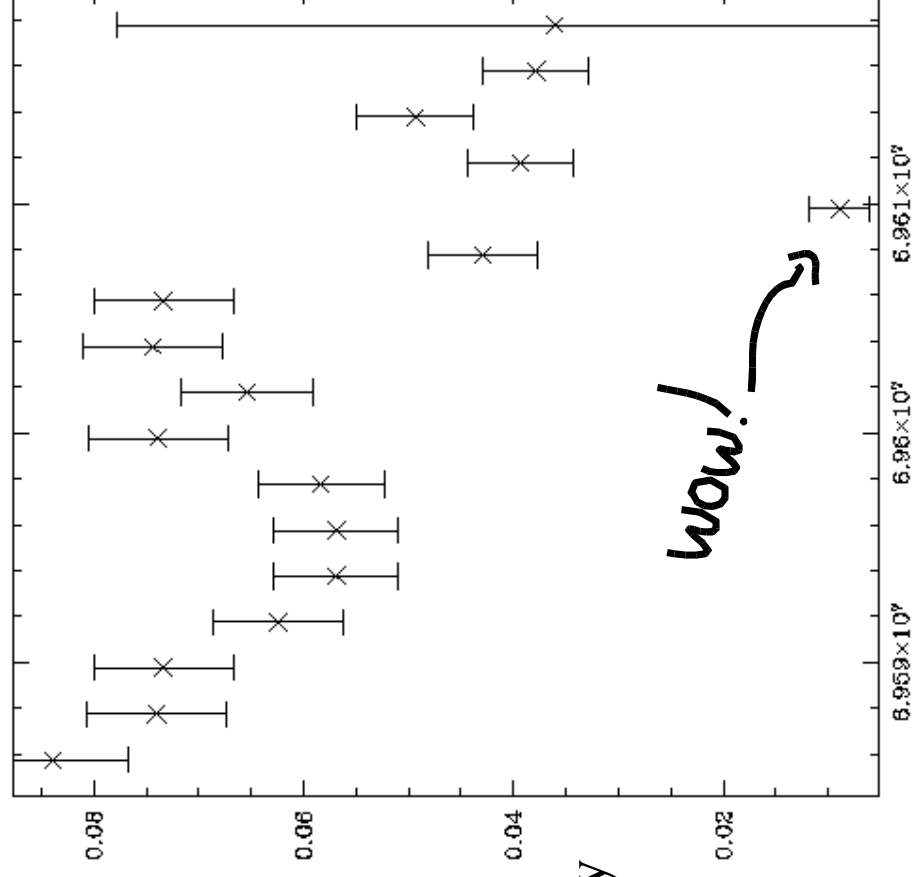




# Timing Analysis with Threads

## Outline

- I. Using Lightcurve
  - A. simple lightcurve
  - B. bugs/caveats
  - C. aliasing
  - D. intrinsic periods
- II. Phase-resolved Spectroscopy
  - A. create phase column
  - B. extract and fit data
  - C. caveats



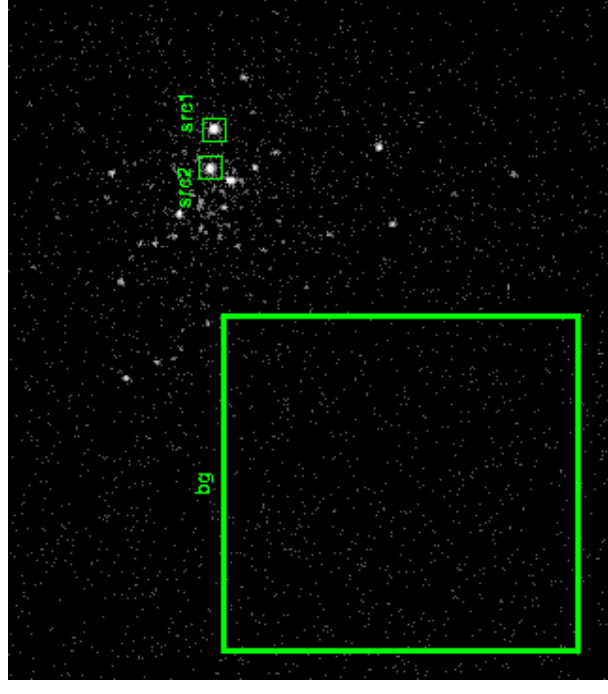


## I. Use and Caveats of Lightcurve

### References:

- [cxc.harvard.edu/ciao/threads/lightcurve](http://cxc.harvard.edu/ciao/threads/lightcurve)
- `ahelp lightcurve`
- [cxc.harvard.edu/bugs/lightcurve](http://cxc.harvard.edu/bugs/lightcurve)

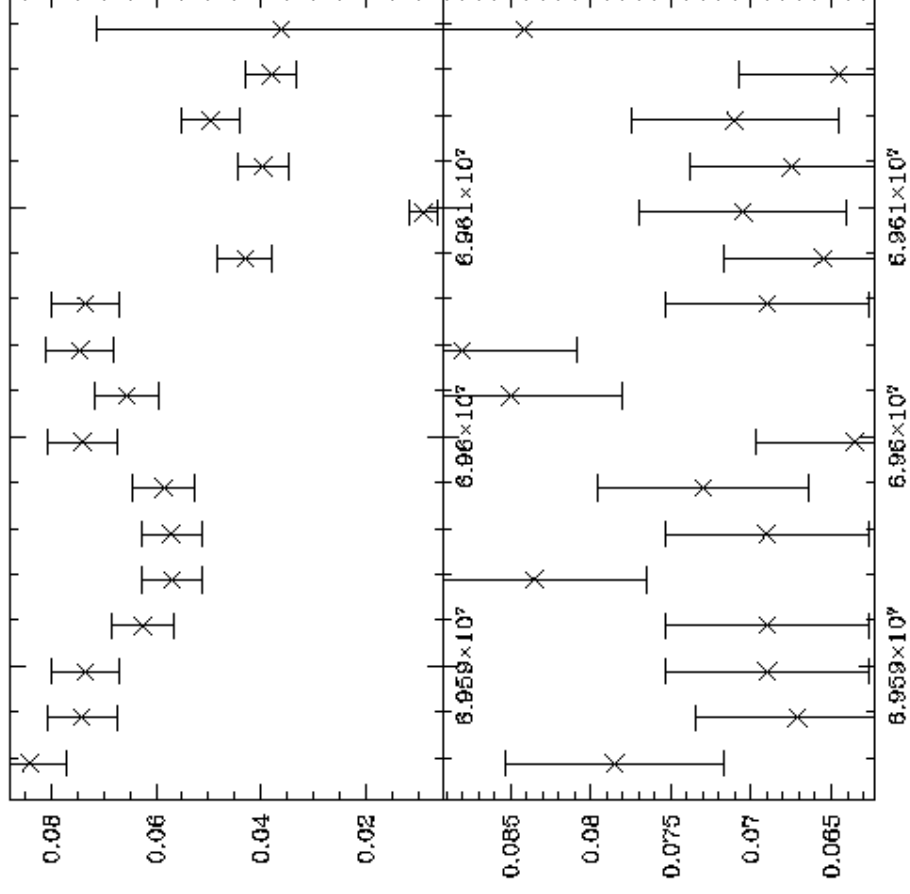
### A. Simple Lightcurve



```
unix% lightcurve infile="evt2.fits[ccd_id=3,sky=region(src1.reg)]"  
bkgfile="evt2.fits[ccd_id=3,sky=region(bg.reg)]"  
outfile="lc_bgsub.fits" binlength="2000"
```



... or we can extract two lightcurves for comparison. Simple, no?



## B. Bugs/Caveats

1. GTIs, [ccd\_id=n] filter
2. area bug, background subtraction

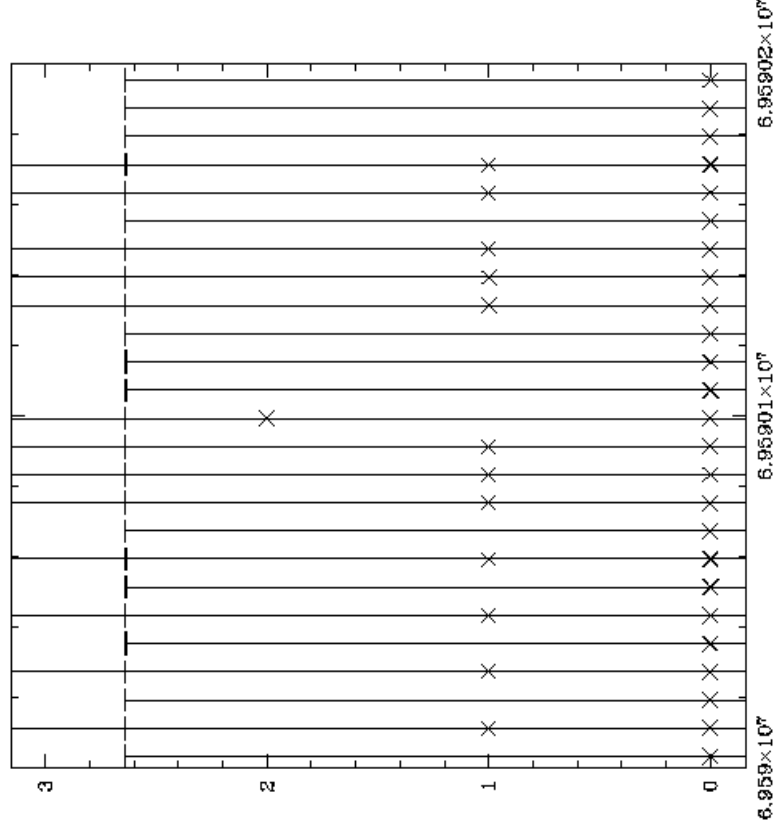


## C. Aliasing: the big caveat

Aliasing is the apparent periodicity formed by the combination of two other periods.

Example 1: Underbinned lightcurve  $\rightarrow$

Example 2: Moiré pattern





A large rectangular area filled with horizontal ruling lines, typical of a composition or writing page. The lines are evenly spaced and cover the entire central portion of the page.



● More properly, lightcurve bins should be **integer multiples** of any natural or intrinsic periodicity. Unfortunately, this can be tricky...

<u>Intrinsic Periods</u>	<u>Dither Periods</u>	
3.2 s, ACIS TE-mode	HRC	ACIS
2.8 ms, ACIS CC-mode	Pitch	768.6 s
~40 $\mu$ s, ACIS readout streak	Yaw	1087.0 s
~0.5 s, HRC (new!)		1000.0 s

● A good way to verify a suspected period is to rebin the lightcurve with a different resolution. Any aliasing with the lightcurve binsize will change with the binsize.



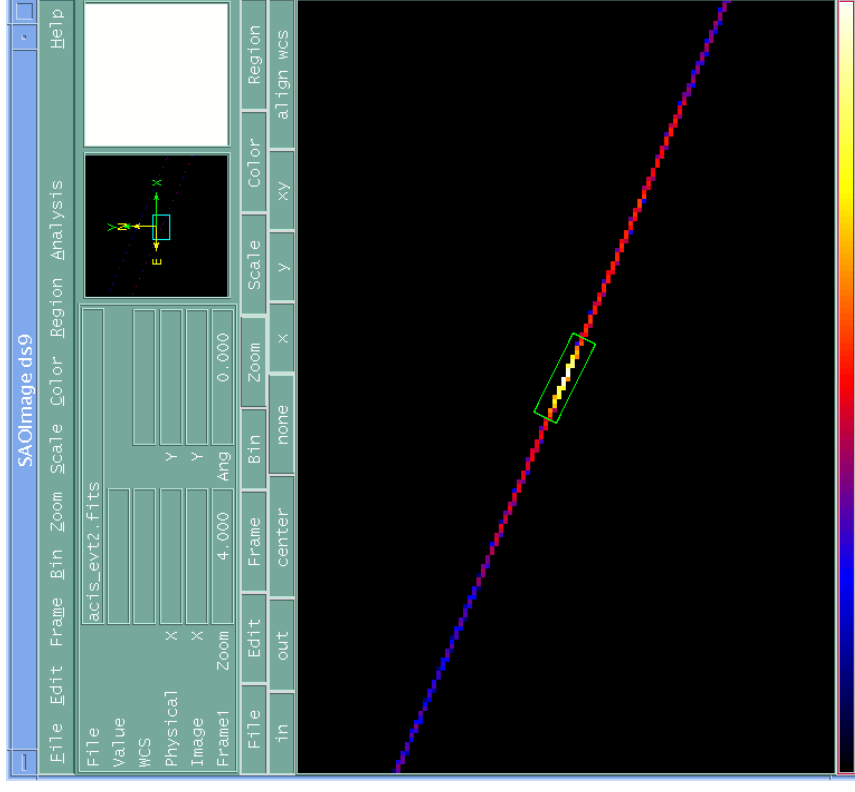
## II. Phase-resolved Spectroscopy

### References:

- [cxc.harvard.edu/ciao/threads/phase\\_bin](http://cxc.harvard.edu/ciao/threads/phase_bin)
- ahelp dmtcalc
- see also "faseBin" contributed software

### Steps:

1. Apply barycenter time corrections (if necessary, see thread)
2. Use an assumed period (and  $\dot{P}$ ) to convert TIME to PHASE
3. Extract and plot spectrum



PSR B0540-69  
(ACIS CC-mode)

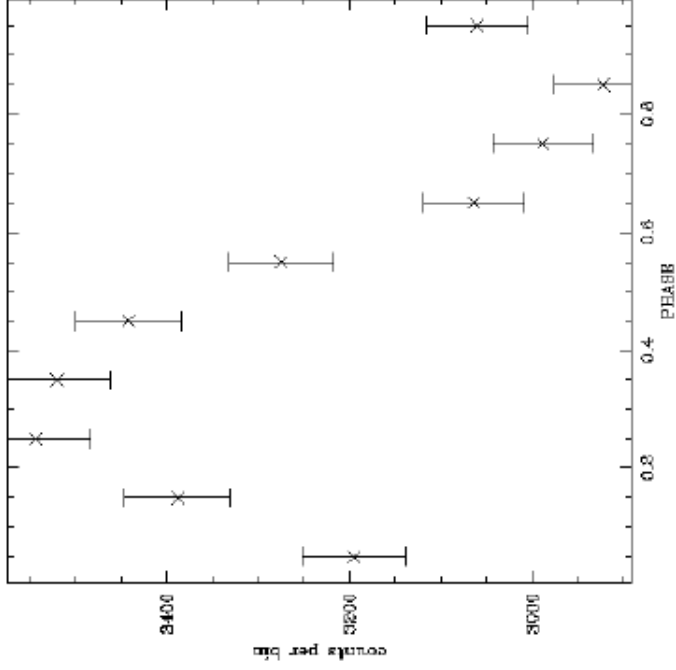


use dmtcalc to calculate the PHASE for each event:

```
dmtcalc infile="evt2.fits" outfile="evt2_phase.fits" expression="@dmtcalc.txt"
```

dmtcalc.txt has the expression. first "phase", then "fold":

```
GPHASE=(time-TSTART)*(-4.79e-13*(time-TSTART))  
PHASE=GPHASE-(int)GPHASE
```



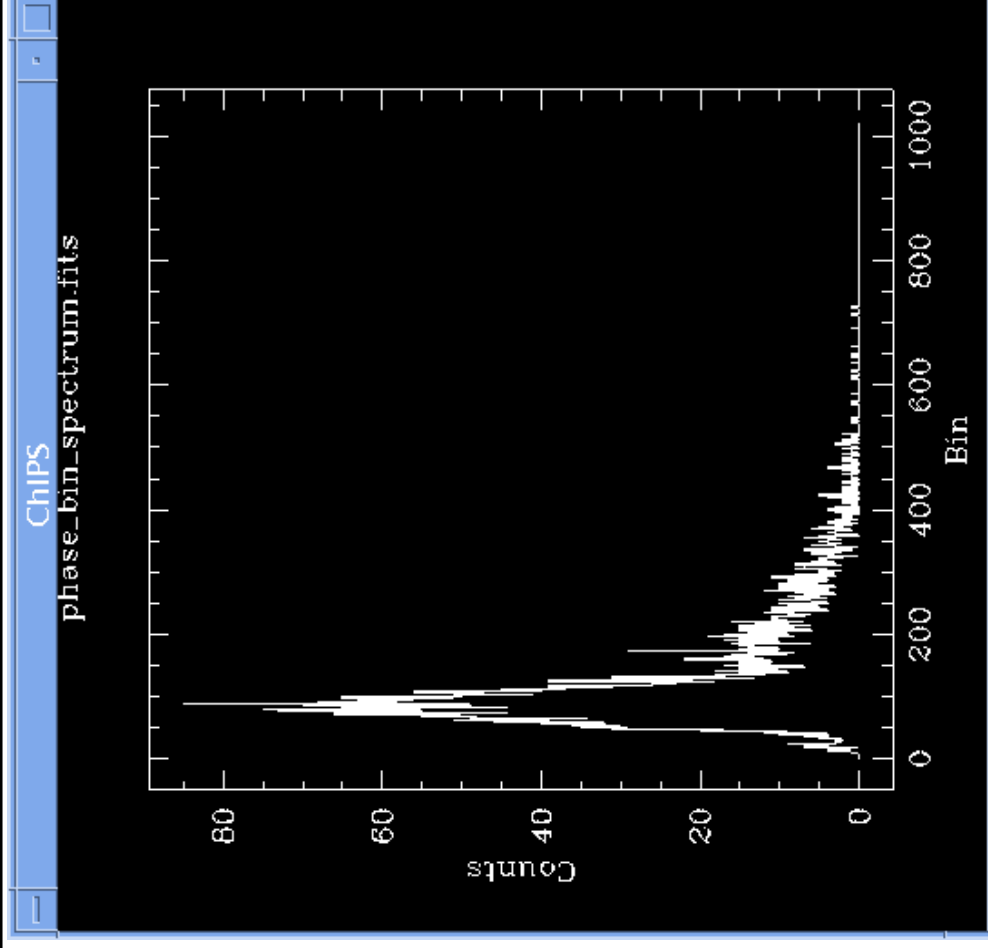
to check shape of oscillation,  
use dmextract to "[bin phase=0:1:0.1]"





Now we extract a spectrum and plot...

```
dmextract "evt2_phase.fits[sky=region(source.reg),phase=0.2:0.4][bin pi]"
```



Caveat: no GTIs used