

*Young, thin, & on the move:*  
HETG Observations of SNRs

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Dan Dewey  
MIT Kavli Institute

# Overview

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- SNR Science
- SNRs with Gratings
- Three SNR-HETG examples:
  - E0102, Cas A, SN1987A
- For the Future

# SNR Science

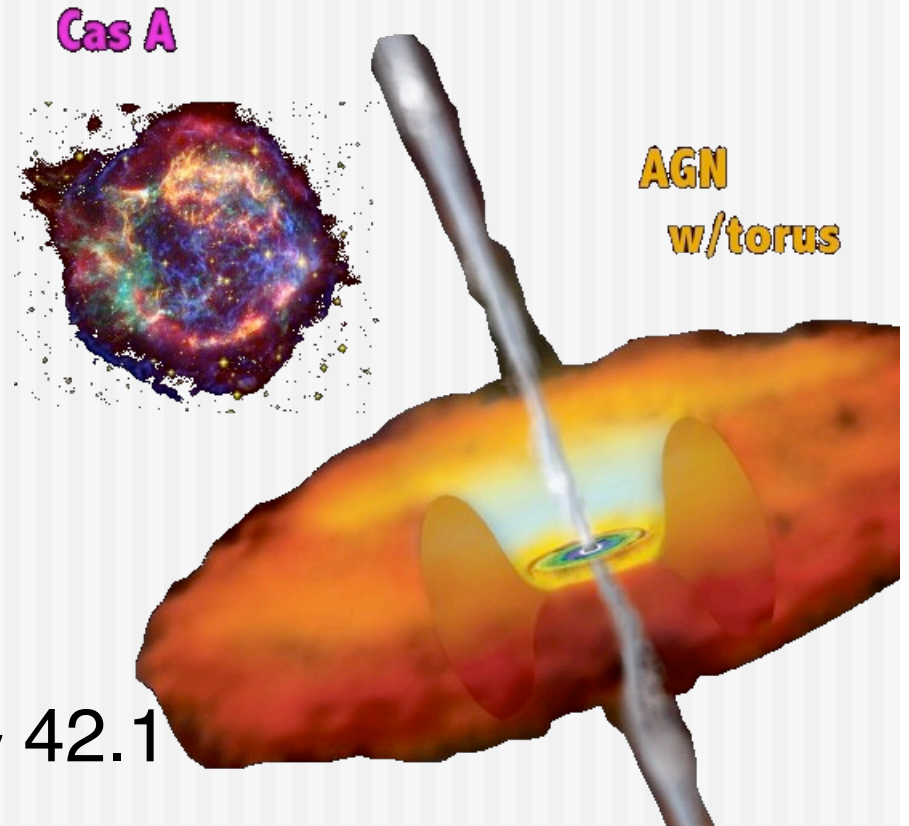
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- Progenitor and CSM
- SN Explosion - Core collapse & Type Ia
  - Mechanism, asymmetry, jets, mixing
  - Compact object formation
  - Nucleosynthesis
  - GRB connection
- ISM enrichment: dust, hot plasma, CRs

For fun...

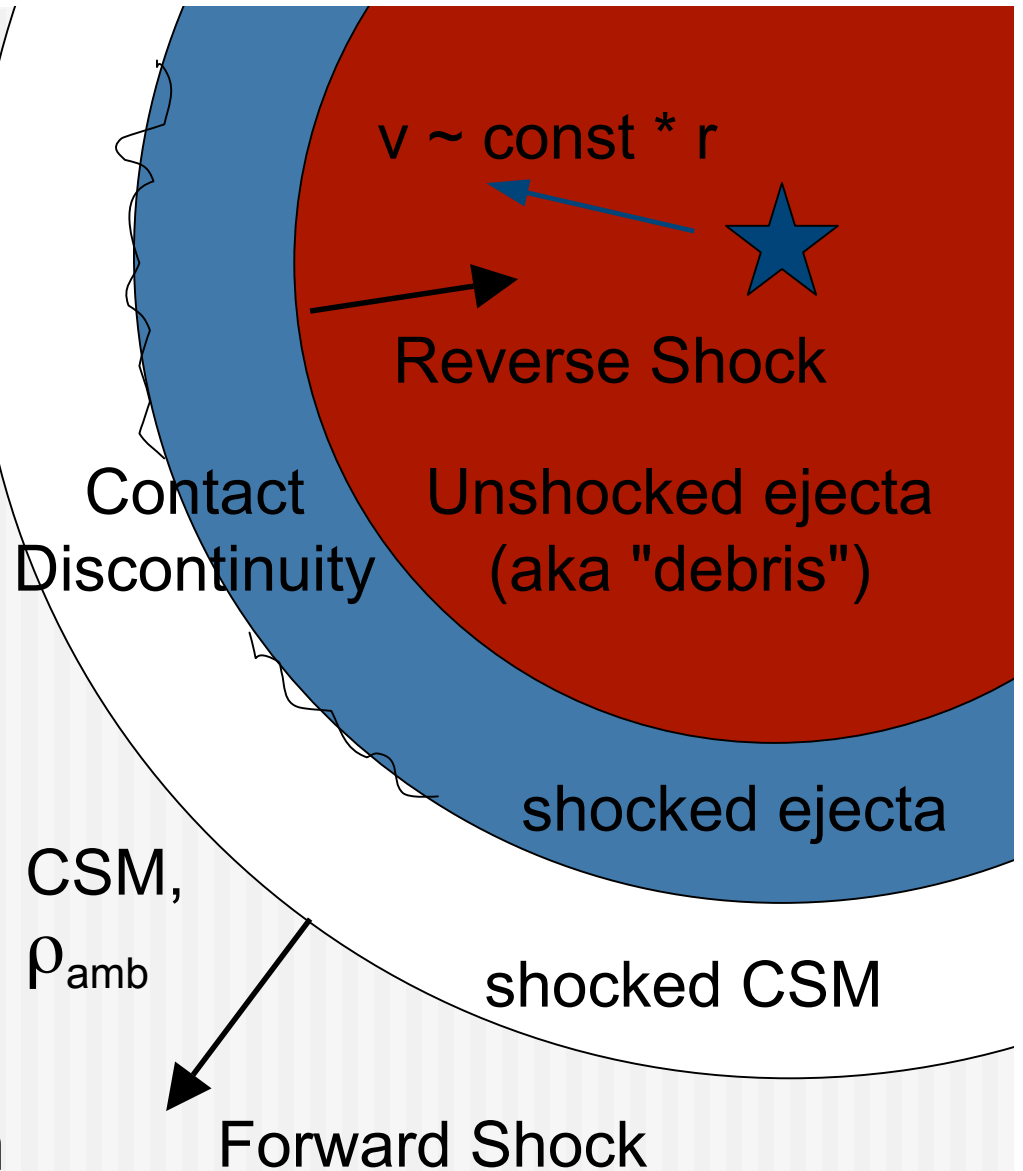
# SNRs and AGN

- Size - similar:
  - few parsecs scale
- Mass:
  - SNR: 10's  $M_{\text{solar}}$
  - torus:  $\sim 10^5 M_{\text{solar}}$
- $L_{\text{X-ray}}$  [ $\log_{10}(\text{erg/s})$ ]
  - SNR  $\sim 37.5$ , AGN  $\sim 42.1$



# SNR cartoon

- Sedov self- similar solution
- Reality adds:
  - R-T/mixing at CD
  - Non-uniform CSM
  - Asymmetric ejecta
  - Particle acceleration



# SNR Plasma

- Collision-less shocks: FS, RS, reflected
  - $V$  forward shock: 3k - 40k km/s
  - $V_{\text{ejecta}} @ \text{RS} \sim r_{\text{RS}} / \text{age}$   
SN1987A=7800; CasA=4800; E0102=2800 [km/s]
- NEI evolution, cooling
  - Electron & ion heating at shocks

# SNR Plasma, cont.

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- Shock-cloud interactions
  - Non-uniform CSM
  - Thermal conduction effects
- Particle acceleration & Cosmic rays
  - Efficiency and effect on hydrodynamics

# SNR Plasma Emission

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- Emission across wide  $\lambda$ -range
- Thermal
  - shocked ejecta; shocked CSM; dust
- Non-thermal
  - Synchrotron,  $T_e$  max, magnetic fields
  - non-thermal brems., inverse Compton
- Photoionized (unshocked, precursor)



# SNRs with Gratings

- High(er) spectral resolution than CCDs
  - Line separation
    - From nearby lines: He-like triplets
    - From continuum: low-E lines: O VIII, Fe-L, Ne
  - Doppler velocities can be measured
- Extended and/or multiple source
  - More complex analysis
  - Focus on *line emission*

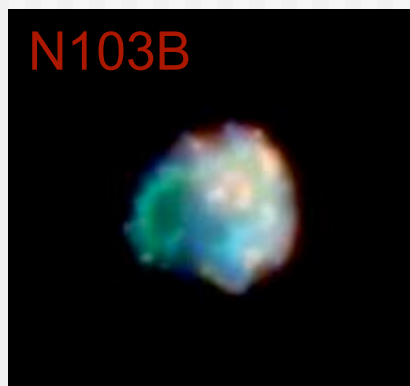
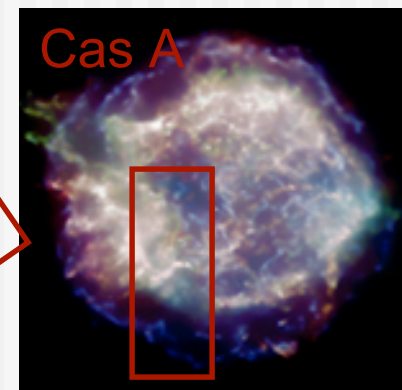
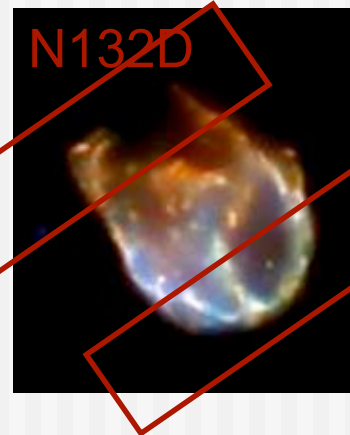
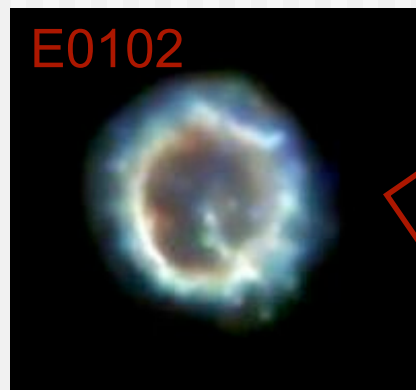


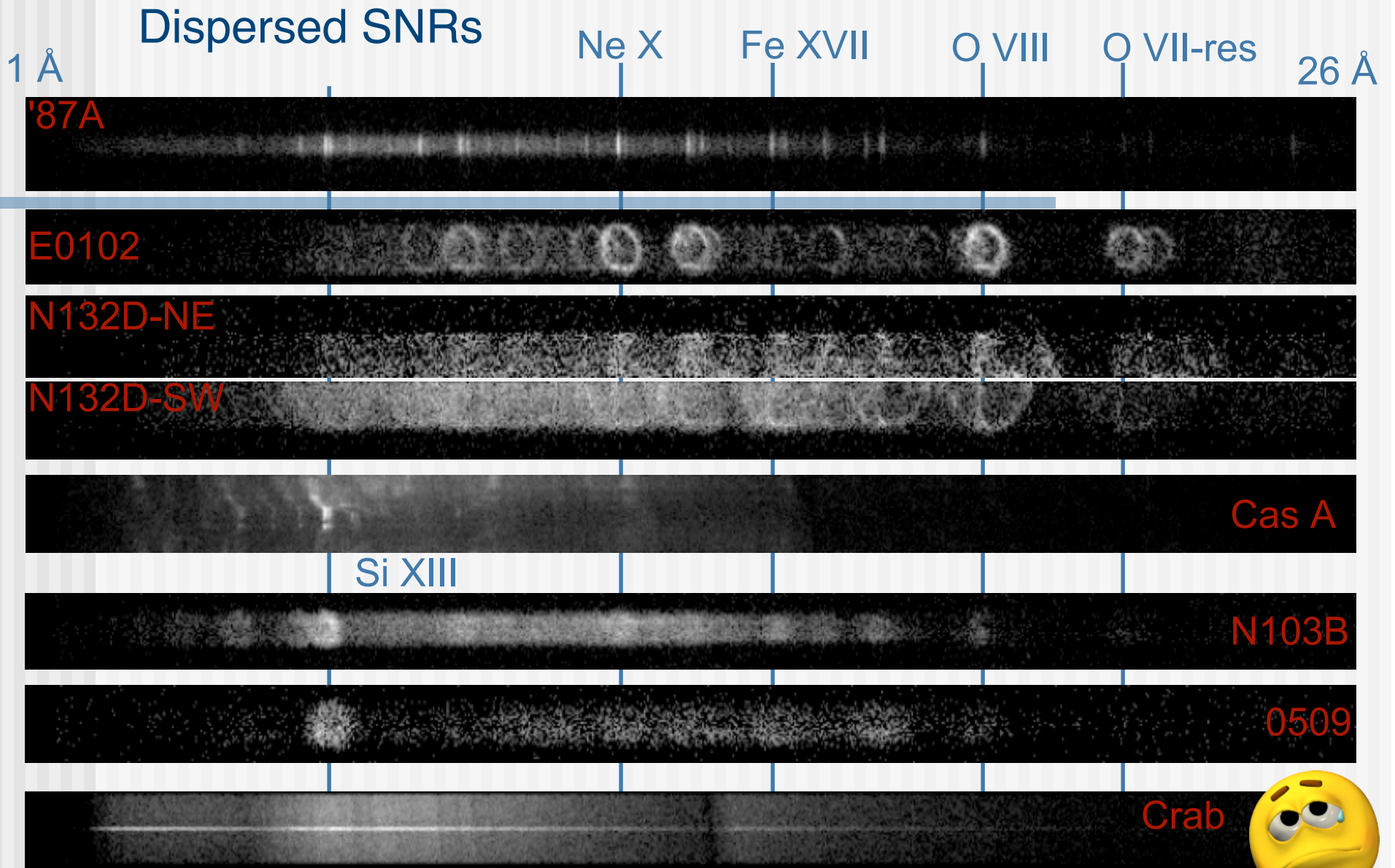
# SNRs observed w/HETG

<i>SNR</i>		<i>Age (yr) [Date]</i>	$n_e$ *	$ v_{bulk} $ *
SN 1987A	cc	20 [1987]	~ 800	150--500
E0102	cc	~ 1000	~ 1	~ 2000
N132D	cc	~ 2500 -- 3400		
Cas A	cc	~ 337 [~1670]	~ 100	~ 4000
N103B	la	~ 860		
0509-67.5	la	~ 400 [~1600]		
( Crab )	cc	953 [1054]		

\* Values are for the bright X-ray emitting plasma.

# SNRs w/HETG: images

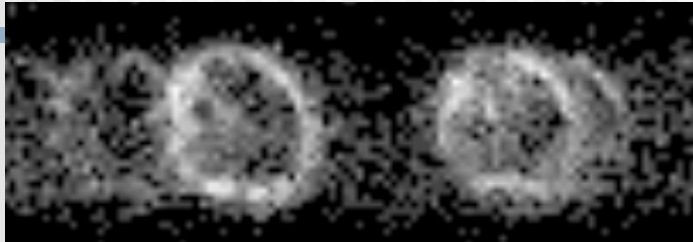




Not as useful for continuum

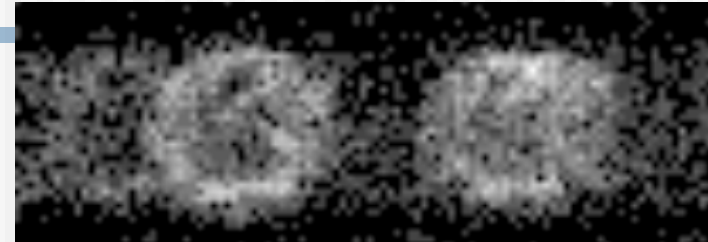
# E0102 w/HETG

MEG -1



Data

MEG +1



Model



Distortions between the + and - orders of the Ne lines are qualitatively reproduced with a cylindrical geometry for the ring emission.

[Dewey 2002](#); [Flanagan et al. 2004](#).

Model geometry



# Cas A w/HETG

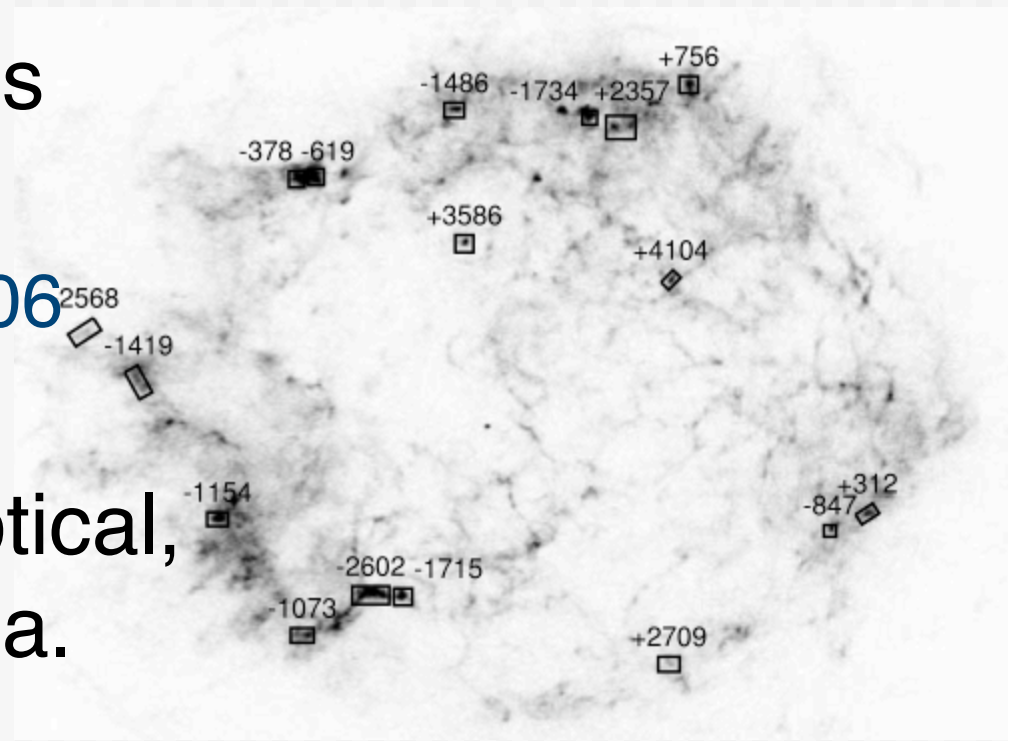
- Doppler velocities

  - 3D spatial info

  - Lazendic et al. 2006

- Combine with optical, IR, and radio data.

  - DeLaney 2007

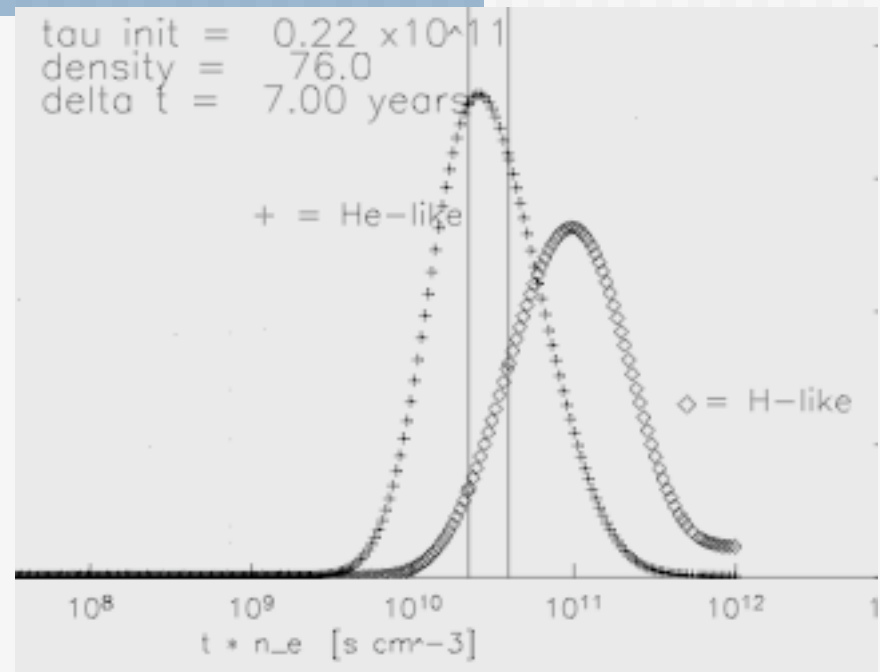
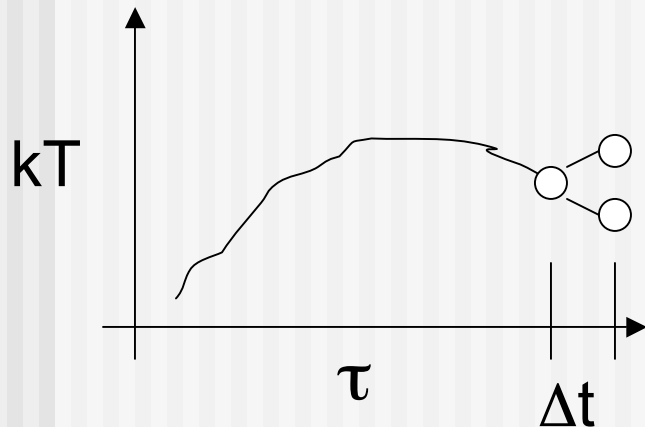


# Cas A w/HETG, cont.

## ■ Line ratios

→ NEI parameters

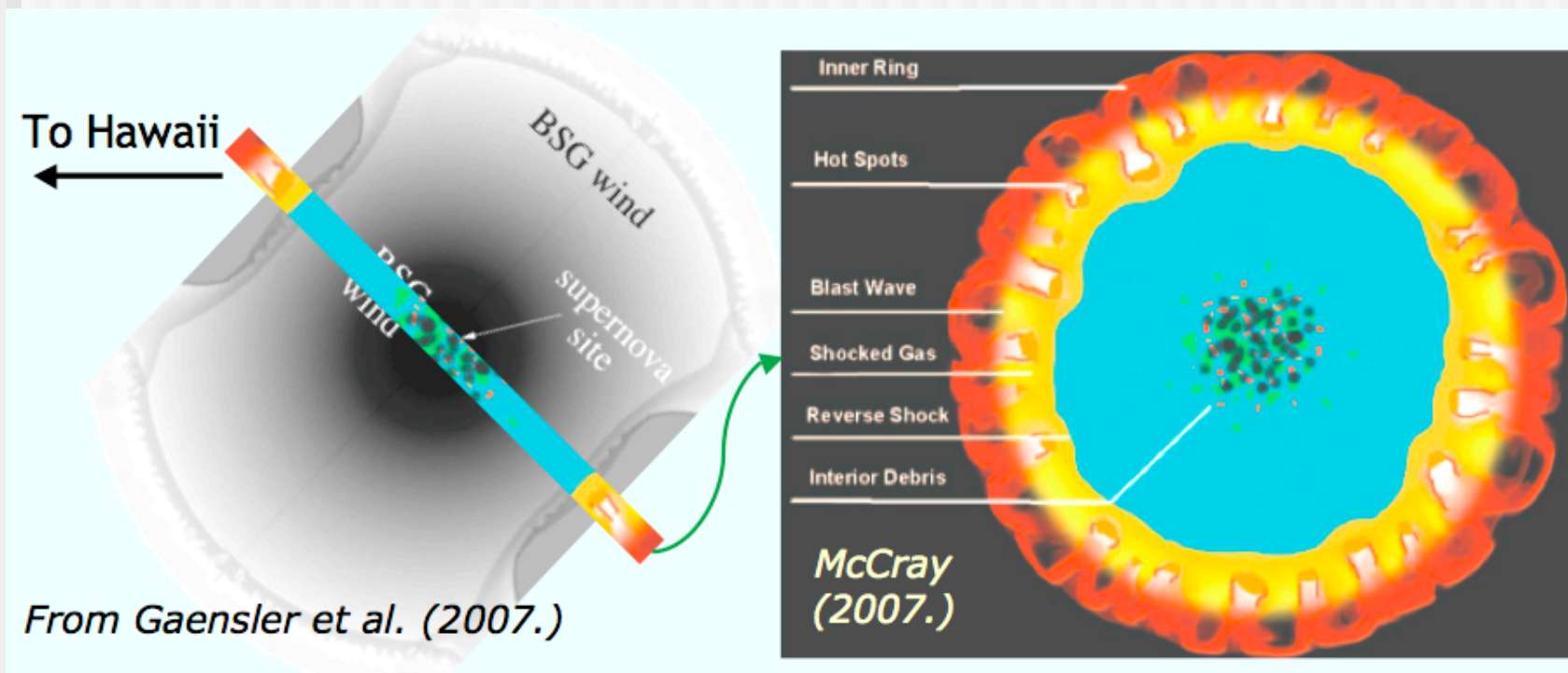
$\Delta$  params /  $\Delta t$  :



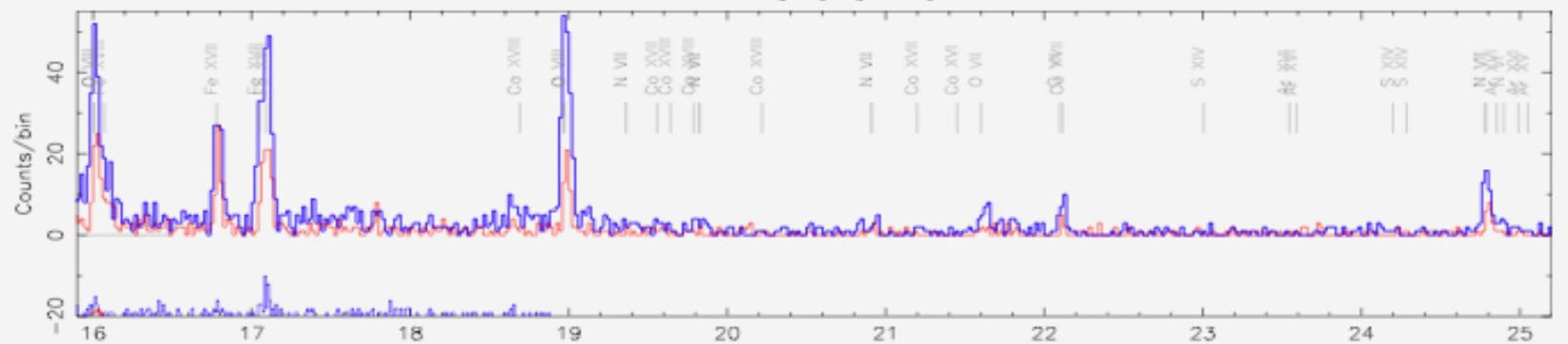
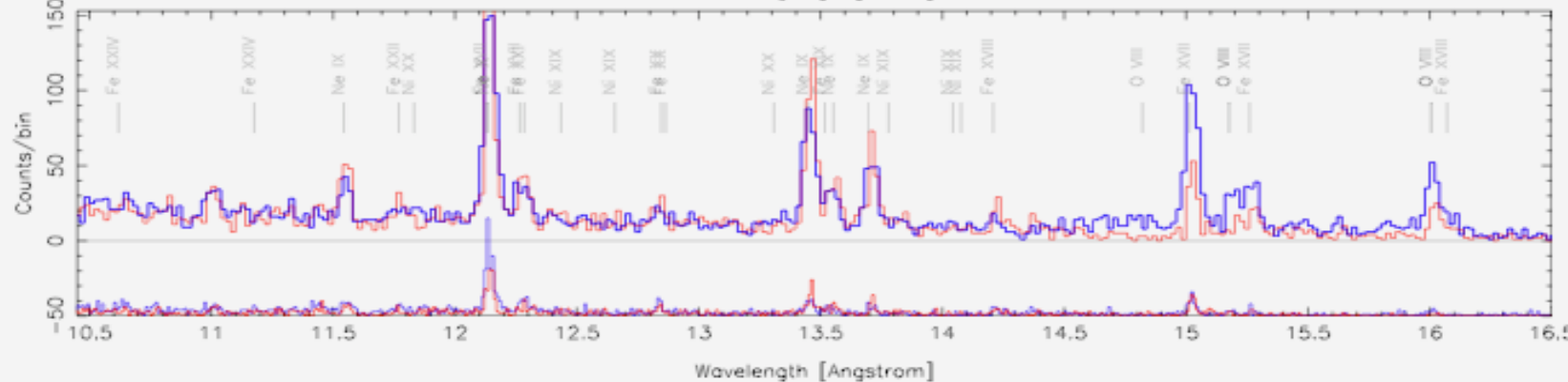
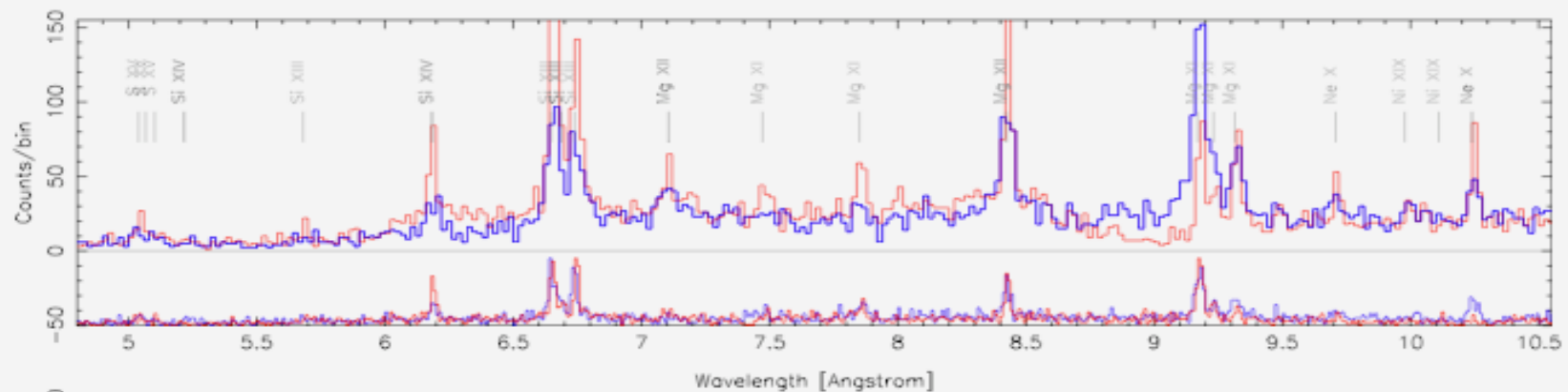
Change in plasma parameters over many years baseline can constrain models.

# SN 1987A w/HETG

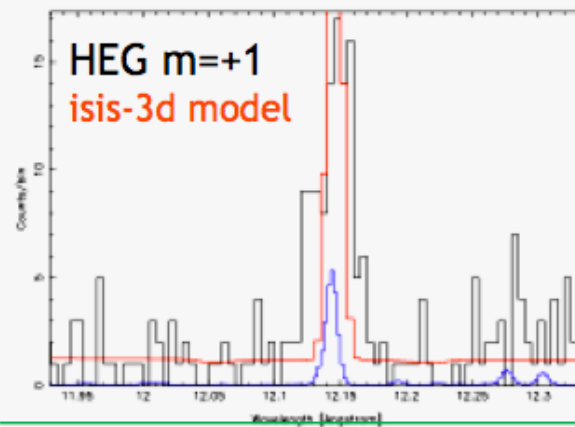
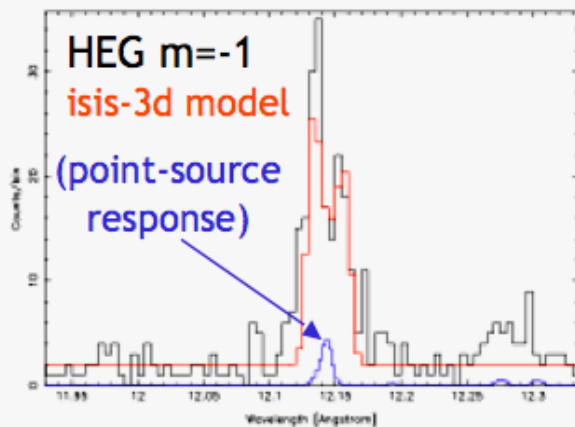
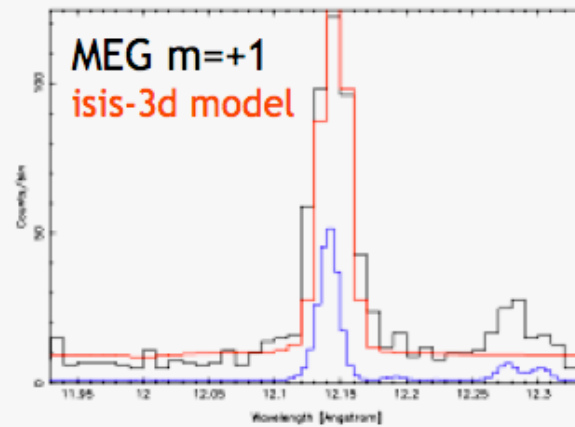
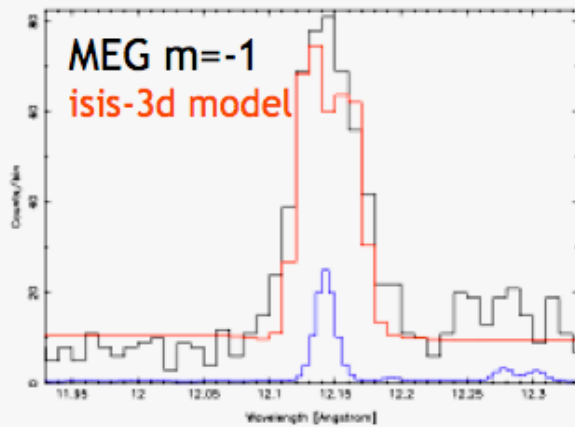
- The blast wave is hitting a dense equatorial ring, tilted 45 degrees to our line of sight



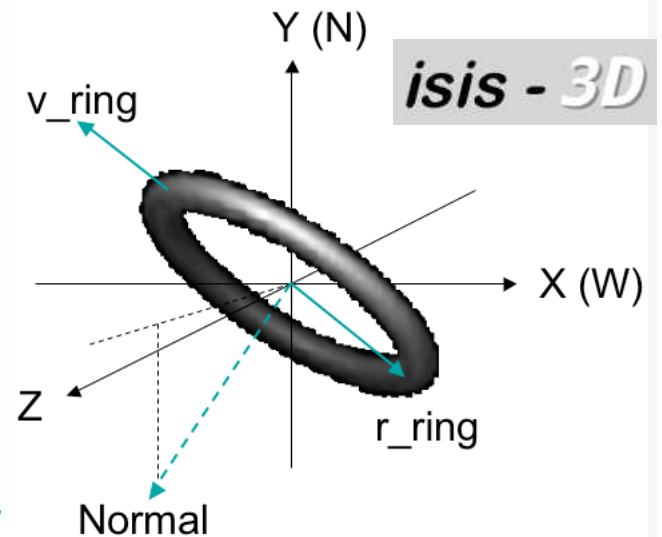




# SN 1987A w/HETG, cont.

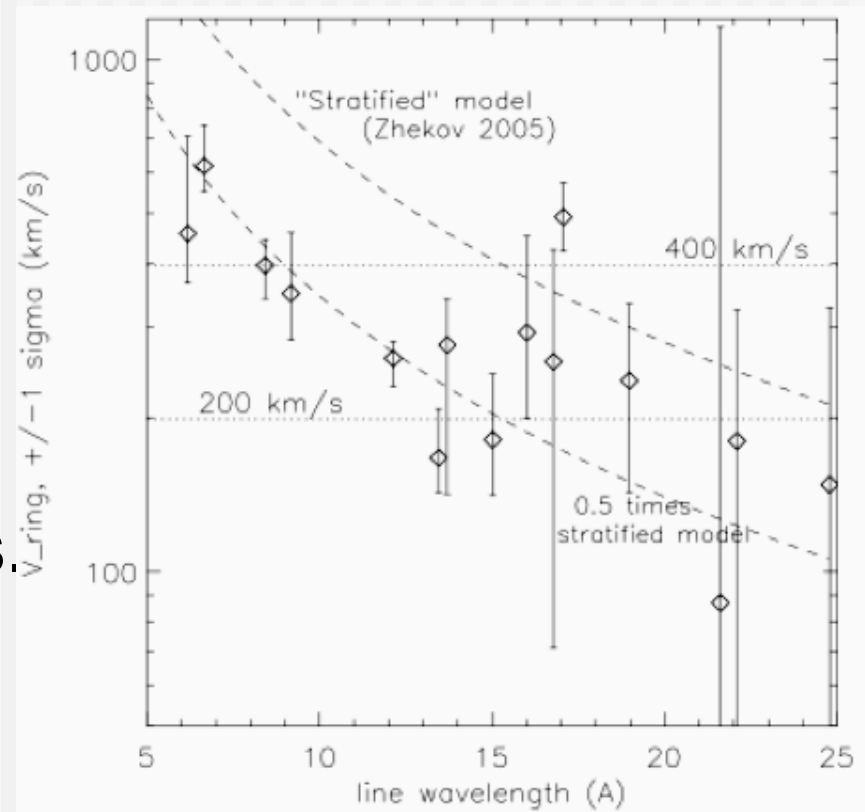


A 3D model is used to fit the HETG line shapes; a key free parameter is  $V_{\text{ring}}$ .



# SN 1987A w/HETG, cont.

- $V_{\text{ring}}$  varies with  $\lambda$ .
- Seen in LETG/ACIS data (McCray PI.)
- Zhekov et al. 2005
- Range of shock (and hence bulk) velocities.
- Higher ionization with faster velocities.



# For the future

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- Can get information from HETG-SNR observations, so ...
- Further HETG observations...
  - O-rich SNR
- Future X-ray instruments for SNR
  - Long slit spectrometer

# Oxygen-rich SNRs

In SMC/LMC:

- **SN 1987A** (LMC) ?O-rich?
- **N132D** (LMC)
- **SNR 0540-69.3** (LMC)
- **E0102.2-7219** (SMC)
- SNR 0103-72.6 (SMC)
- IKT6, SNR 0049-73.6 (SMC)

Galactic & extra-Galactic:

- **Cas A**
- Pup A
- **G292+1.8**
  
- SNR in NGC 4449
- SN 1957D in M83

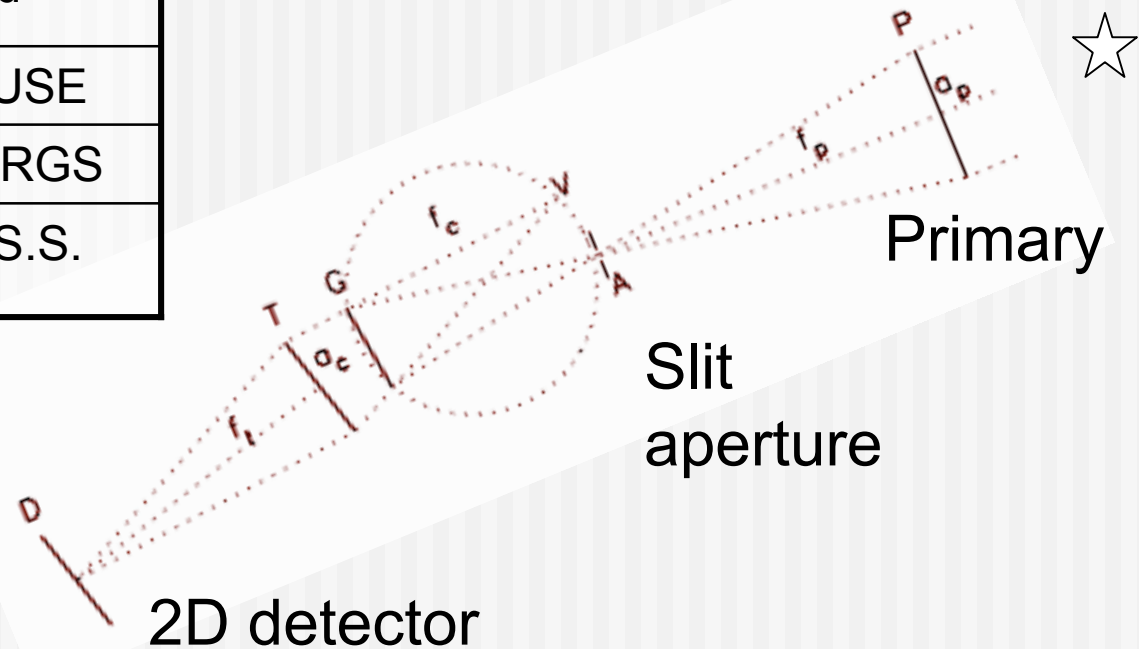
Underline = Grating observed

**Bold** = Deep ACIS data taken

# Multi- $\lambda$ & High Resolution

Radio	VLA
IR	Spitzer-IRS
NIR, Opt.	Ground-based
UV, EUV	HST-STIS, FUSE
X-ray	HETG, LETG, RGS
$\gamma$ -ray - TeV	GLAST, H.E.S.S.

Long-slit instruments  
from IR to ... X-ray ?



Willingale 2005  
SPIE Vol. 5900

# Thanks to

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## Claude Canizares and MKI colleagues

Hydra project:

<http://space.mit.edu/hydra>

## Dick McCray and '87A collaborators

See the program of the Aspen '87A workshop:

<http://astrophysics.gsfc.nasa.gov/conferences/supernova1987a/>

## Kelly Korreck and many SNR enthusiasts

See the Endpoints and Interactions workshop:

<http://www.cfa.harvard.edu/~kkorreck/snrconf.htm>