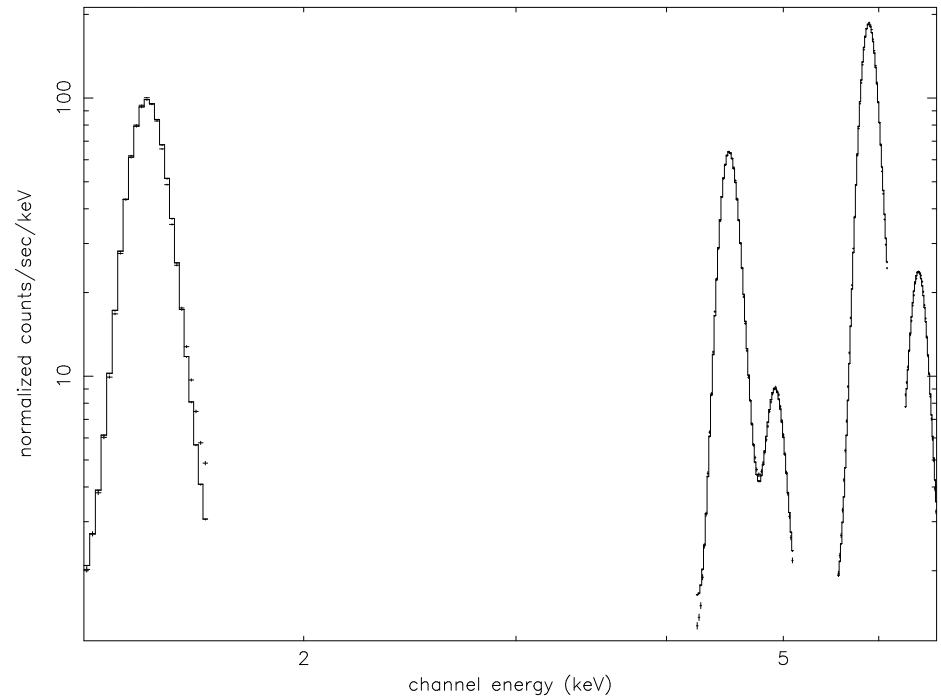
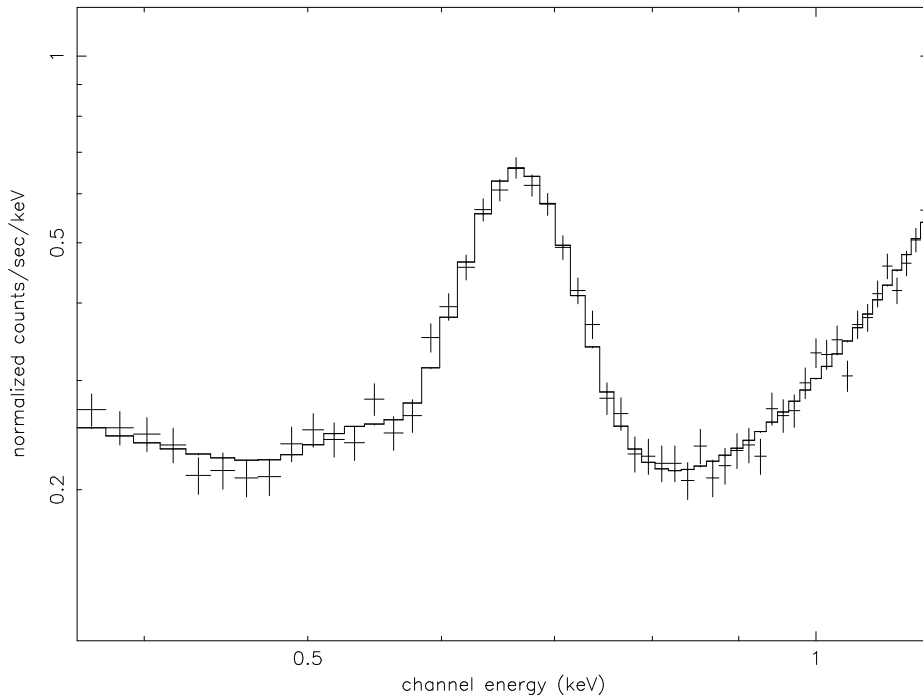


# **SPATIAL STRUCTURE OF ACIS CONTAMINATION**

# Scope

- **Flux ratio for Mn-K and (Mn+Fe)-L gives  $\tau$  near 700 eV**
- **Results presented many times since 2002**
- **This talk:**
  - **final report (hopefully)**
  - **science tests using A1795 pointings**

# Data



- ECS data over 3 months (t-dependence) or 1 year (x,y-dependence)

Observed cnt/s

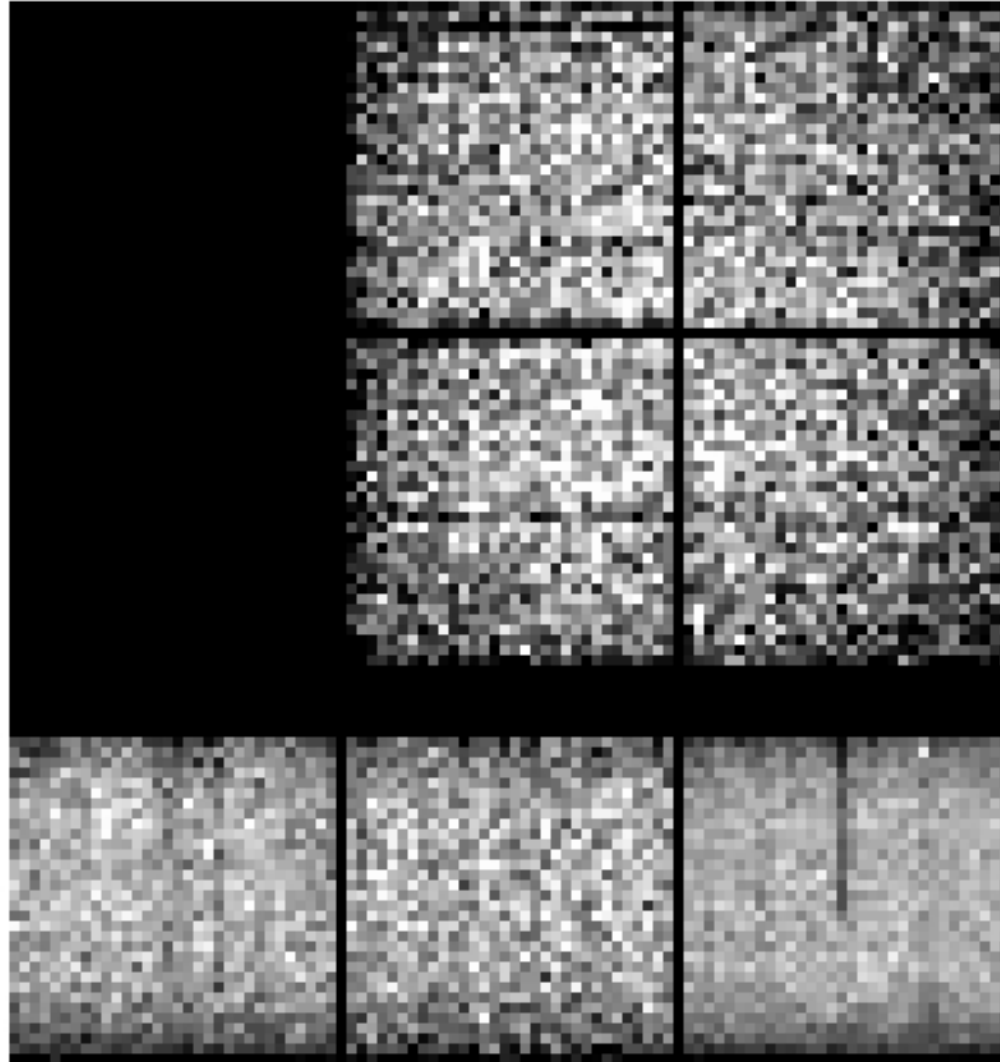
- $\text{flux} = \frac{\text{Observed cnt/s}}{\text{QE} \times \text{OBF transmission} \times \text{QEU}}$

$$\tau_L = -4.687 - \log(f_L/f_{\text{Mn-K}\alpha}) \quad \text{for S3}$$

$$\tau_L = -4.925 - \log(f_L/f_{\text{Mn-K}\alpha}) \quad \text{for ACIS-I and S2}$$

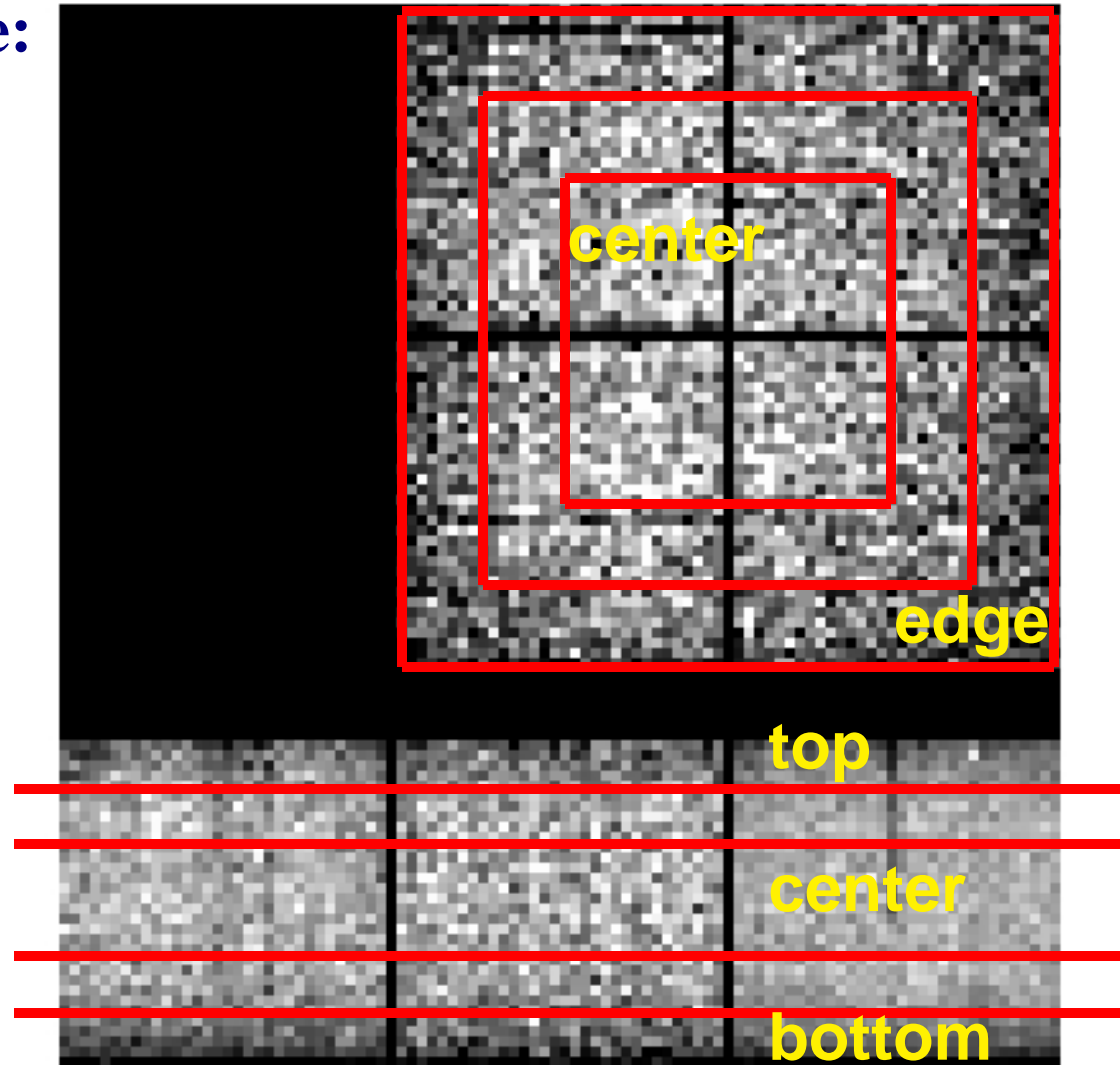
# Qualitative spatial distribution

ACIS, L-line image:

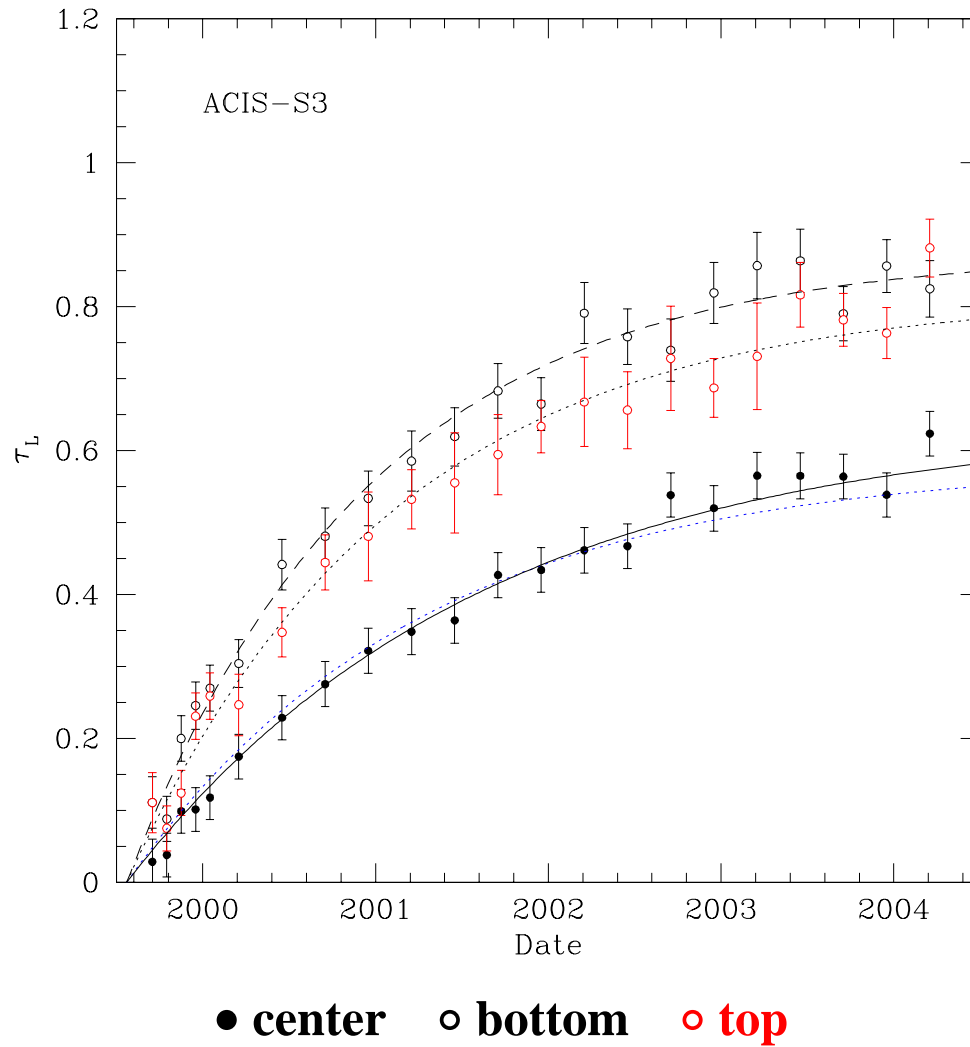


# Qualitative spatial distribution

ACIS, L-line image:



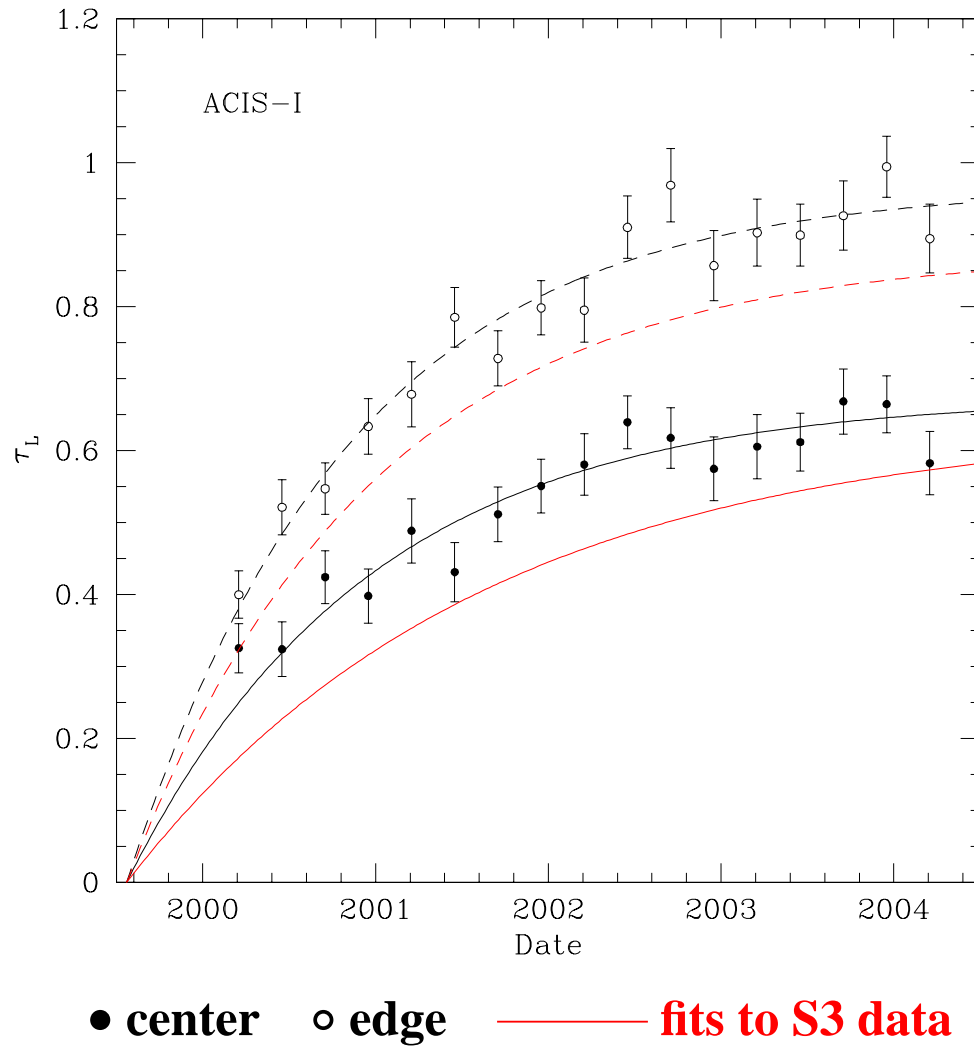
# Time dependance, S3



➤ fits of the form  $\tau = \tau_{\infty} (1 - \exp(-t/T))$

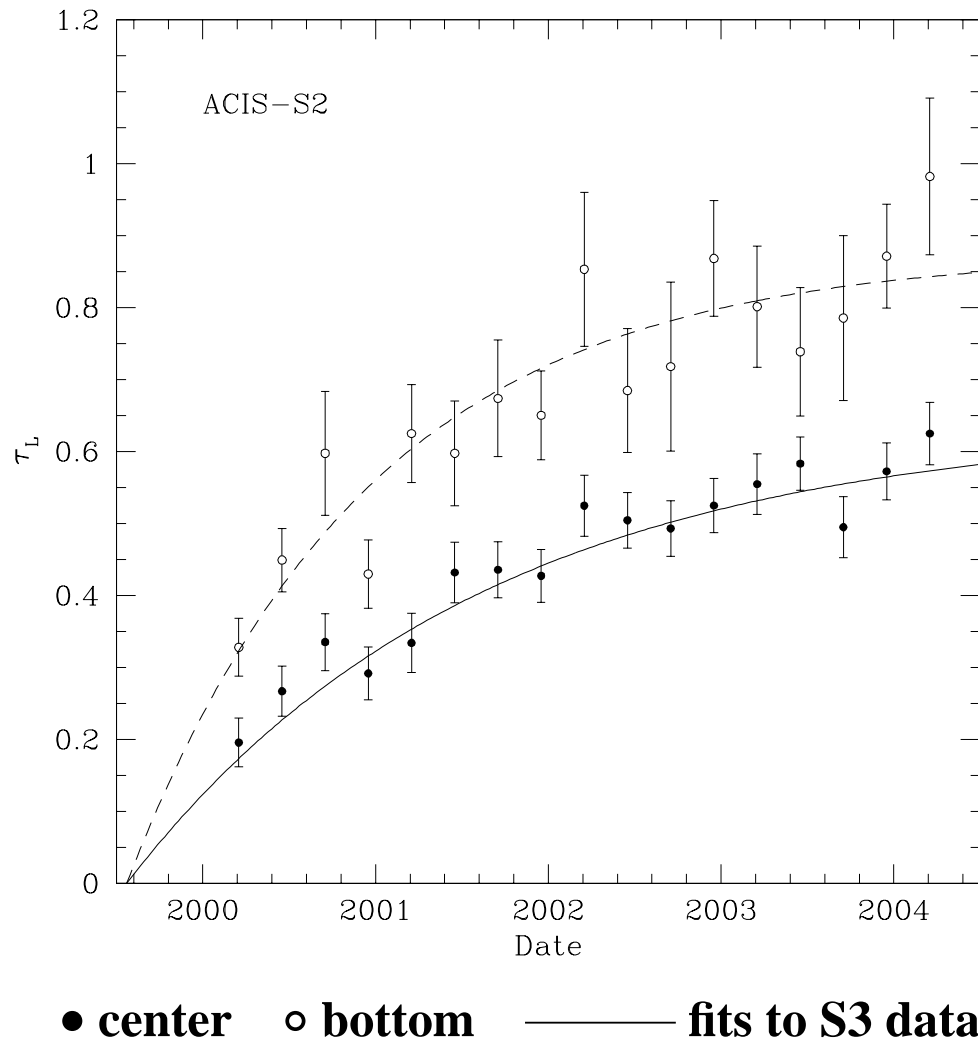
➤ ... — fit to chip-averaged data (C. Grant)

# Time dependance, ACIS-I



➤ stronger contamination in ACIS-I ( $\Delta\tau_L \simeq 0.08$ )

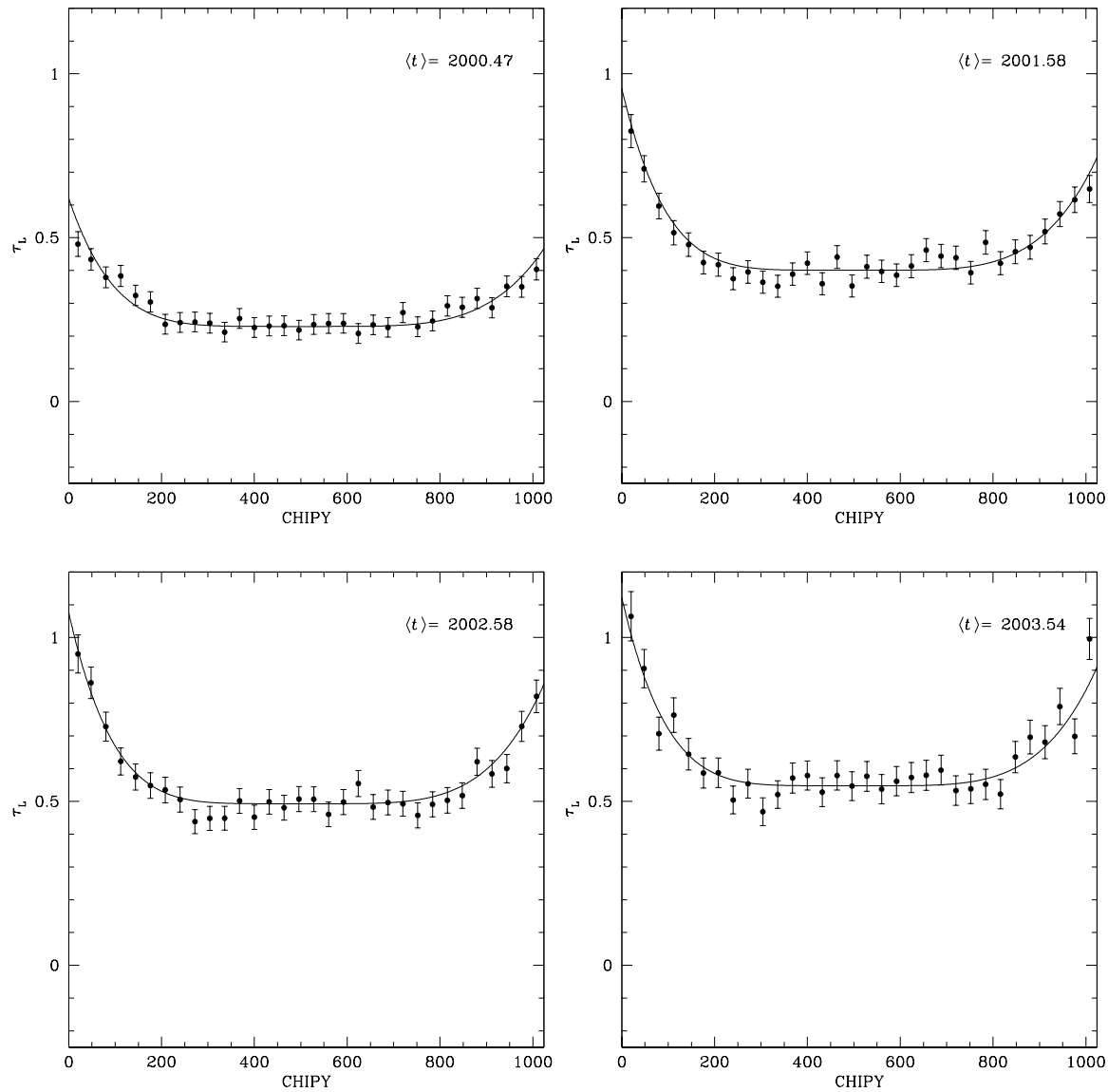
# Time dependance, S2



➤ (S3 vs. I difference not due to FI/BI crosscalibration)

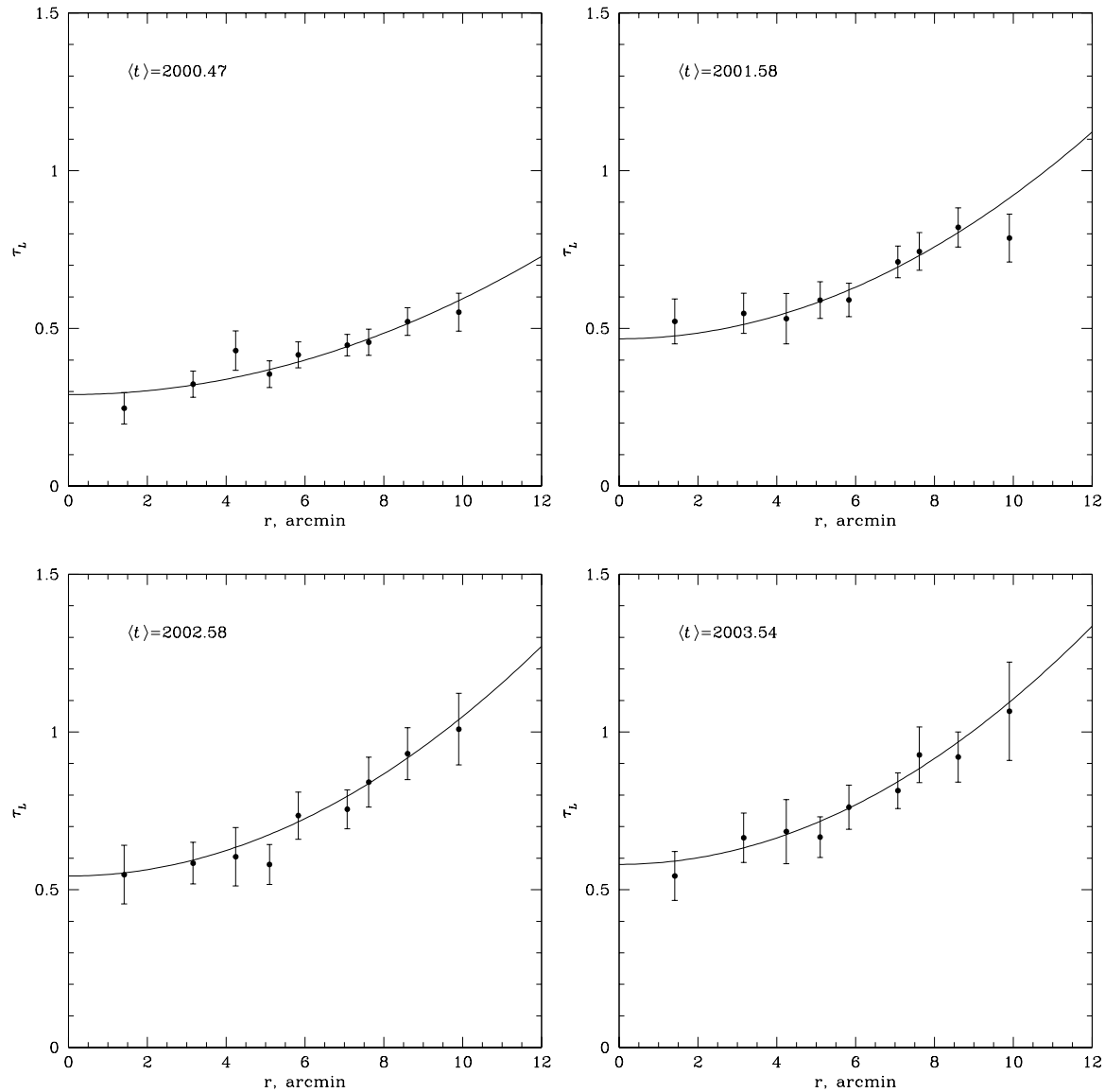


# Spatial dependence, S3



fits of the form  $\tau(y) = \tau_0 + \tau_1 |y - 512|^\alpha$ , with  $\tau_0, \tau_1$  from  $t$ -dependence

# Spatial dependence, S3



fits of the form  $\tau(y) = \tau_0 + \tau_1 r^{2.0}$ , with  $\tau_0, \tau_1$  from  $t$ -dependence

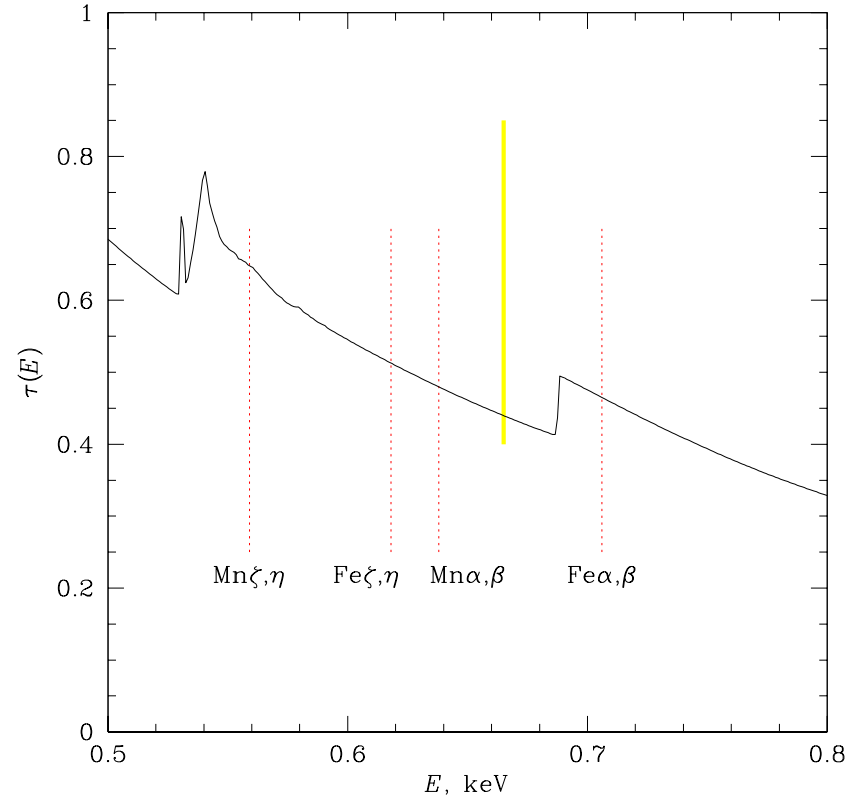
# How to go from $\tau_L$ to $\tau(E)$

- ECS gives model for  $\tau_L(x, y, t)$
- we want  $\tau(E, x, y, t)$
- procedure:  $\tau(E, x, y, t) = \tau_{\text{grat}}(E) \times A$   
 $A$  adjusted so that  $\tau_L(x, y, t)$  is reproduced

# Structure of the L-complex

From central energy of L-line:

Line group	$E$ , keV	$f(\text{FI})$	$f(\text{BI})$
Fe $\alpha, \beta$ .....	0.706	39%	33%
Mn $\alpha, \beta$ .....	0.638	54%	58%
Fe $\zeta, \eta$ .....	0.618	2%	3%
Mn $\zeta, \eta$ .....	0.559	5%	6%
<i>Empirical fit</i>			
X $\alpha$ .....	0.665	93%	
X $\zeta$ .....	0.535	7%	

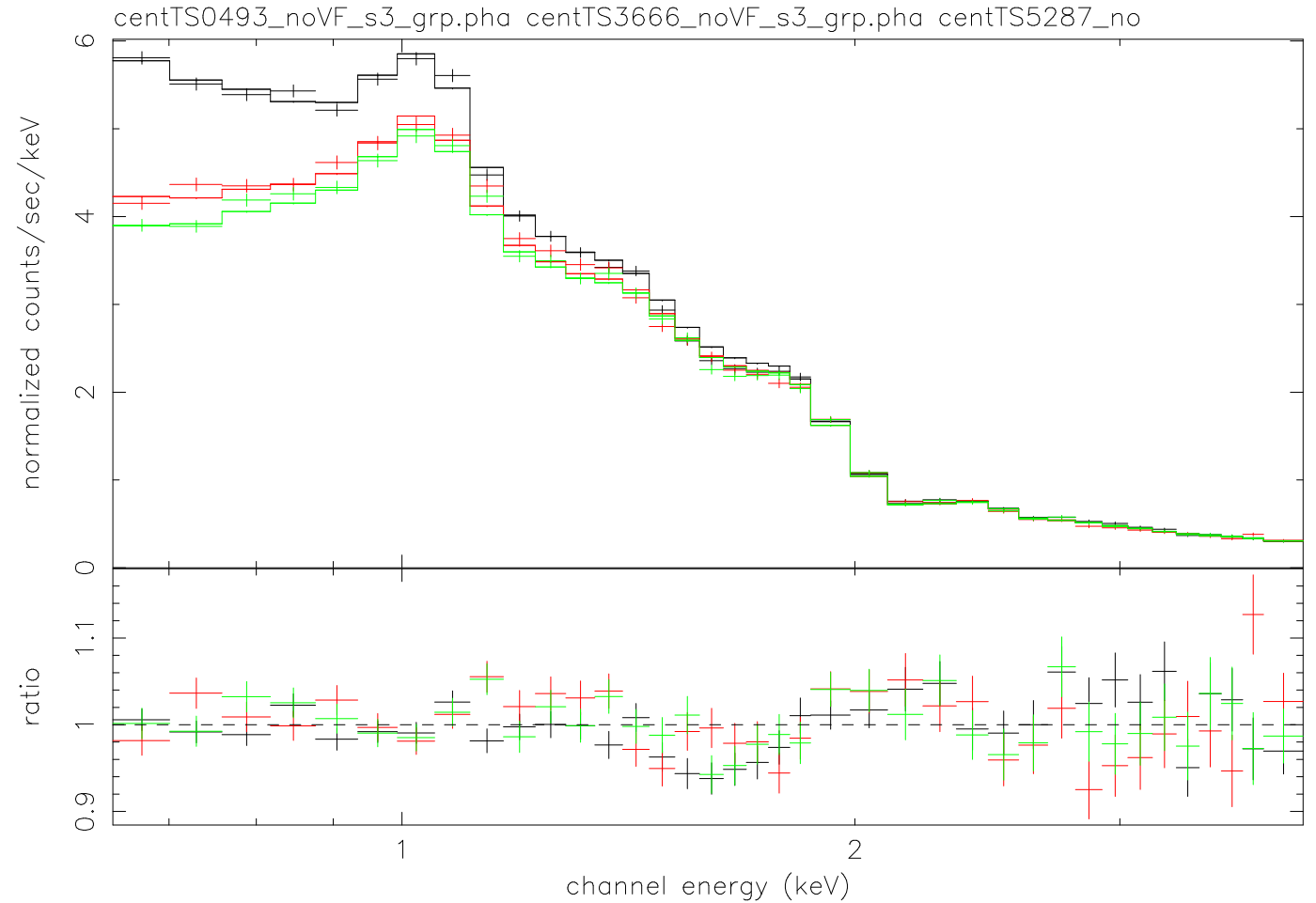
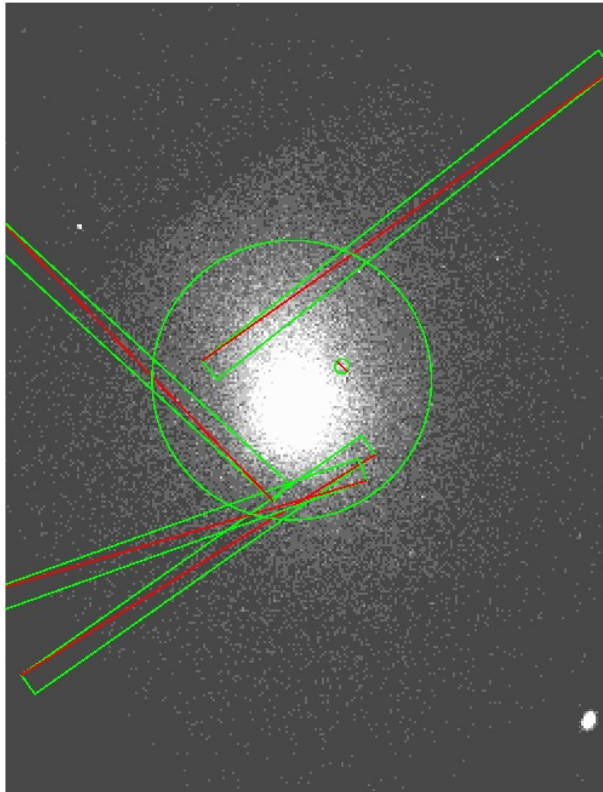


$$A = \frac{\tau_L(x, y, t)}{\sum f_i \tau_{\text{grat}}(E_i)}$$

# Final model

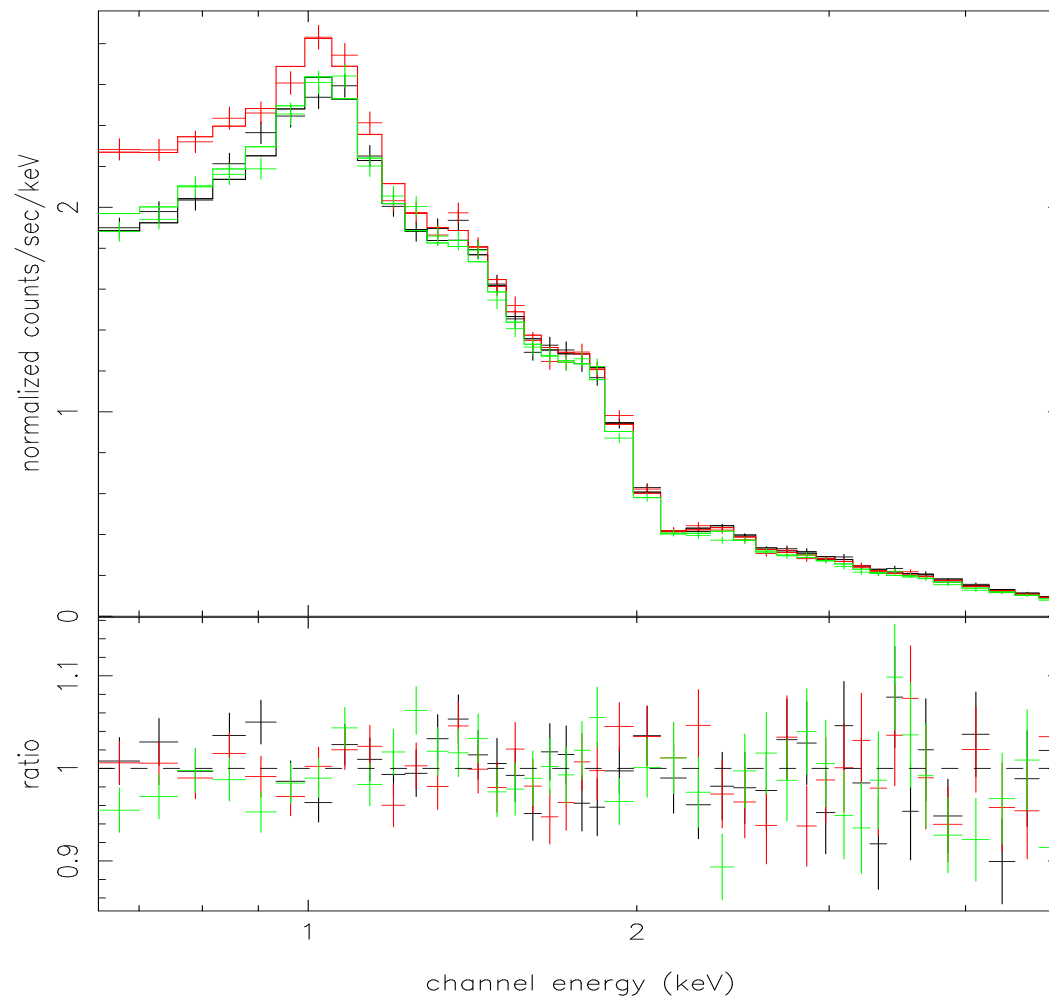
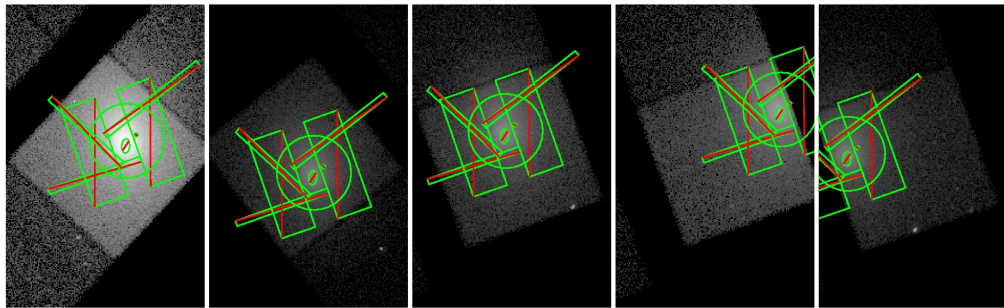
- 1)  $\tau(E)$  from grating measurements
  - 2)  $\tau_L(x, y, t)$  from ECS data
  - 3)  $\tau(E, x, y, t)$  from renormalization of  $\tau(E)$  to match  $\tau_L(x, y, t)$
- Model should be at  $E \gtrsim 0.6$  keV (above L-line)  
can be inaccurate near C-K edge until  $\tau(E)$  finalized
  - **TEST: multiple observations of A1795:**
    - 4 times in the center of S3
    - pointings to bottom & top of S3, center & edge of ACIS-I

# A1795: t-dependence in S3



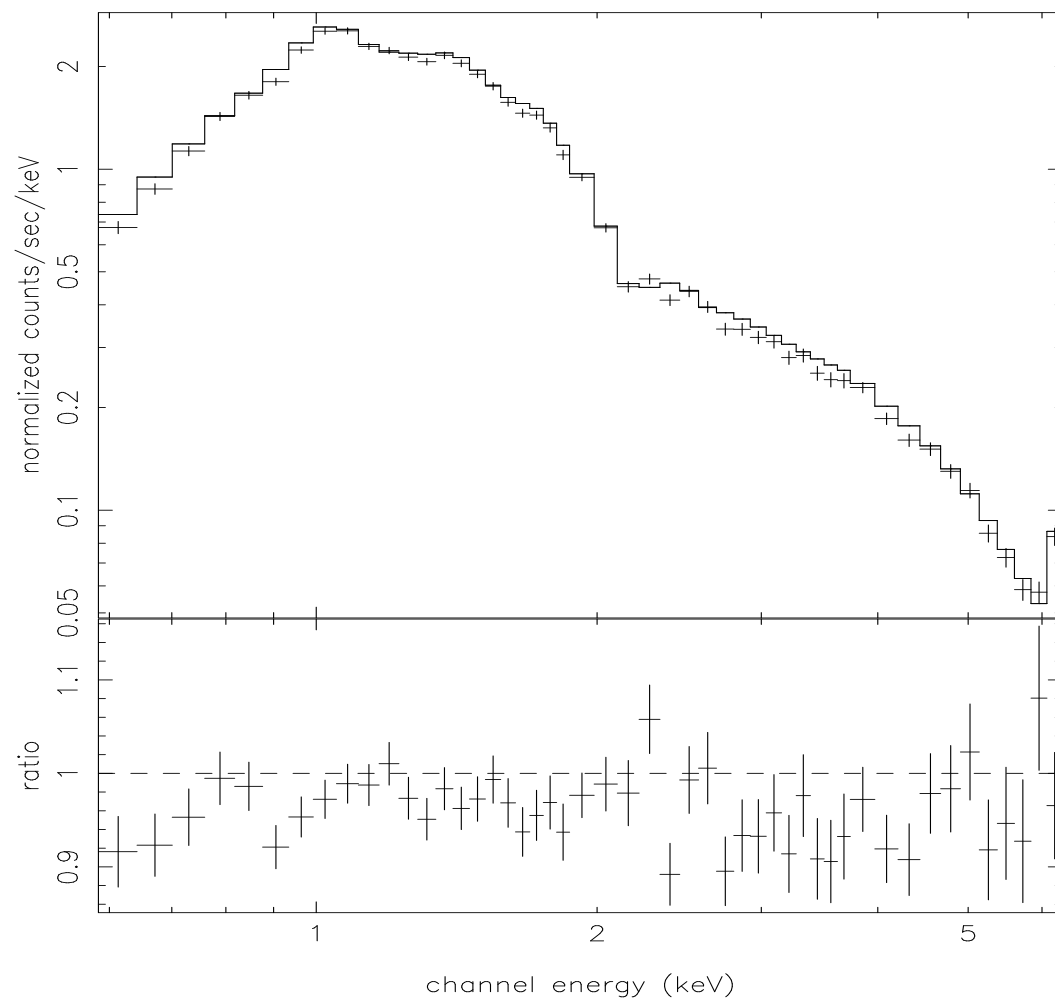
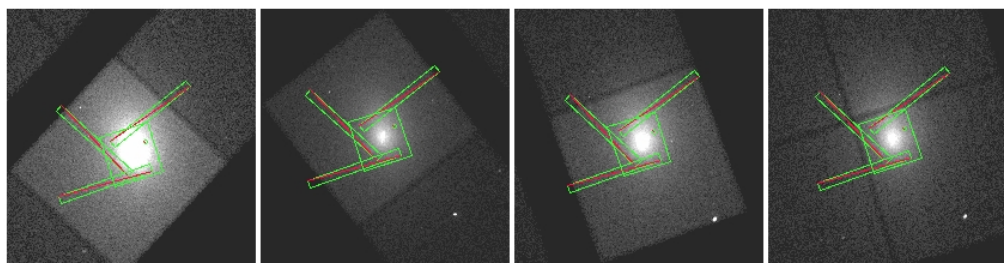
**< 3% residuals, time span 4 years**

# A1795: spatial dependence in S3



(reference fit from all pointings to S3-center)

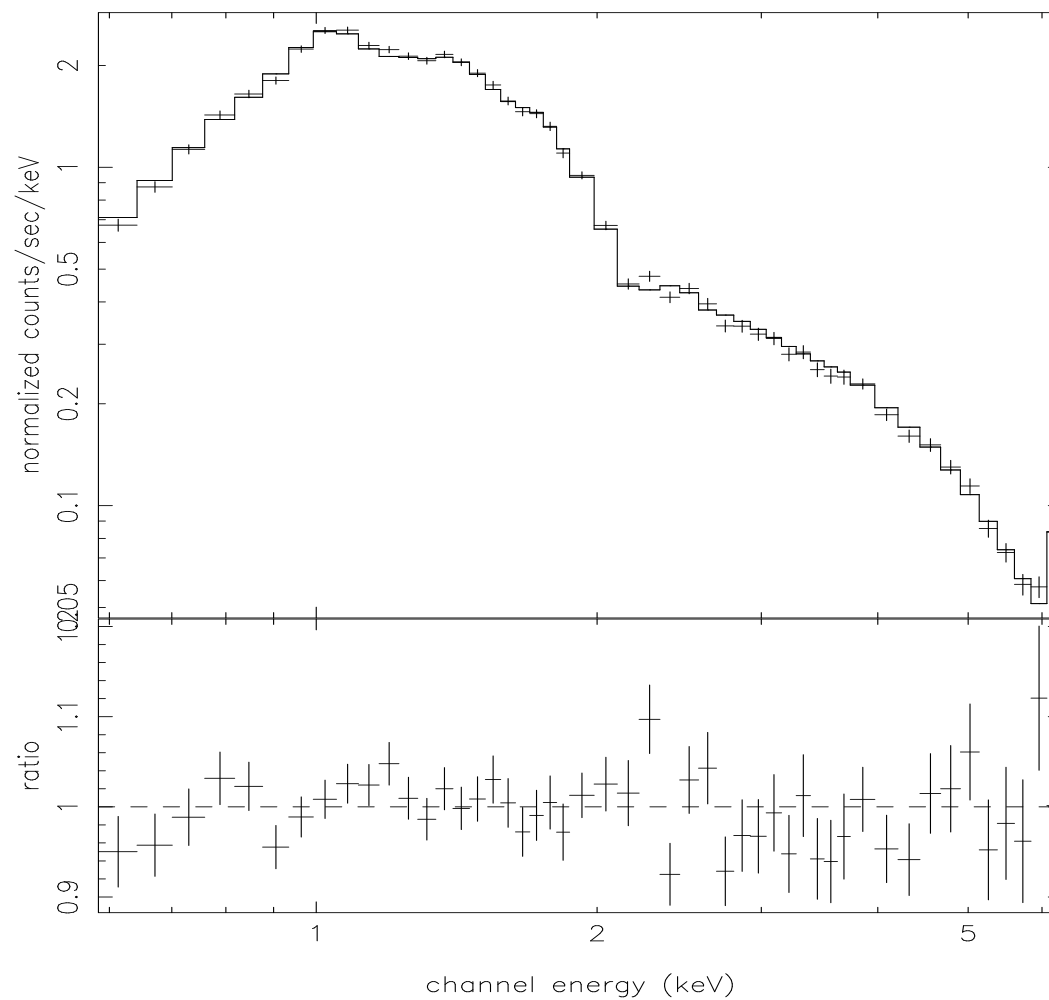
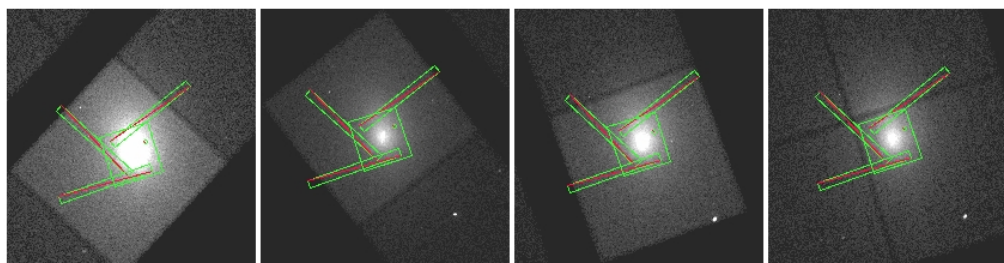
# A1795: cross-calibration between ACIS-S and ACIS-I



reference fit from all pointings to S3-center

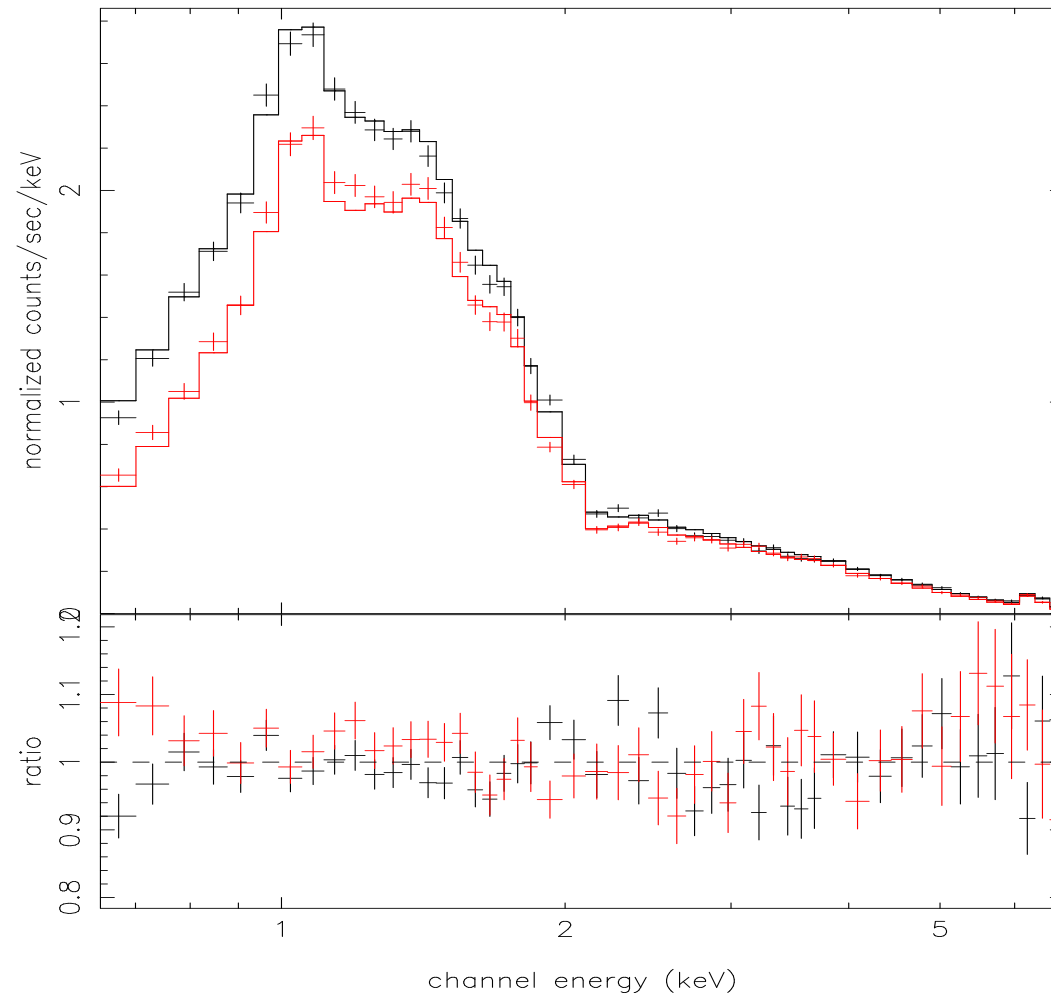
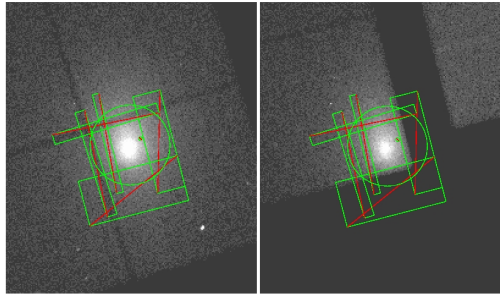


# A1795: cross-calibration between ACIS-S and ACIS-I



reference fit from all pointings S3-center, corrected for 3% dead area

# A1795: spatial dependence in ACIS-I



**(joint fit to center and edge data)**

# Conclusions

- **Accurate contamination model for ACIS imaging**  
(**< 2 – 3% uncertainties above 0.6–0.7 keV**)
- **(A1795 data also confirms cross-calibration of FI/BI QE and QEU)**