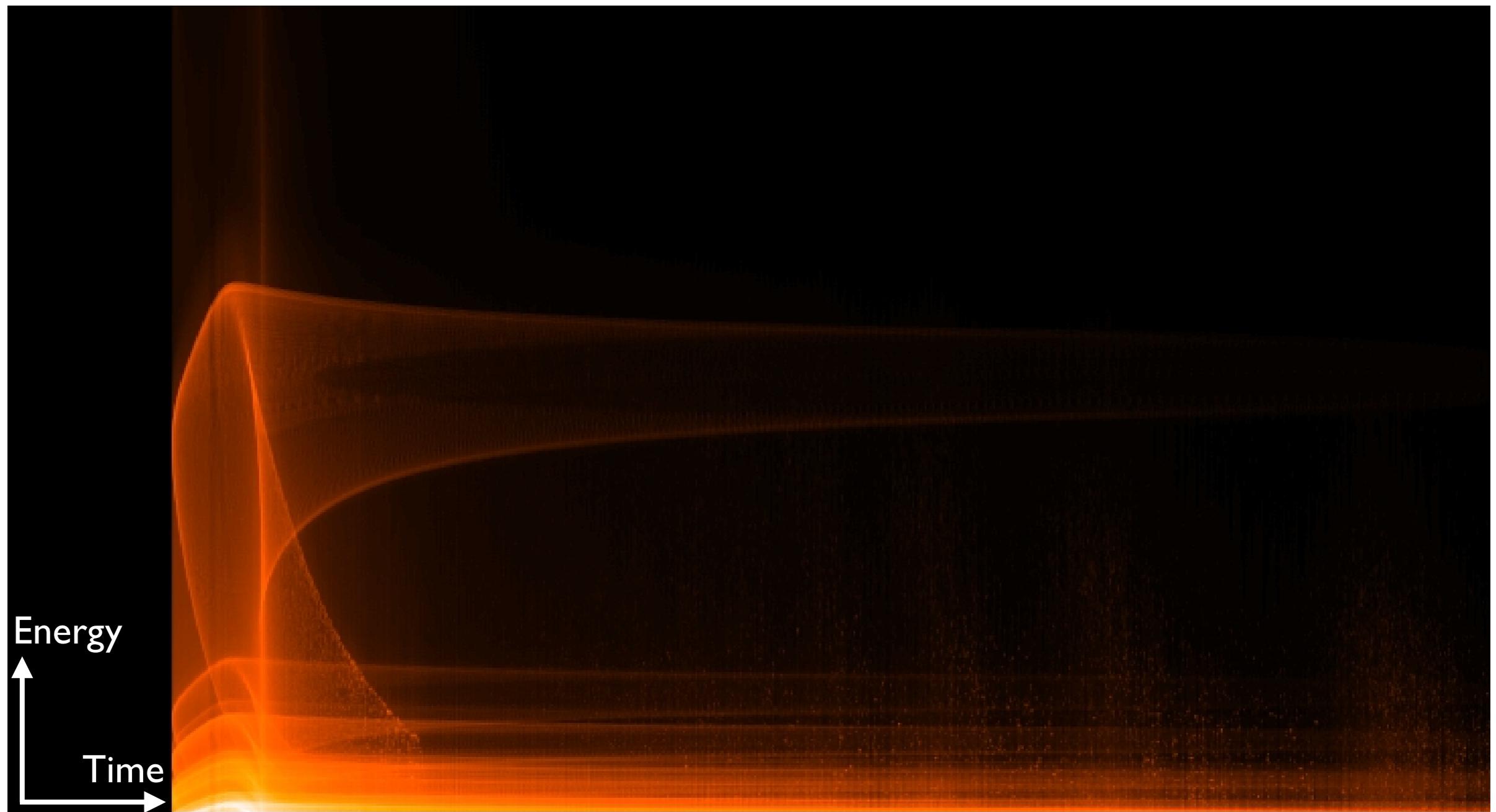
Hard X-rays lag soft (in continuum)



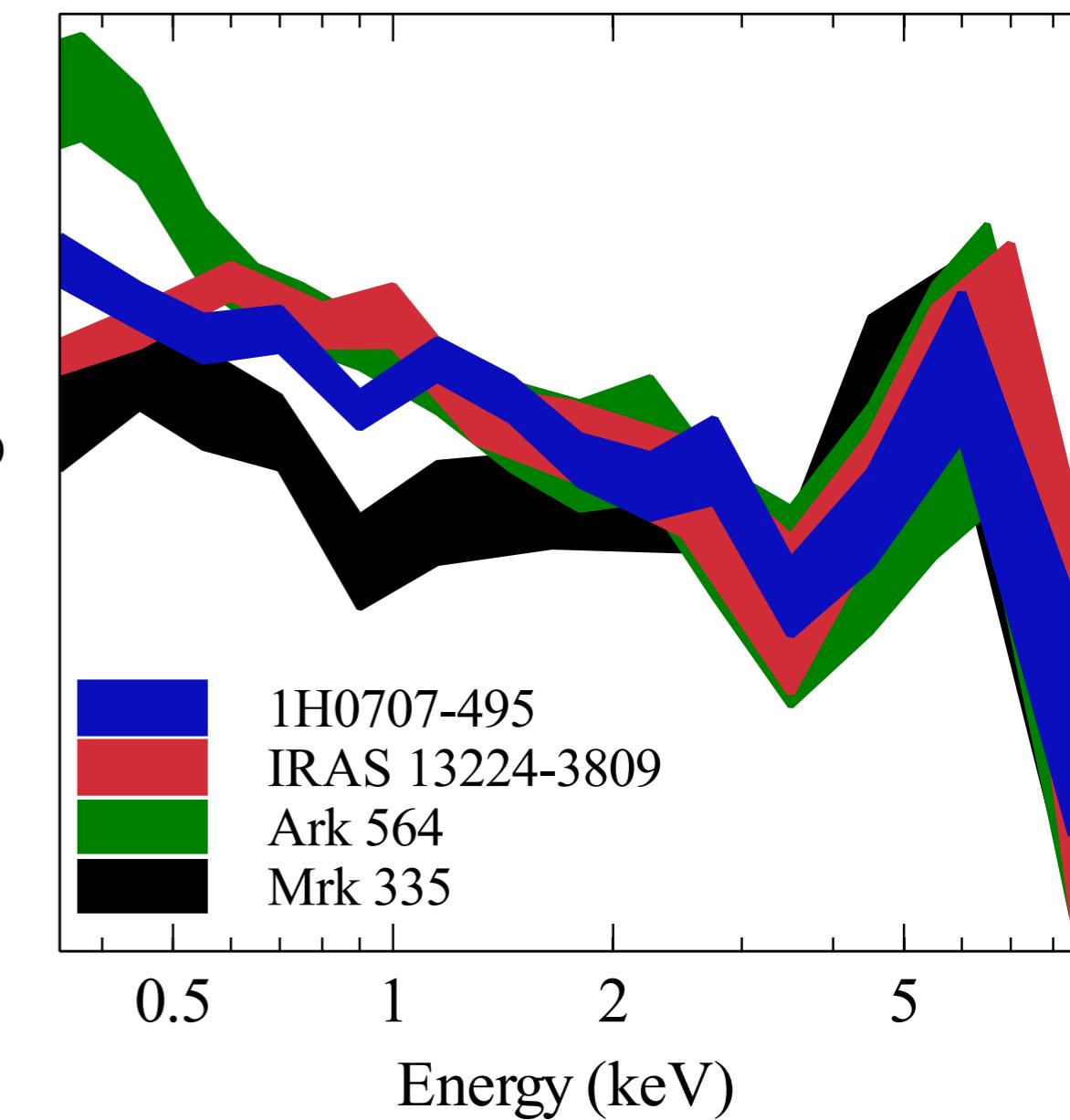
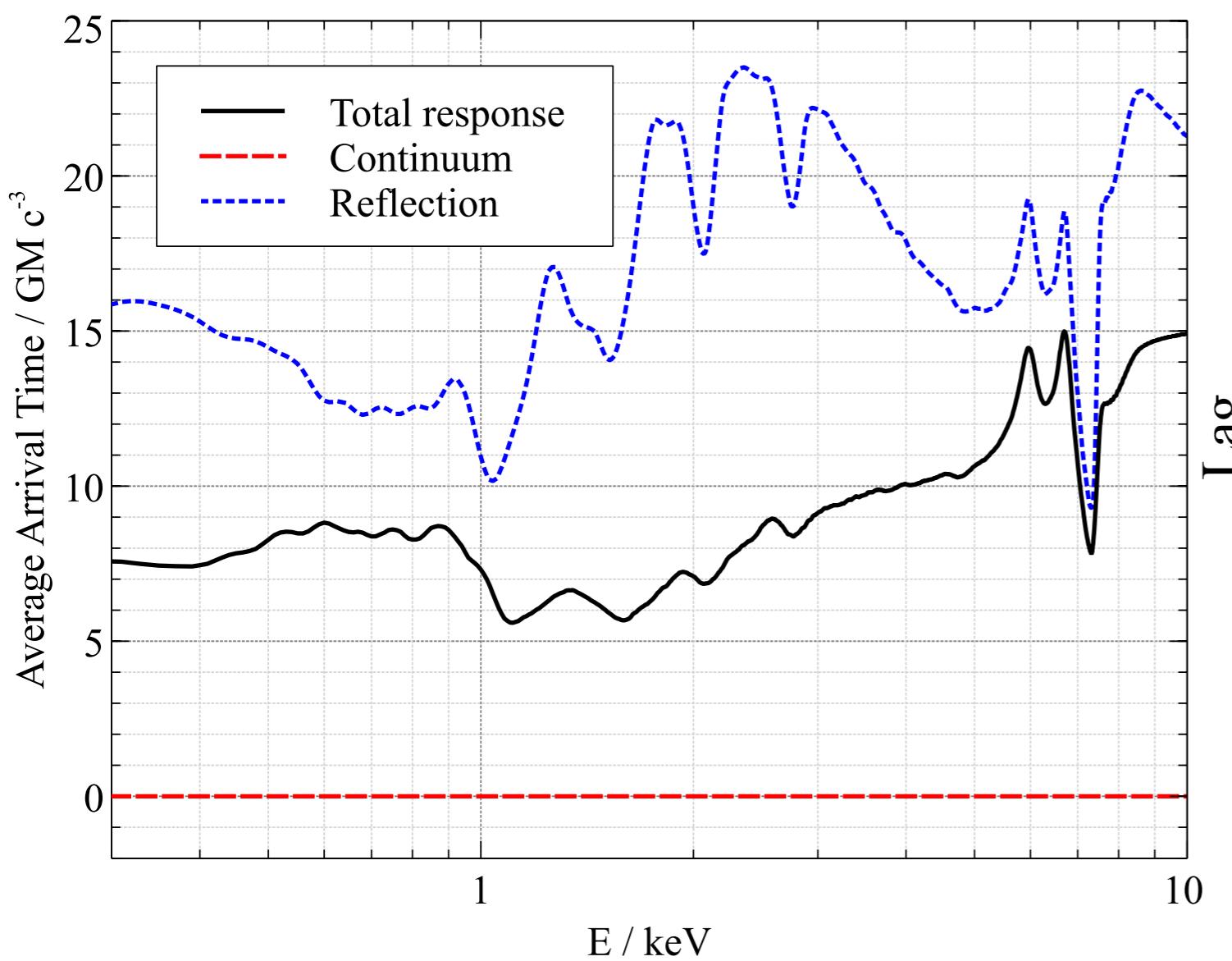
Soft (reflection-dominated)
lags hard (continuum)



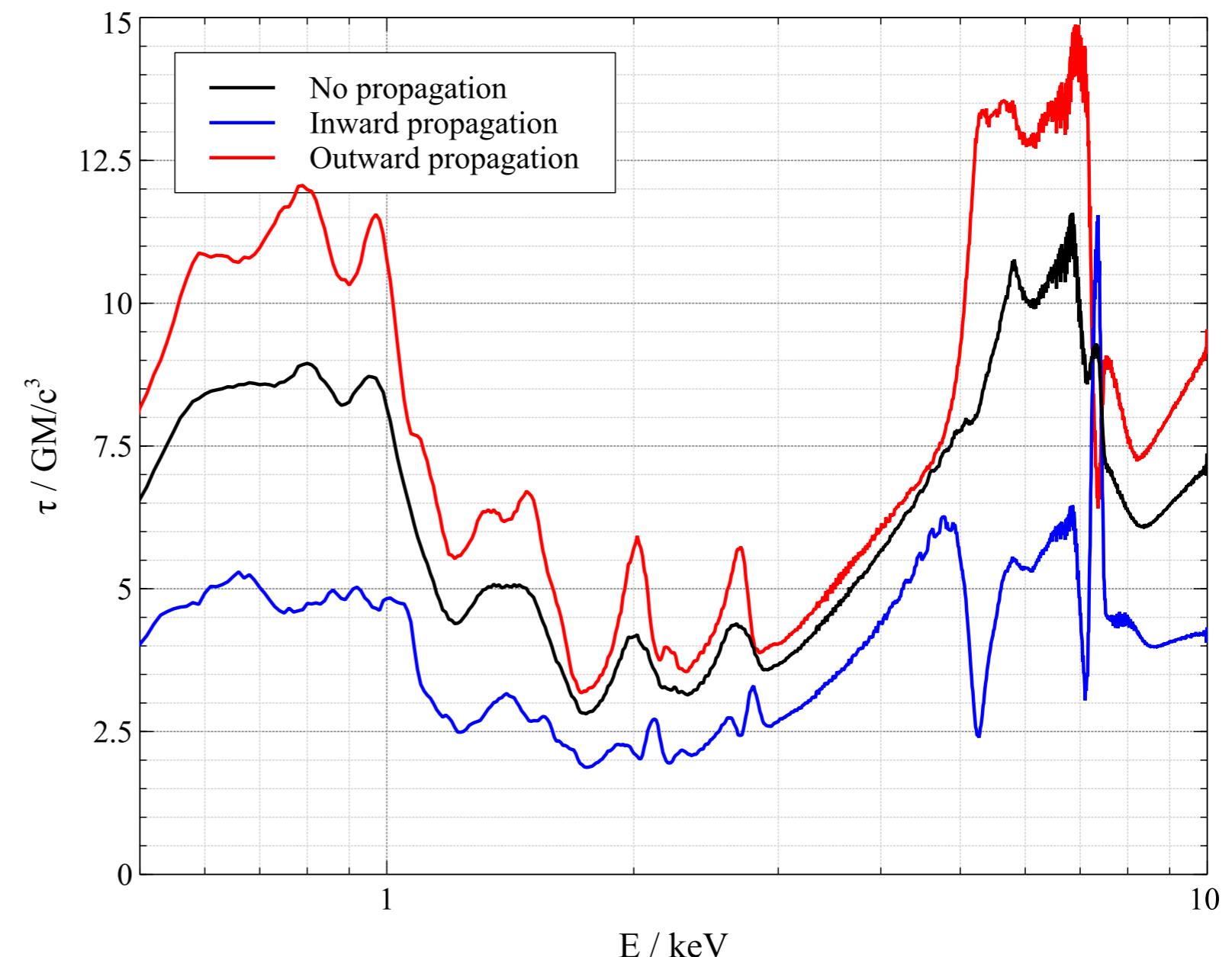
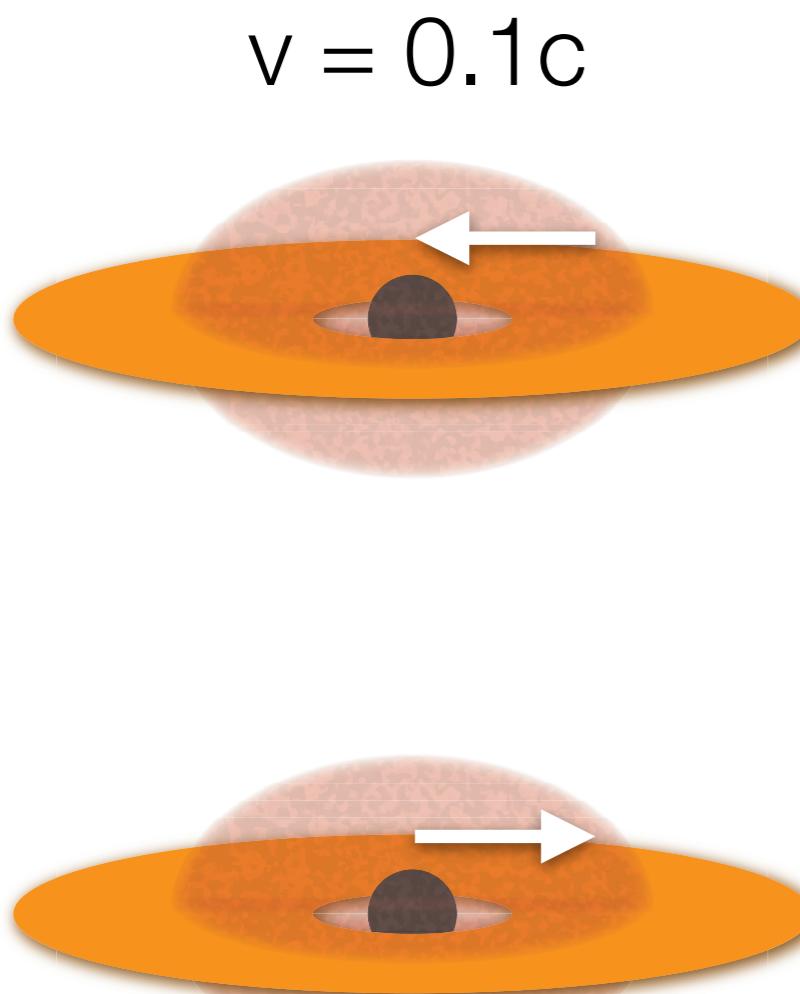
Reverberation Response



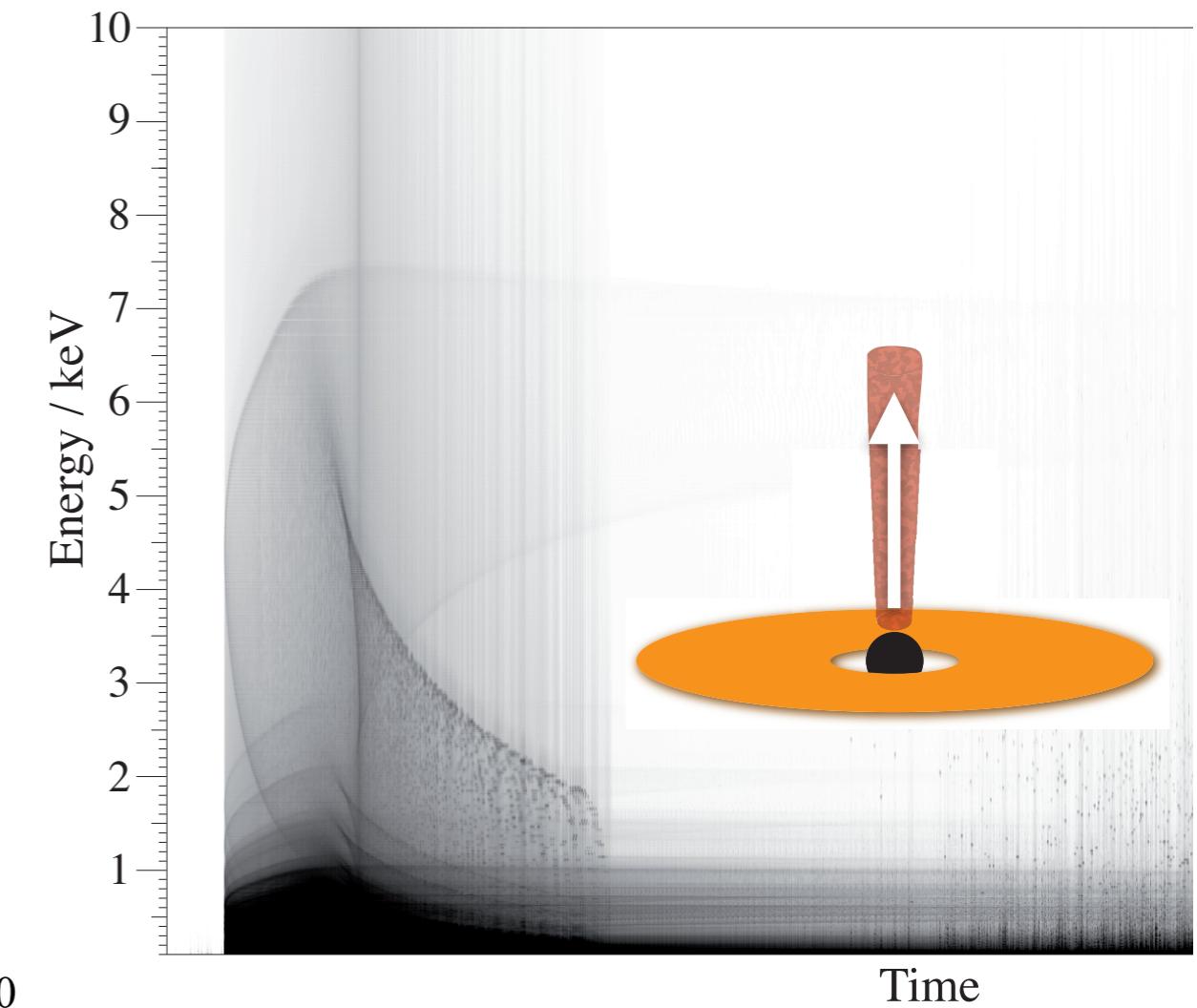
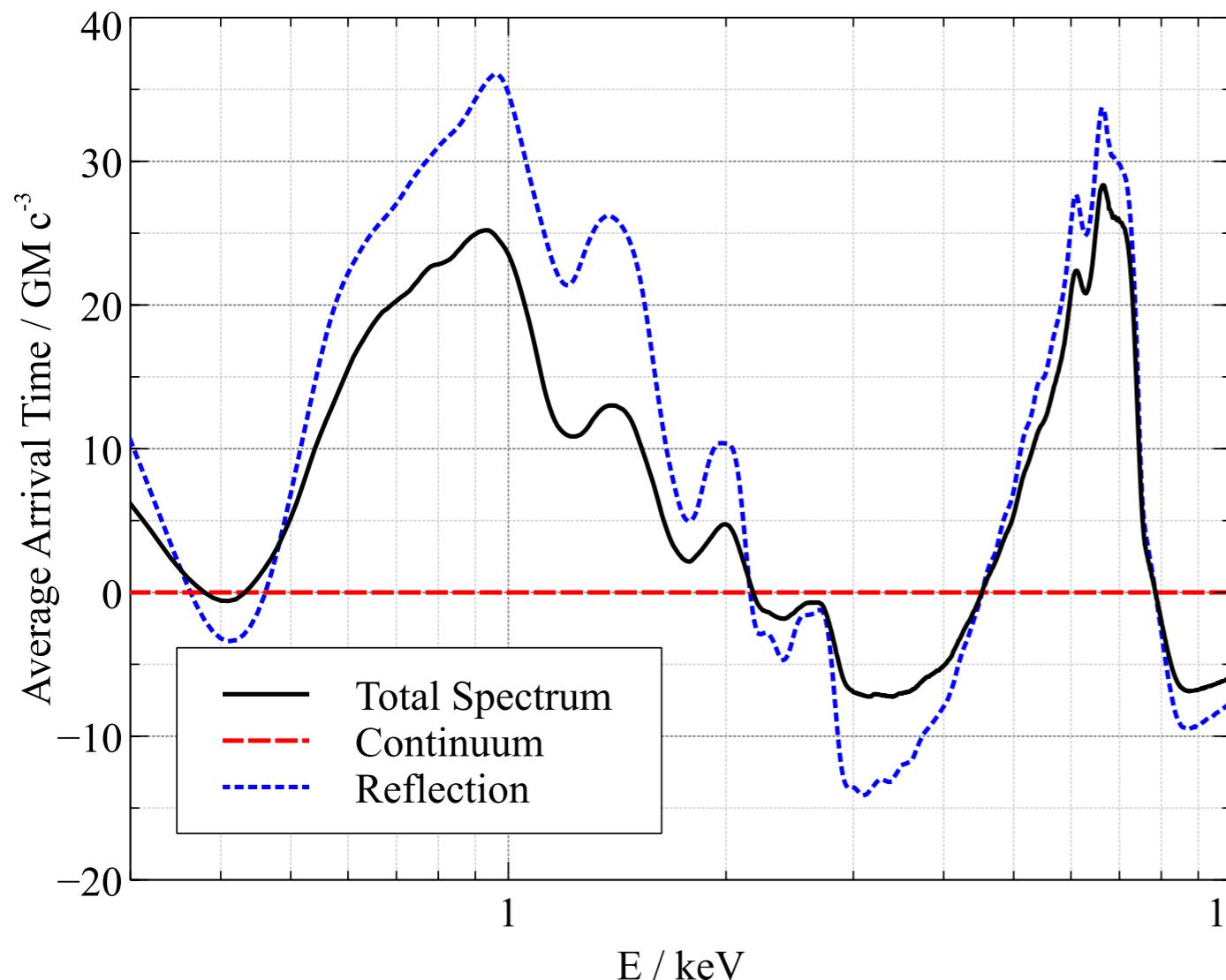
The Lag-Energy Problem



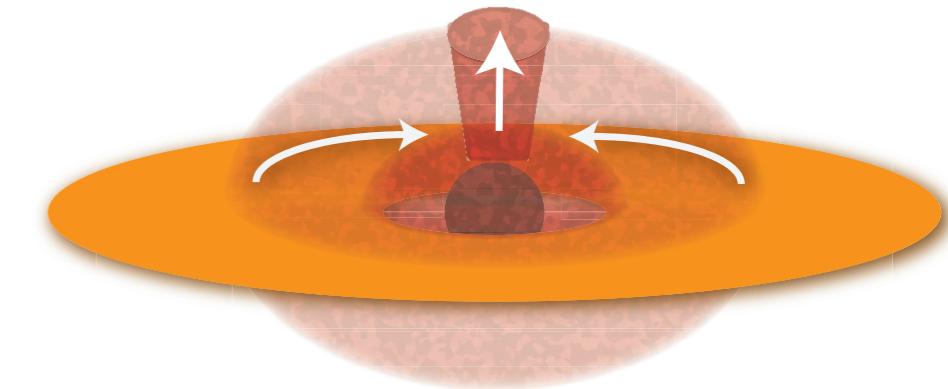
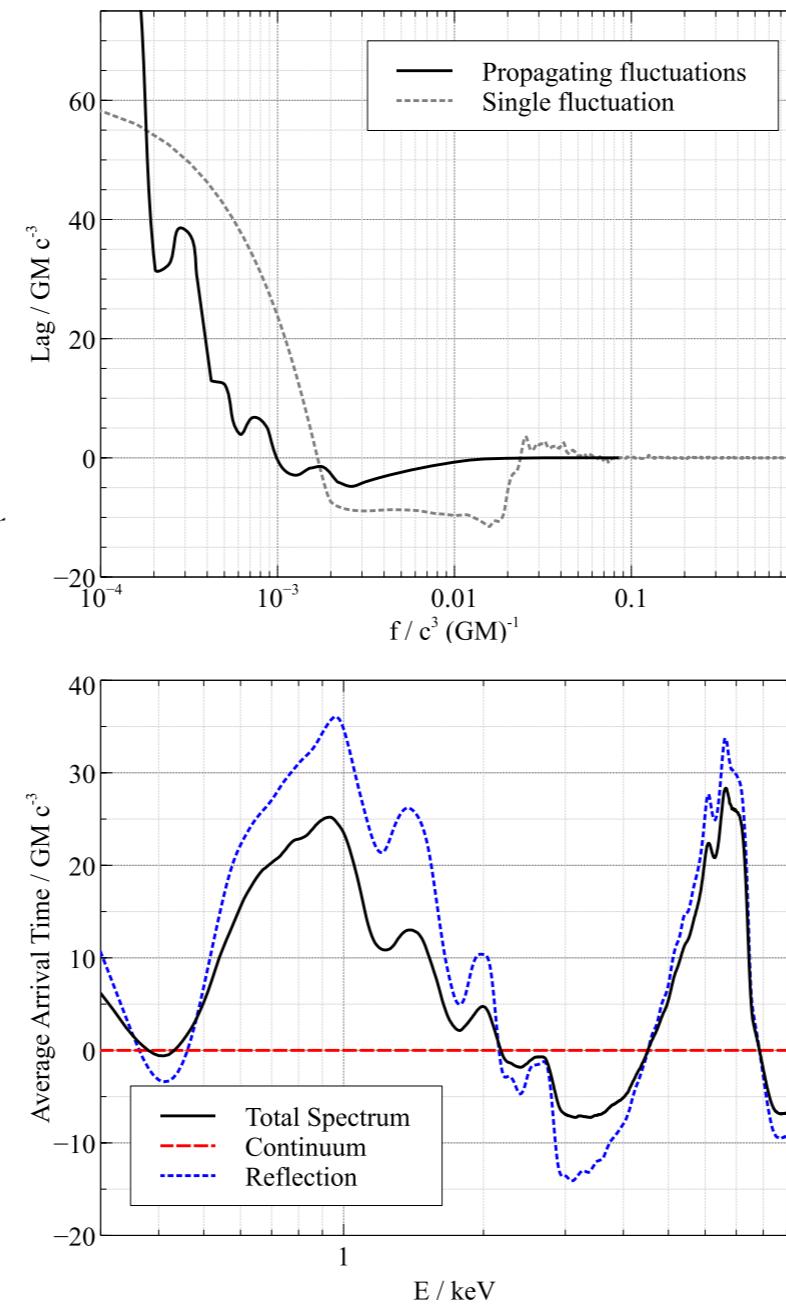
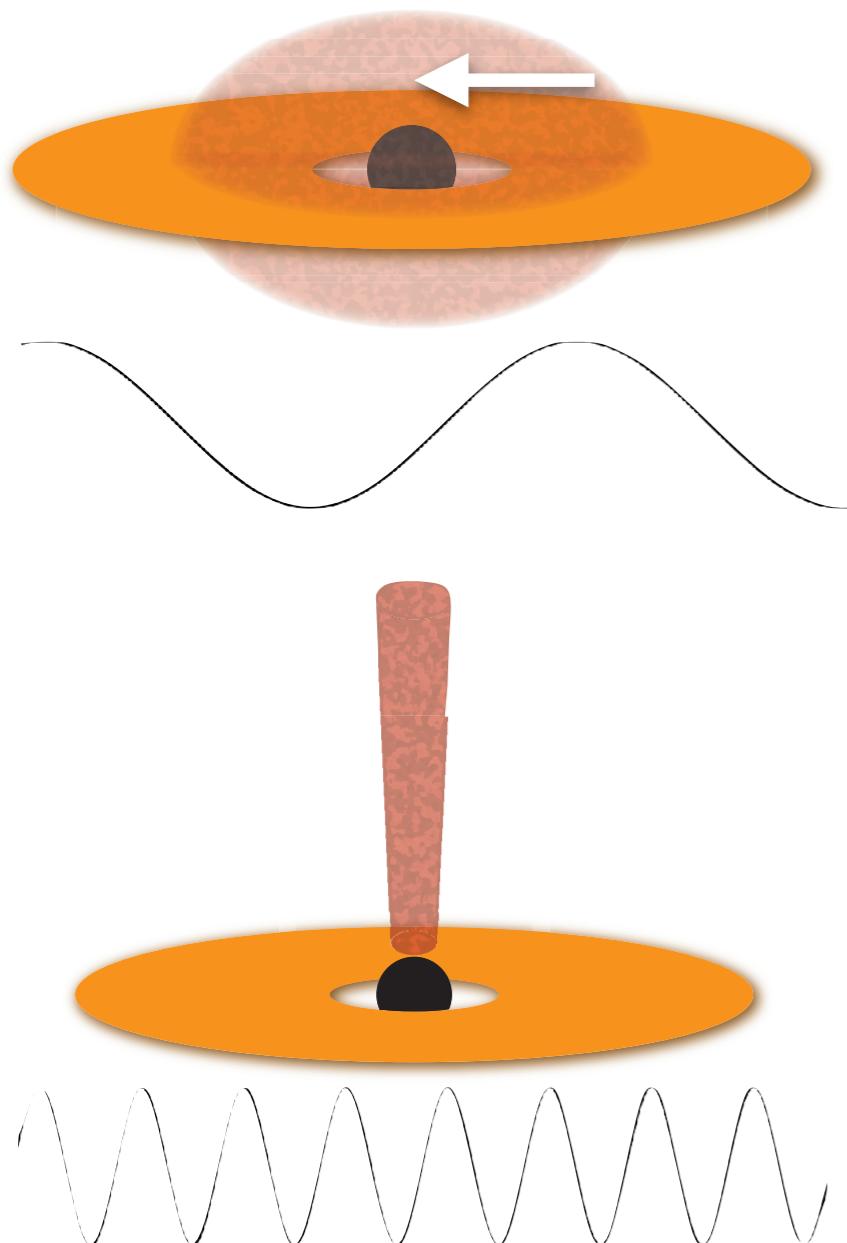
Extended Coronae in Lag-Energy



What is the dip at 3keV?

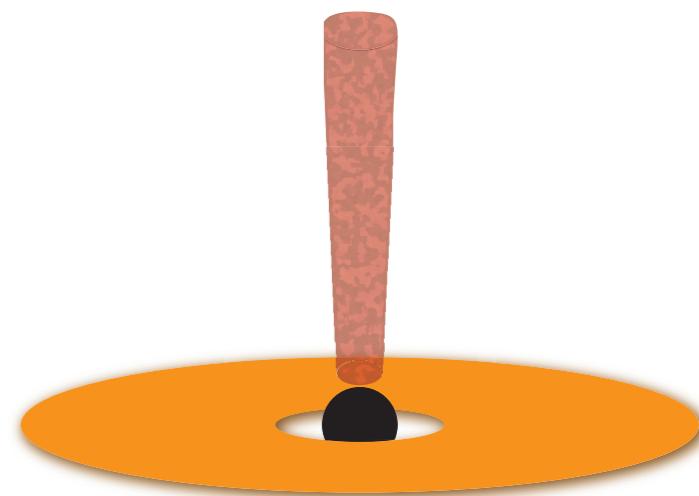
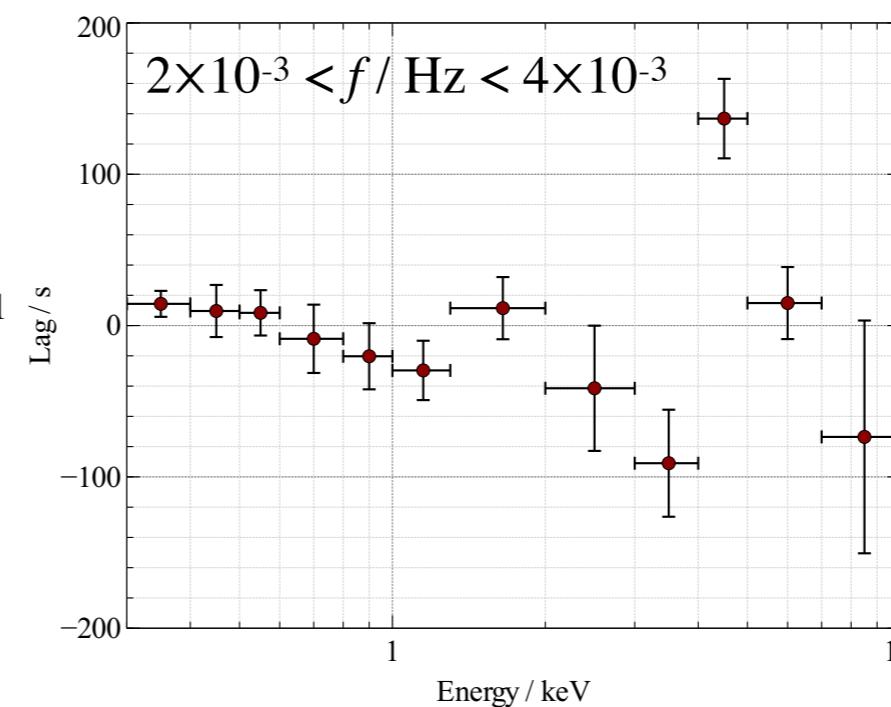
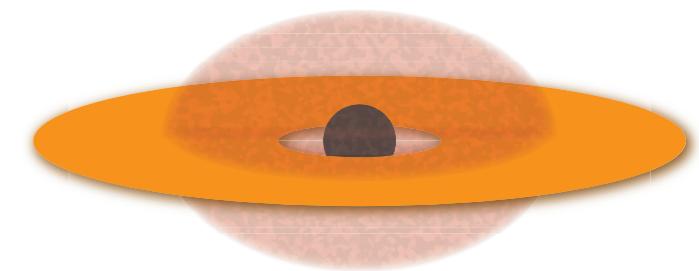
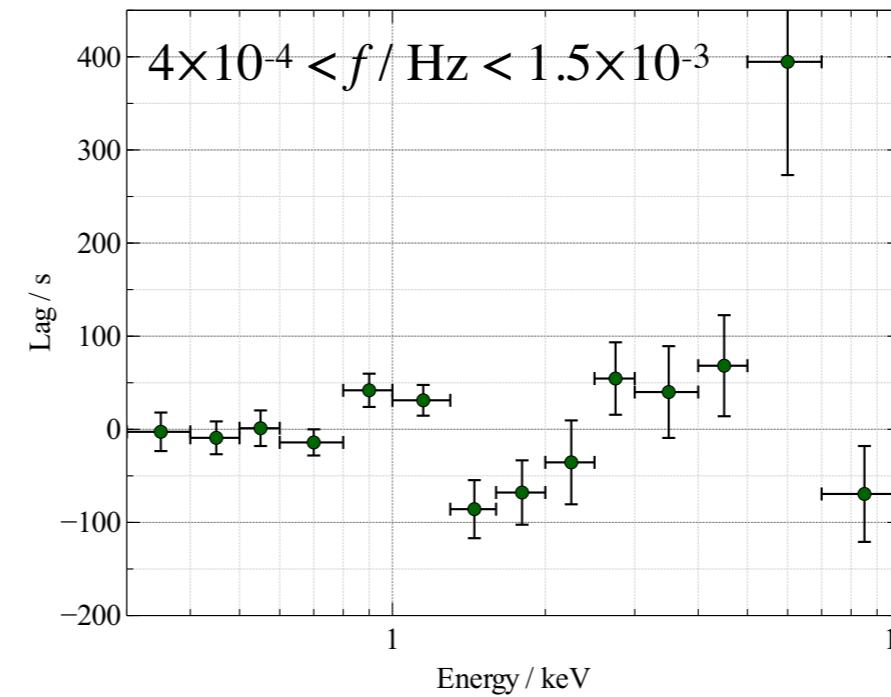
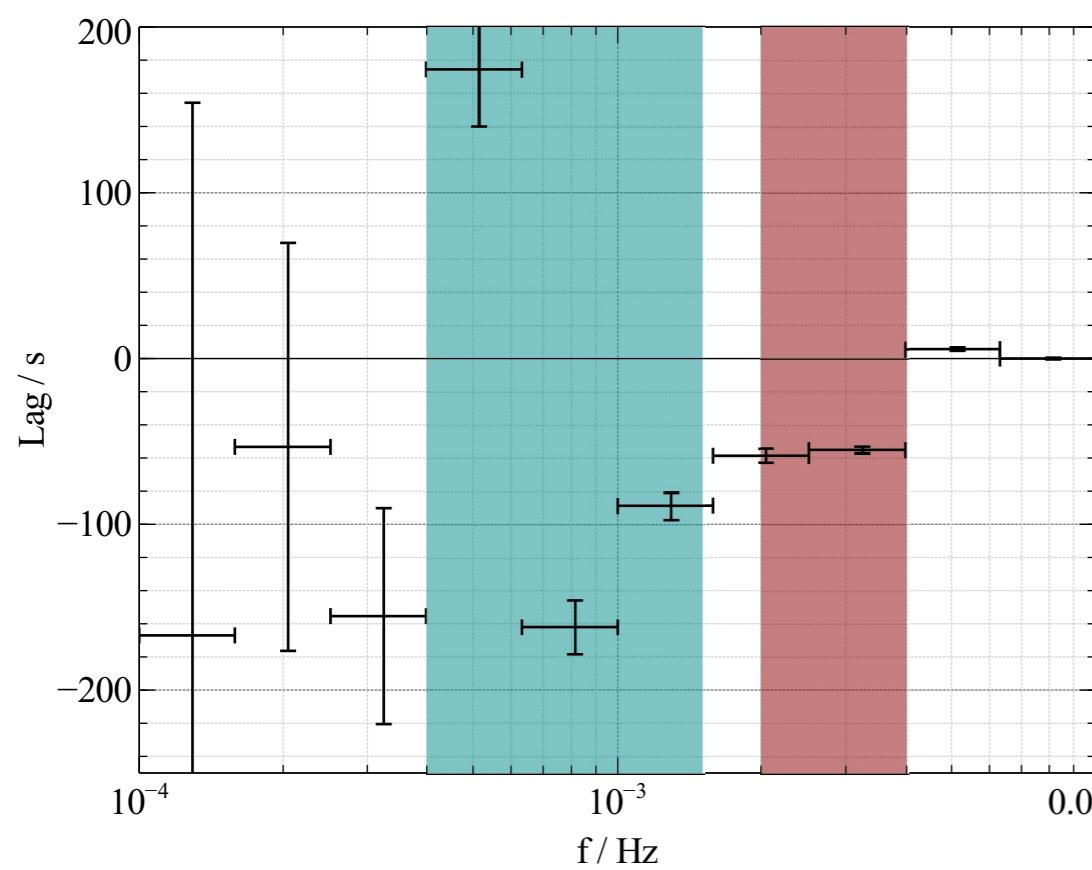


A tale of two coronae?

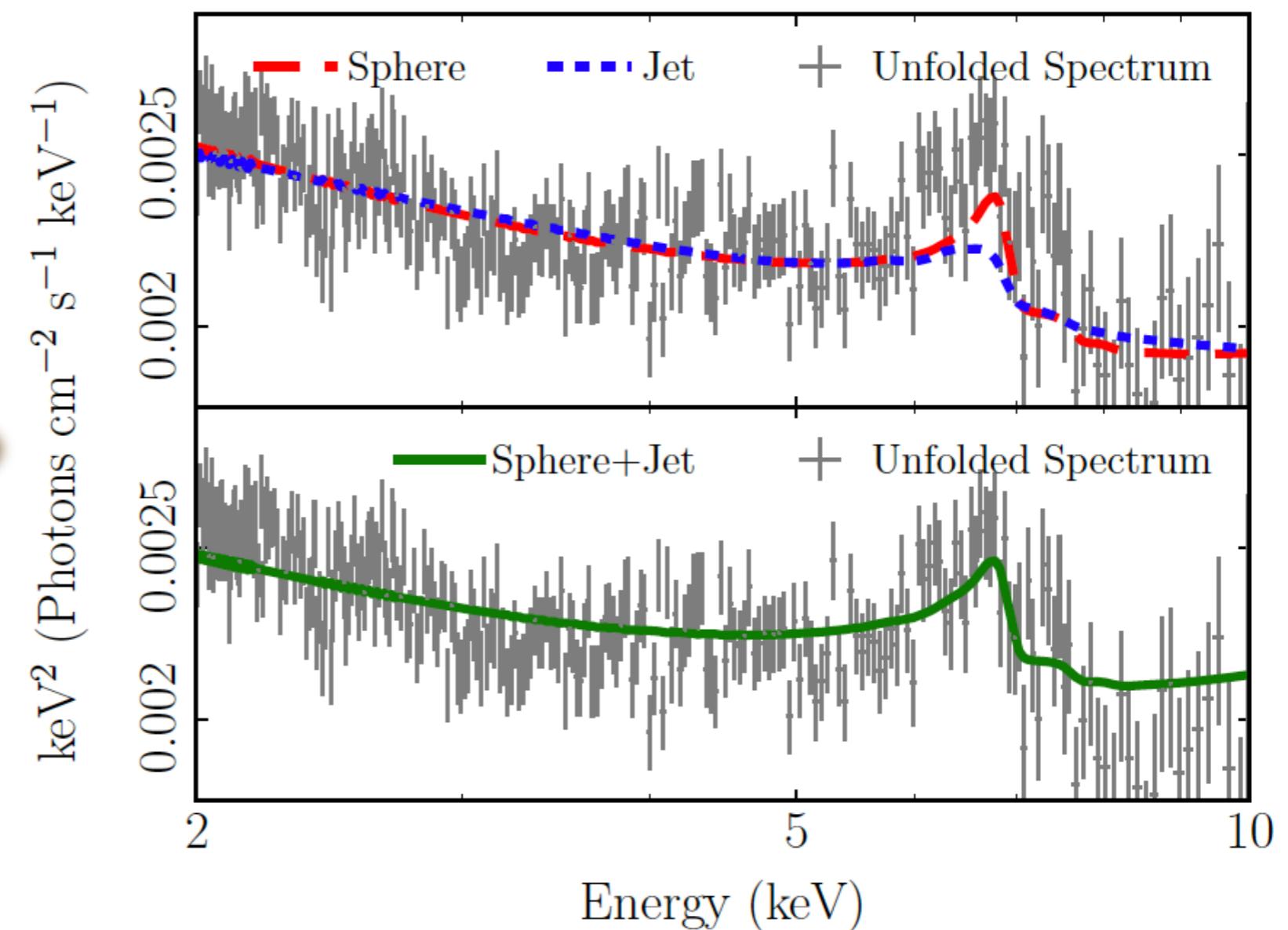
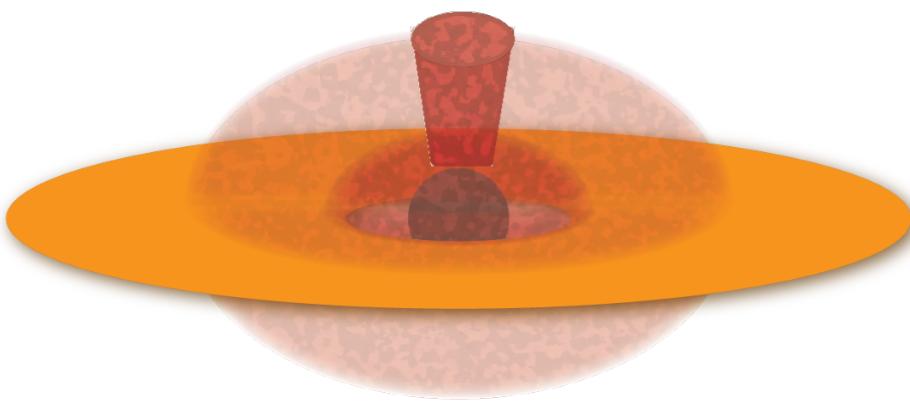


Seeing the structure in 1 Zw 1

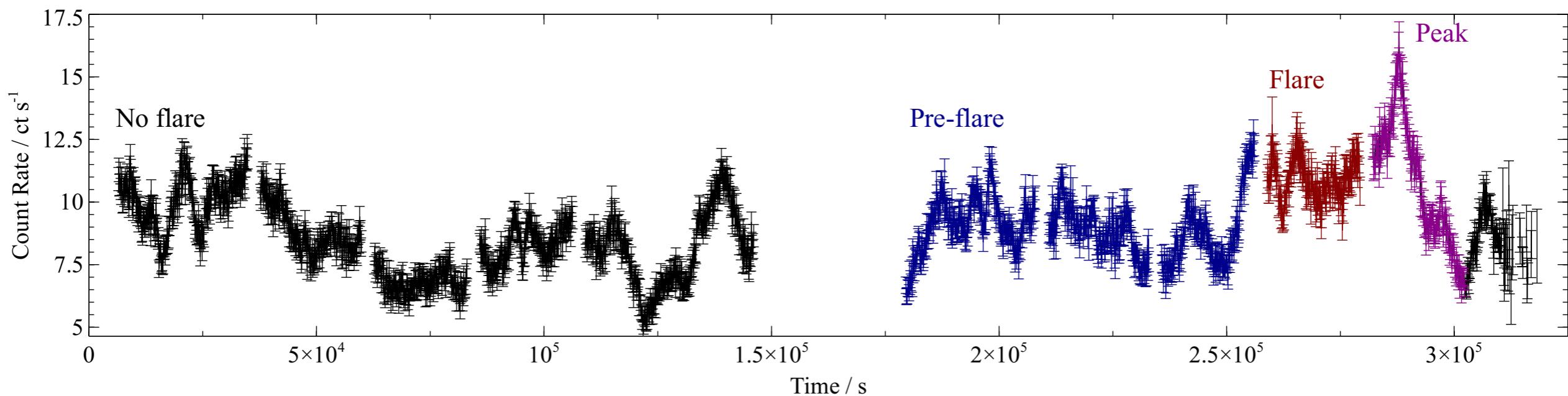
XMM Newton, January 2015
2 orbits (275ks)



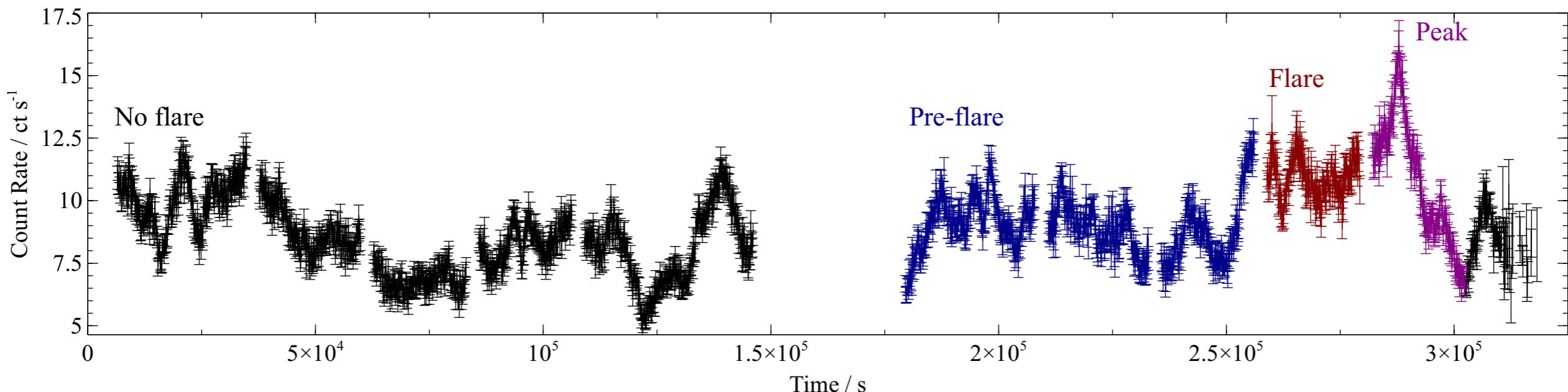
Evidence from the Spectrum



Evolution of the Corona in I Zw 1

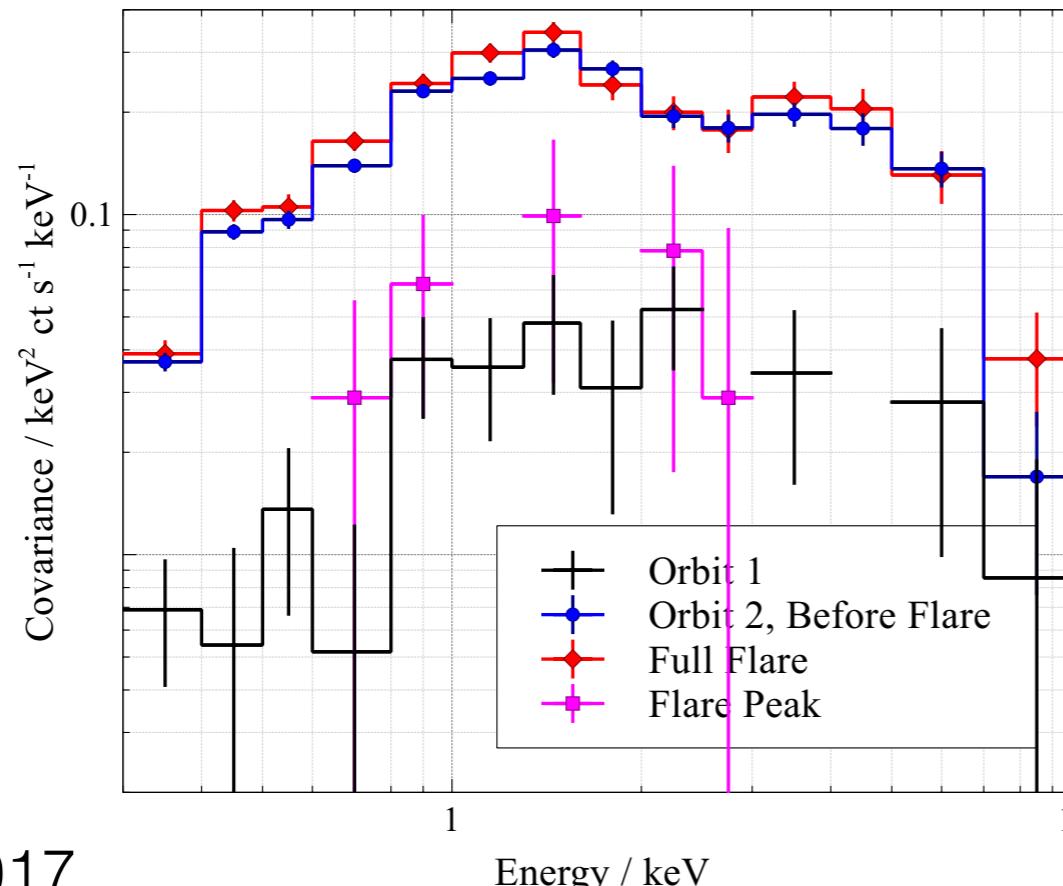


Evolution of the Corona in I Zw 1

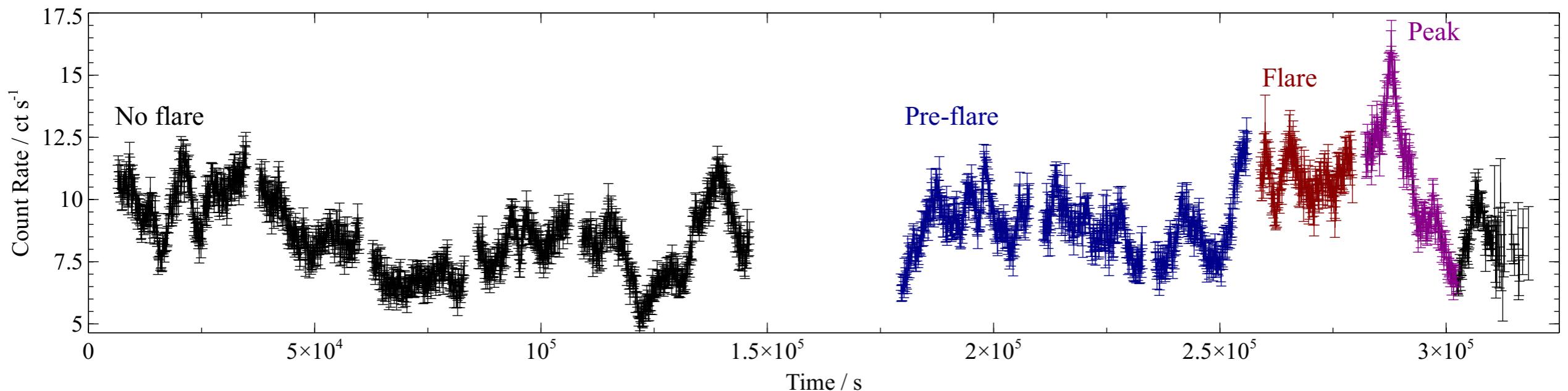


Covariance Spectrum

Left:
Low Frequency
 $4 \times 10^{-4} - 1.5 \times 10^{-3} \text{ Hz}$

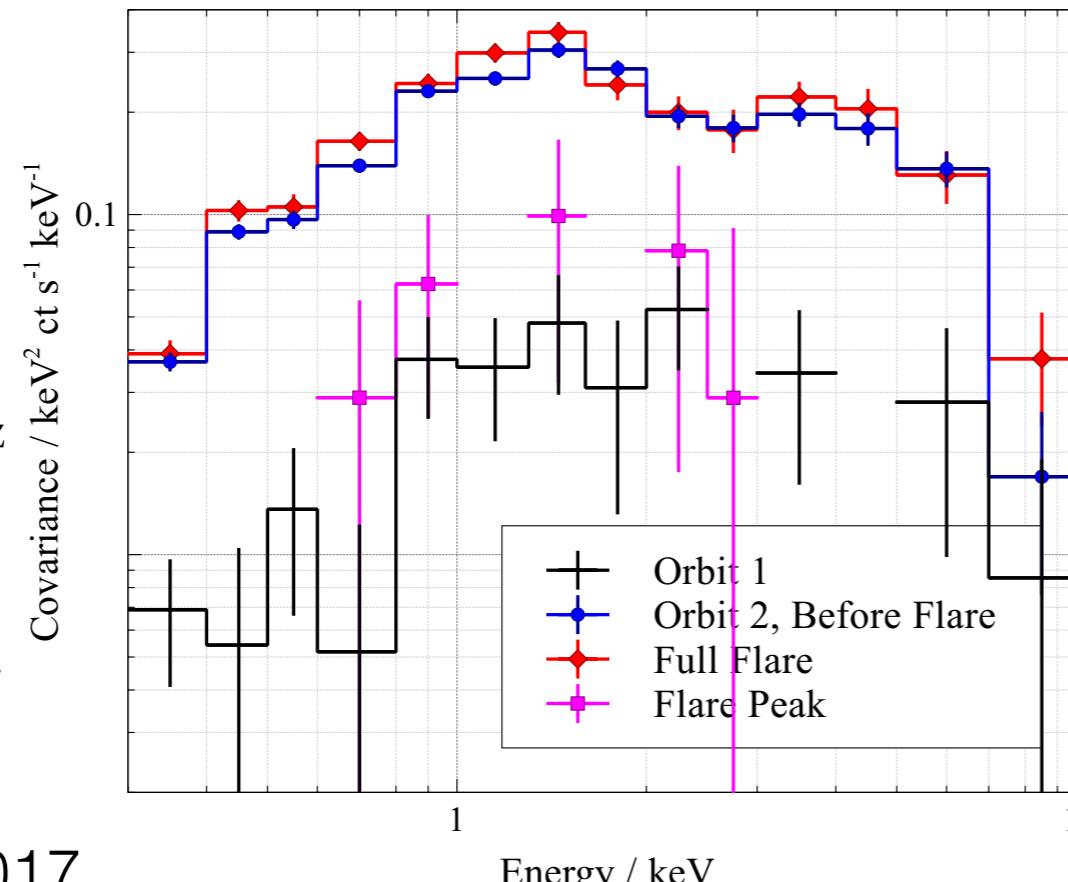


Evolution of the Corona in I Zw 1

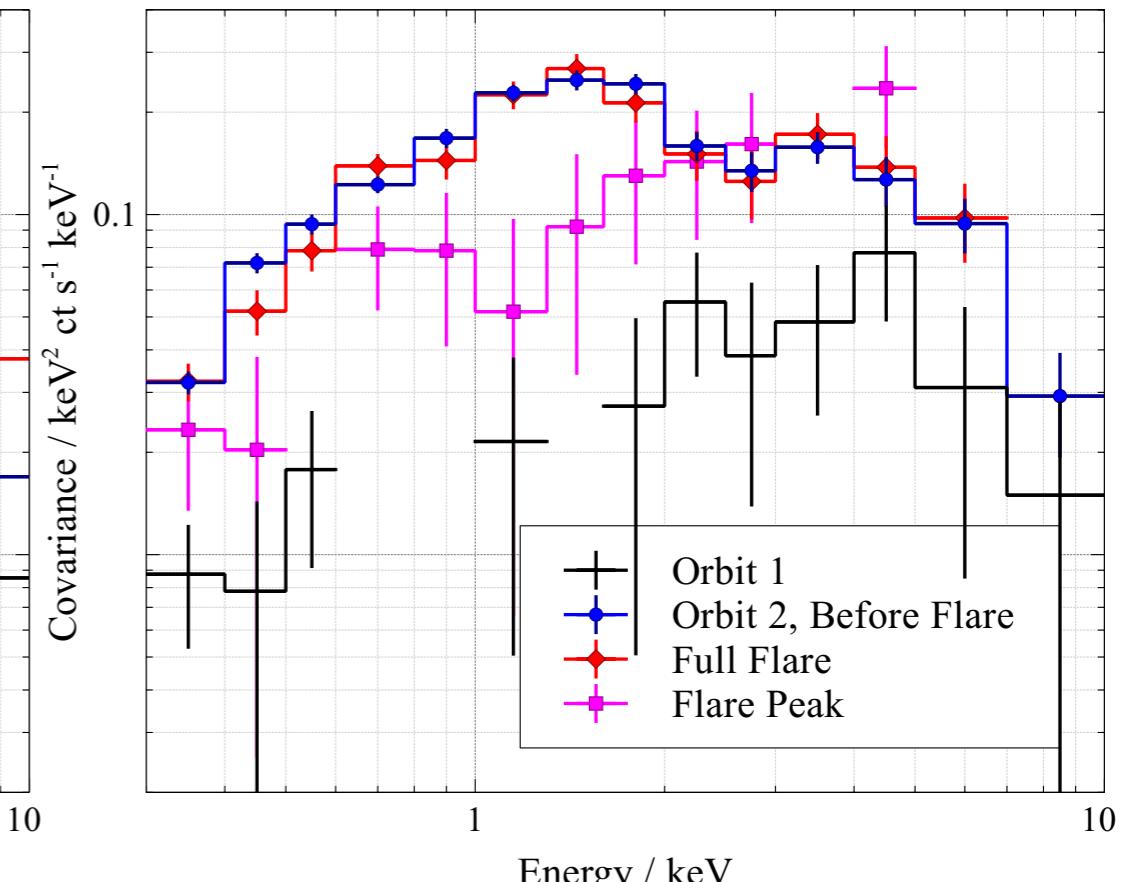


Covariance Spectrum

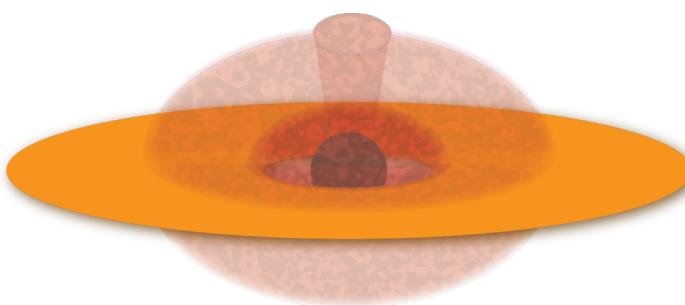
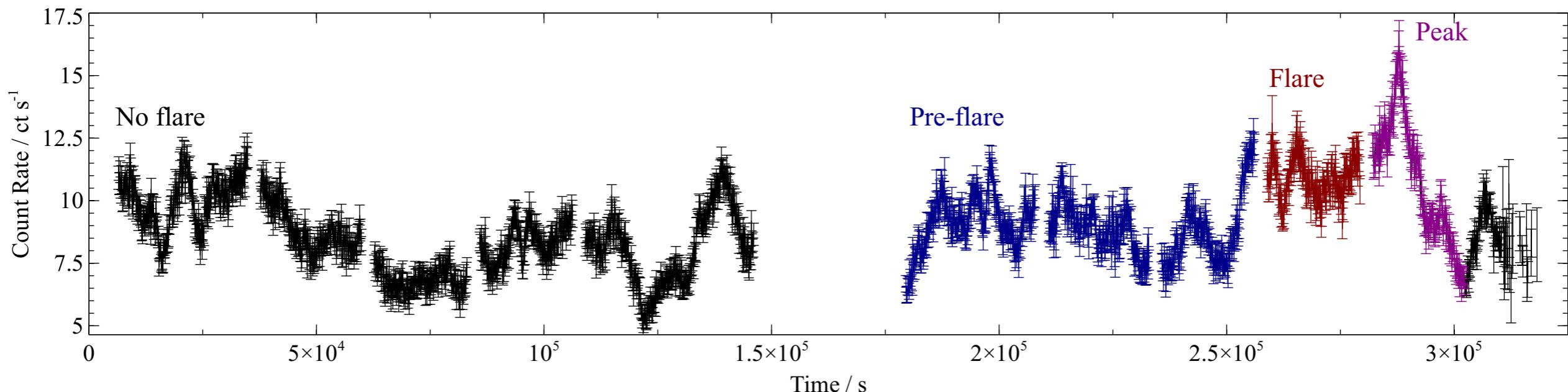
Left:
Low Frequency
 $4 \times 10^{-4} - 1.5 \times 10^{-3}$ Hz



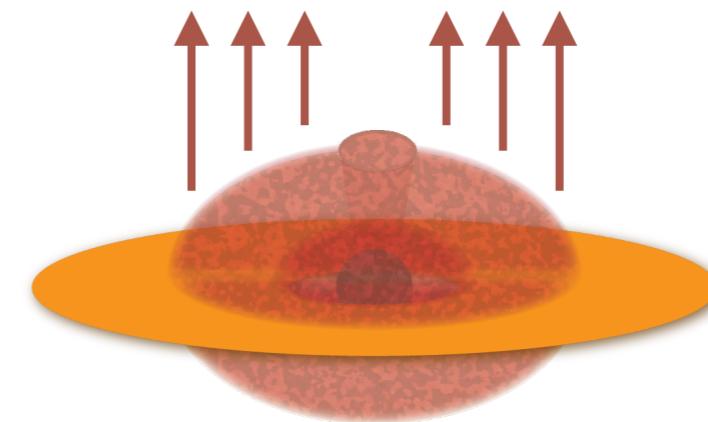
Right:
High Frequency
 $2 \times 10^{-3} - 4 \times 10^{-3}$ Hz



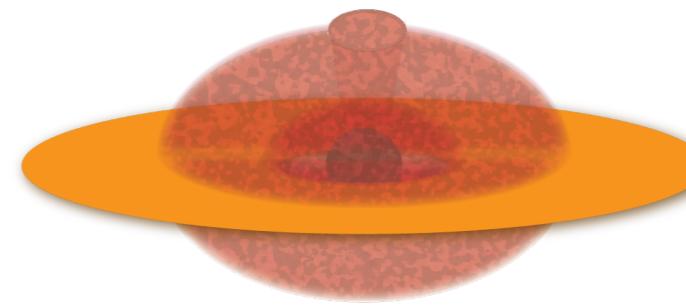
What is the corona doing?



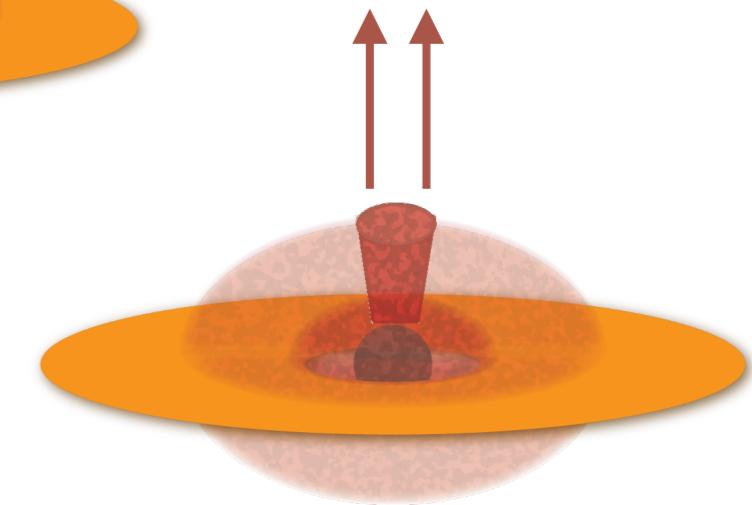
No Flare



Flare

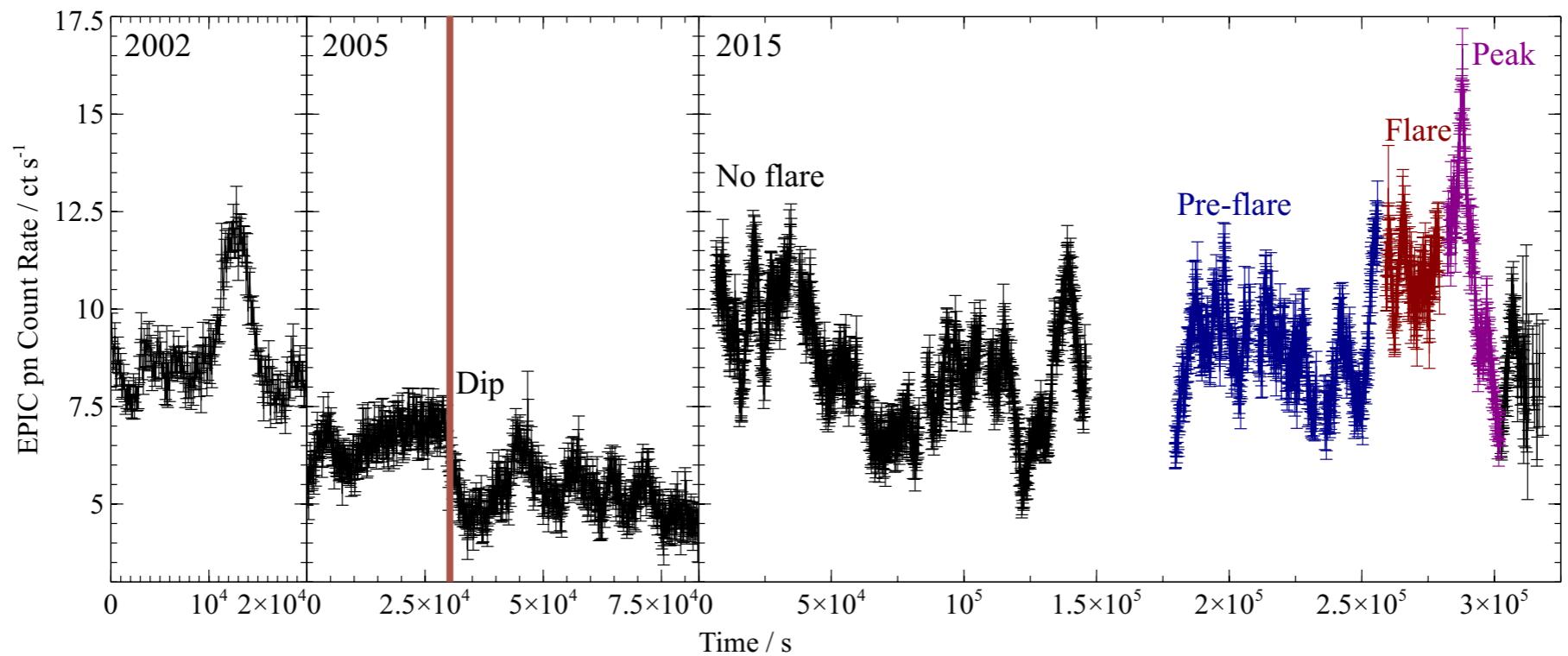


Pre-flare

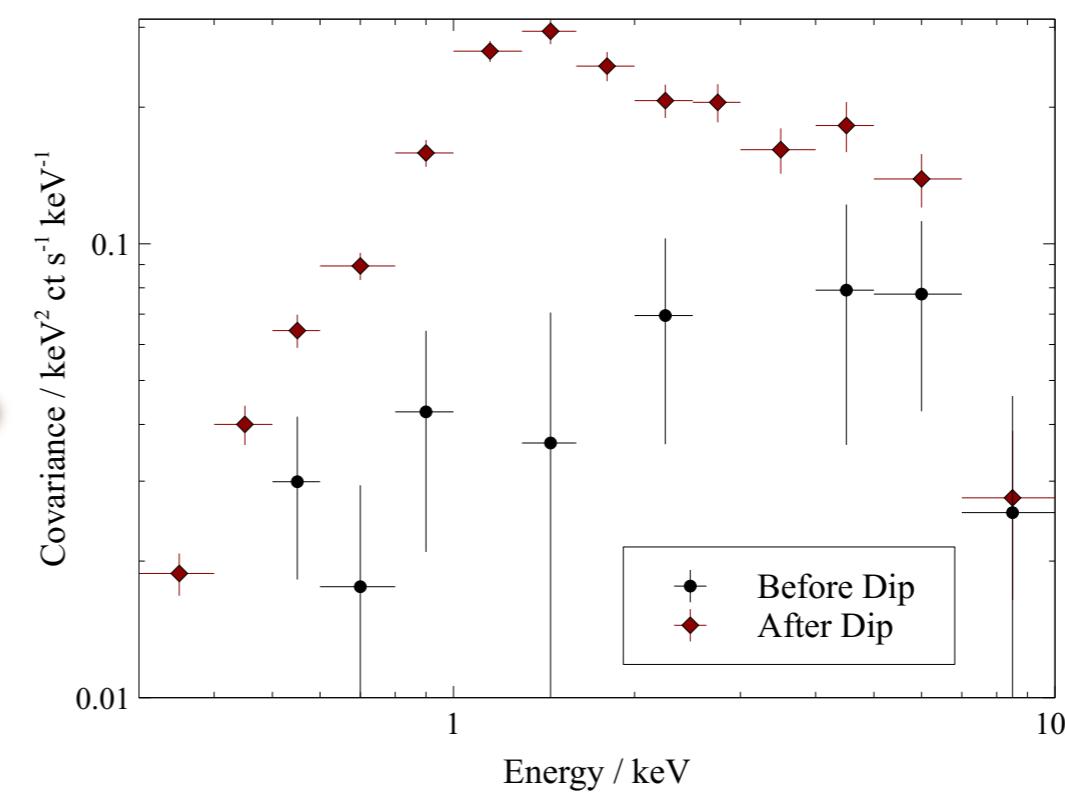
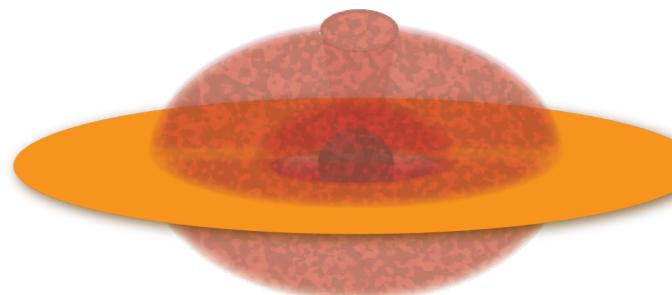


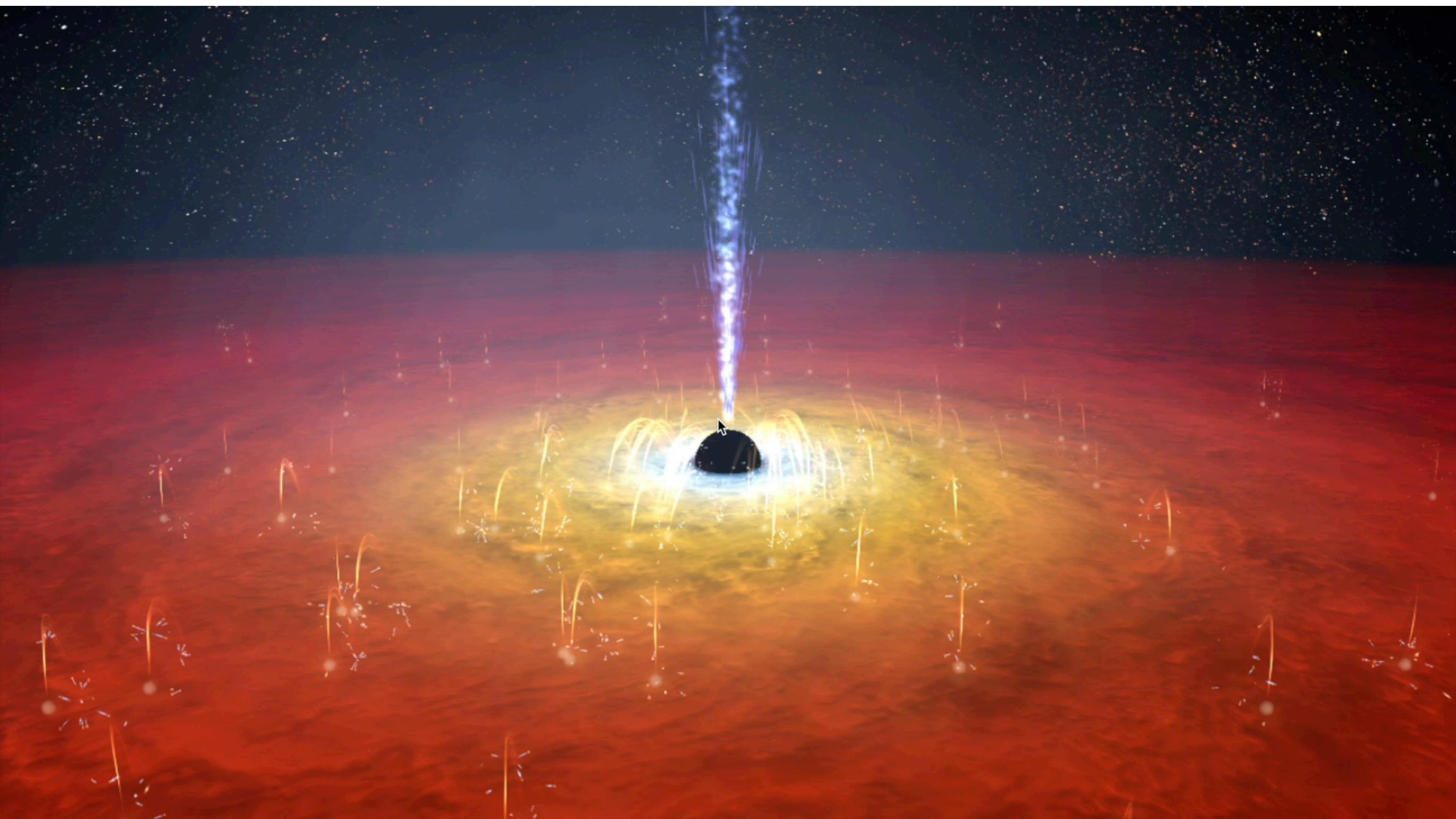
Peak

More From the Disc Corona

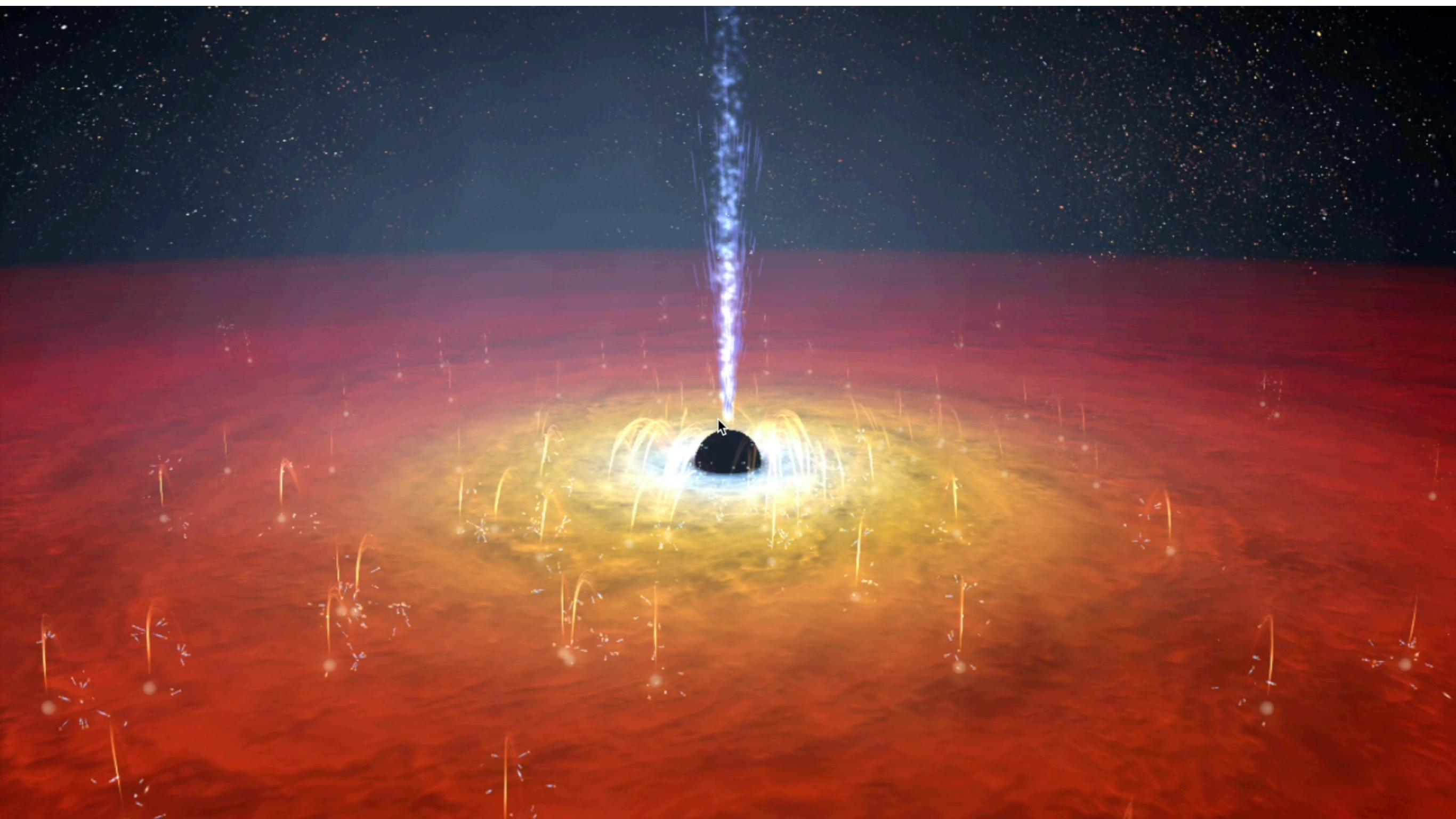


Low Frequency
 $4 \times 10^{-4} - 1.5 \times 10^{-3}$ Hz





Animation: Niel Brandt & Nahks Tr'Ehnl



Animation: Niel Brandt & Nahks Tr'Ehnl

Summary

dan.wilkins@stanford.edu

- X-ray reflection and reverberation reveals structure right down to the innermost stable orbit and even the event horizon
- Time-averaged spectra and low frequency variability are well-described by a corona extended over the disc but reverberation lag-energy spectra require upward propagation through a collimated corona
- Starting to see evidence of a persistent collimated core within an extended corona – the base of a (failed) jet?
- Flaring seen to originate in the disc corona which becomes more turbulent before the count rate increases, then flare passes to core