

# The splashback radius of dark matter halos

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*In collaboration with:*

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Einstein Fellows Symposium • Harvard University • 10/02/2018

# Cosmology?

forward modeling

photometric redshifts

surveys

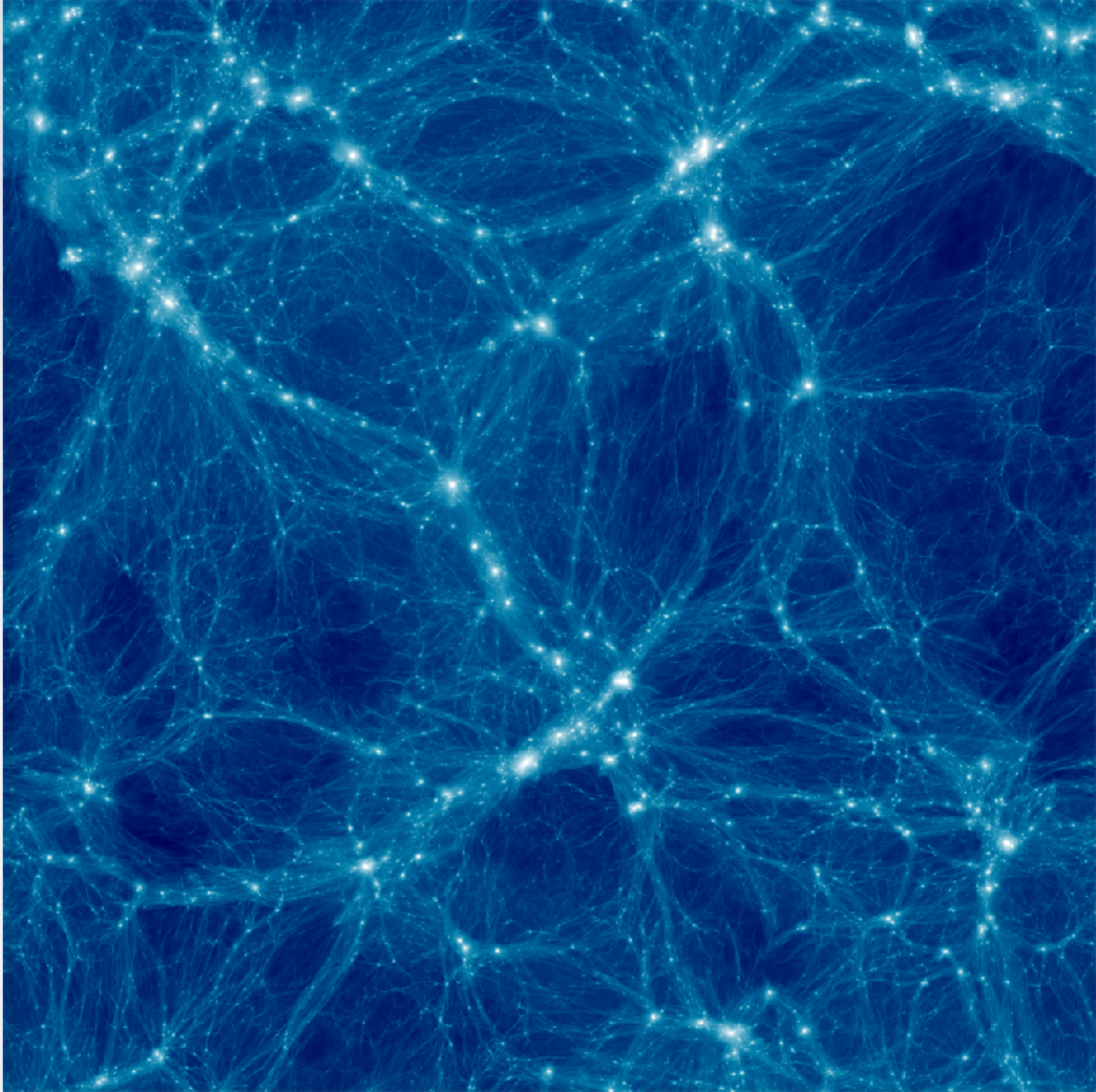
fuzzy dark matter

perturbation theory

**...sort of: structure formation!**



89 Mpc



**Visualization code:**

Phil Mansfield

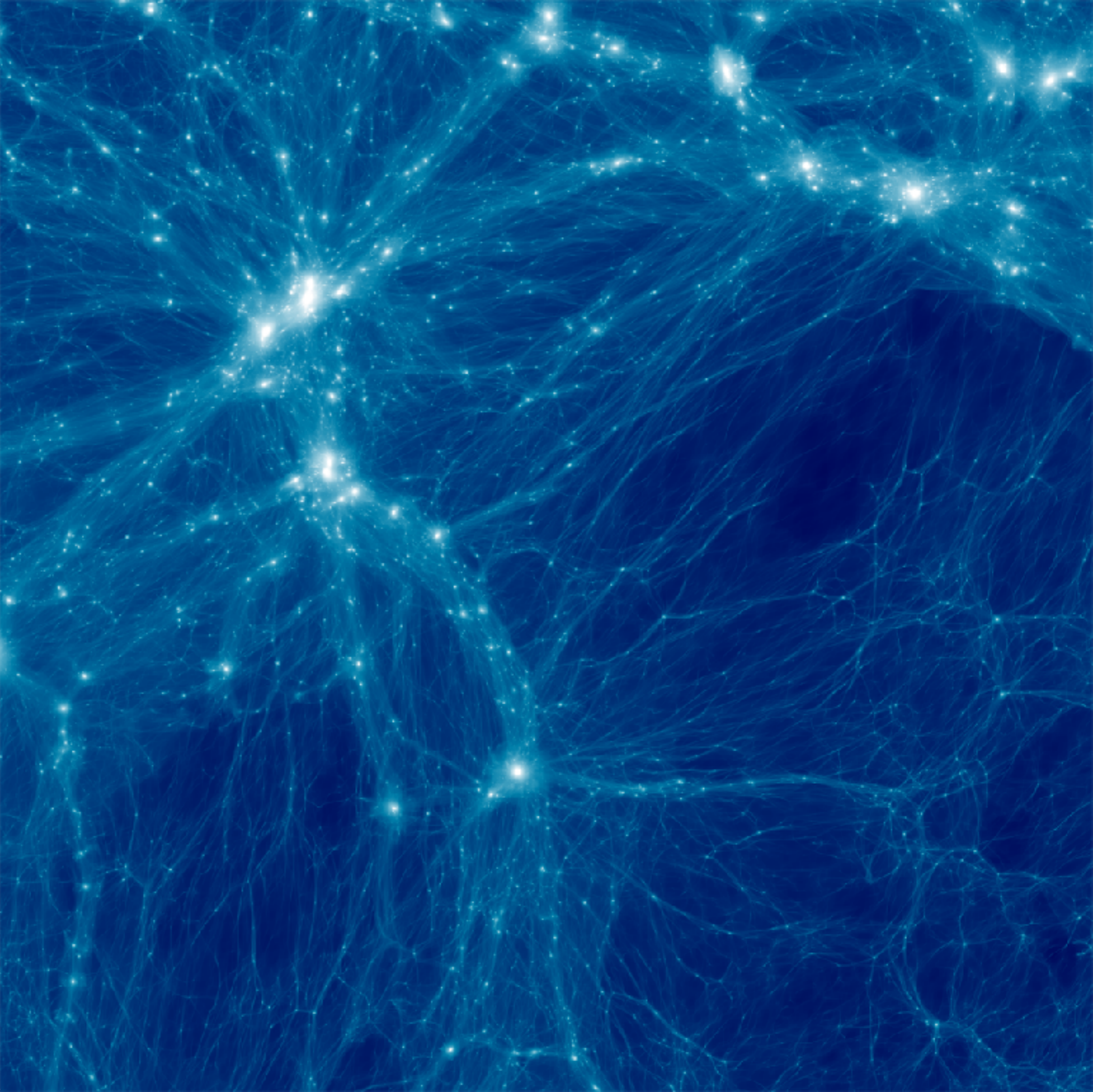
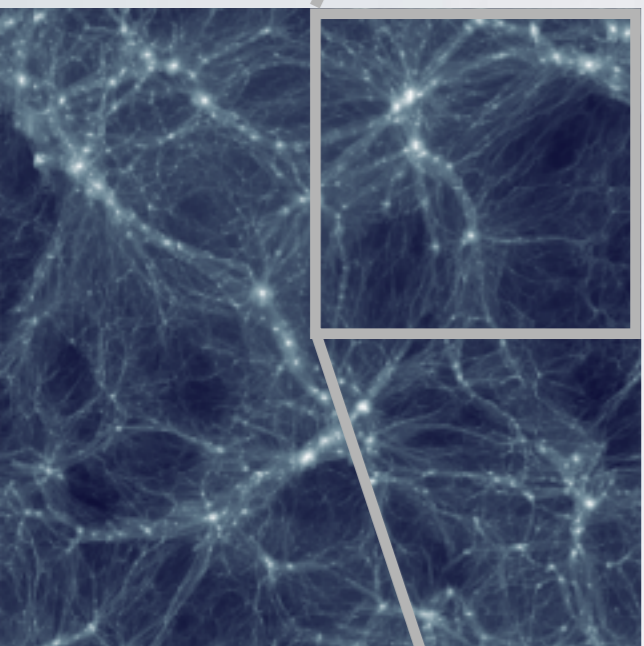
**Algorithm:**

Tom Abel, Oliver Hahn,

Ralf Kaehler

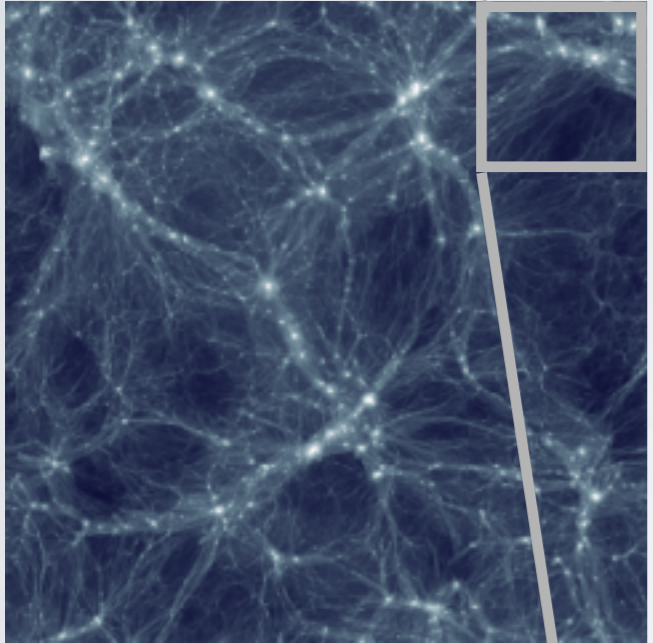
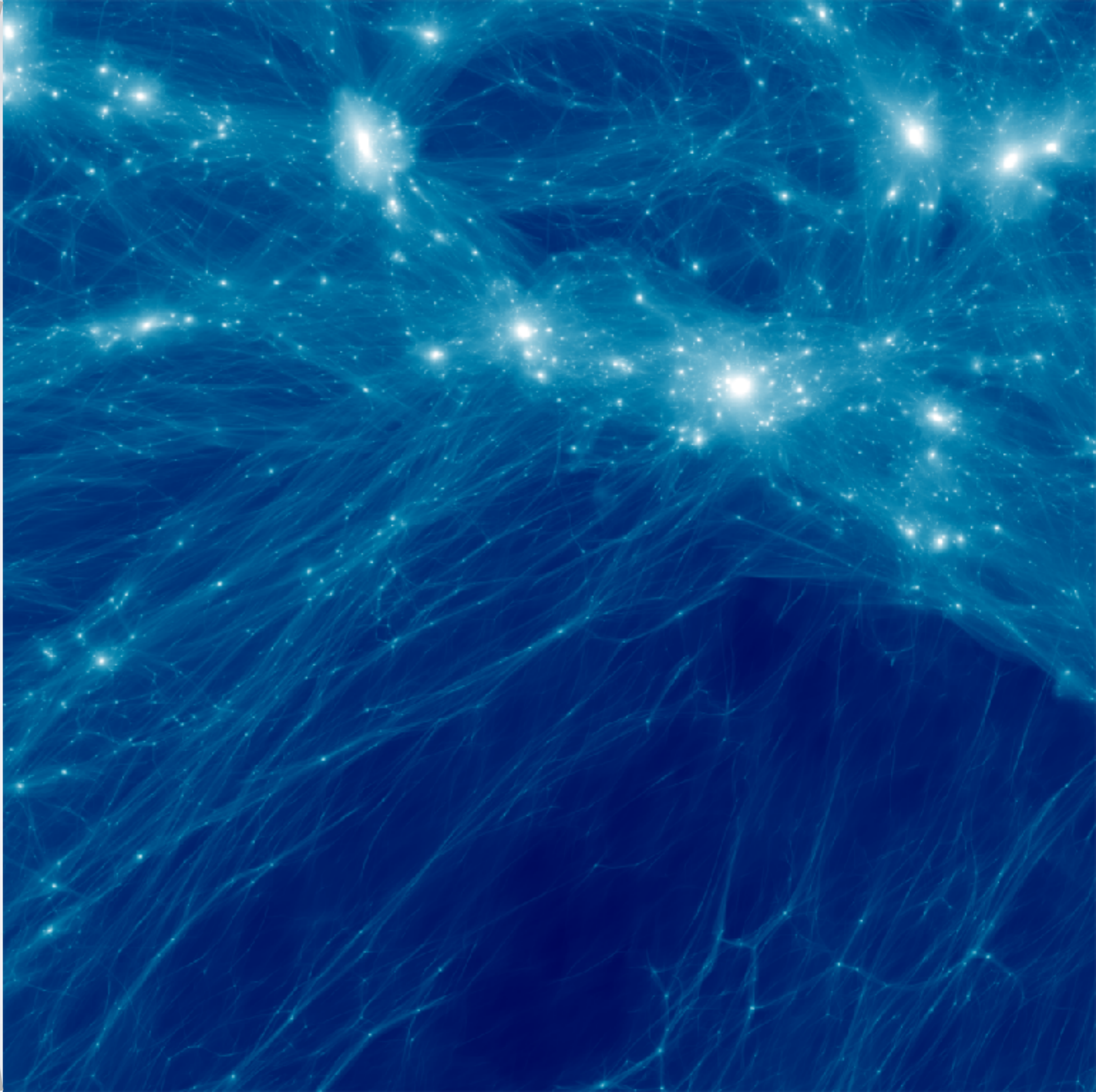


45 Mpc



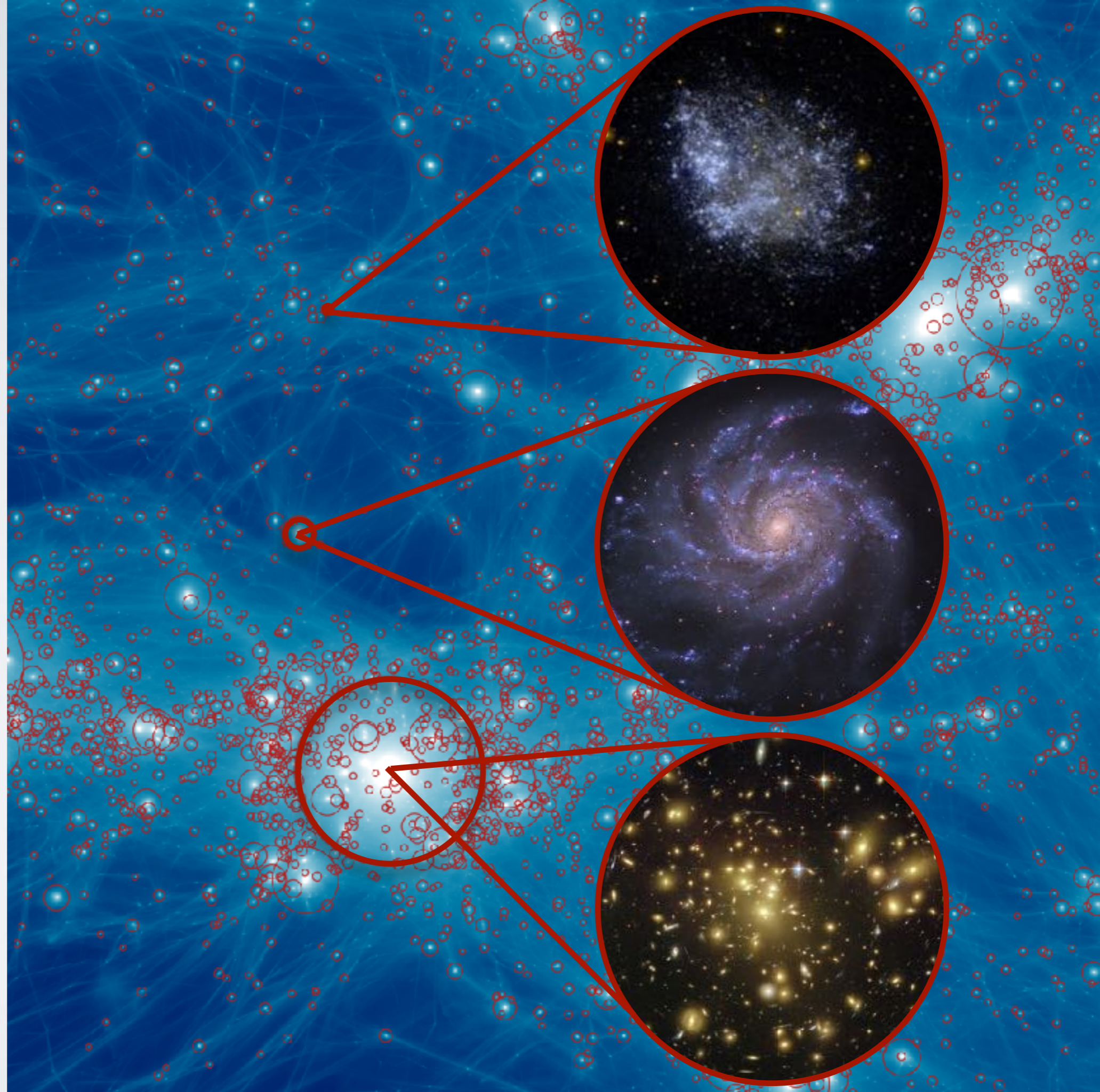


22 Mpc

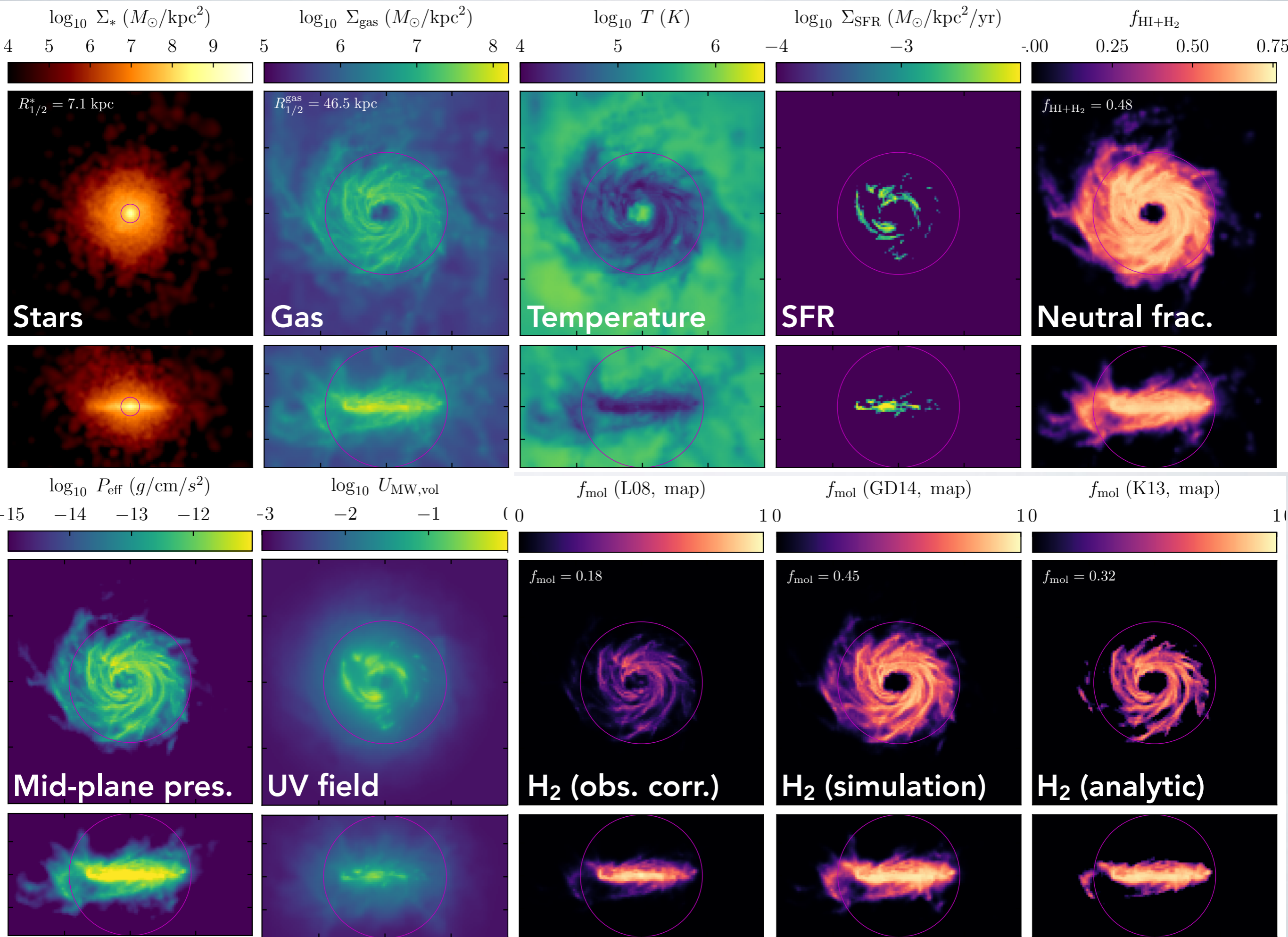




11 Mpc

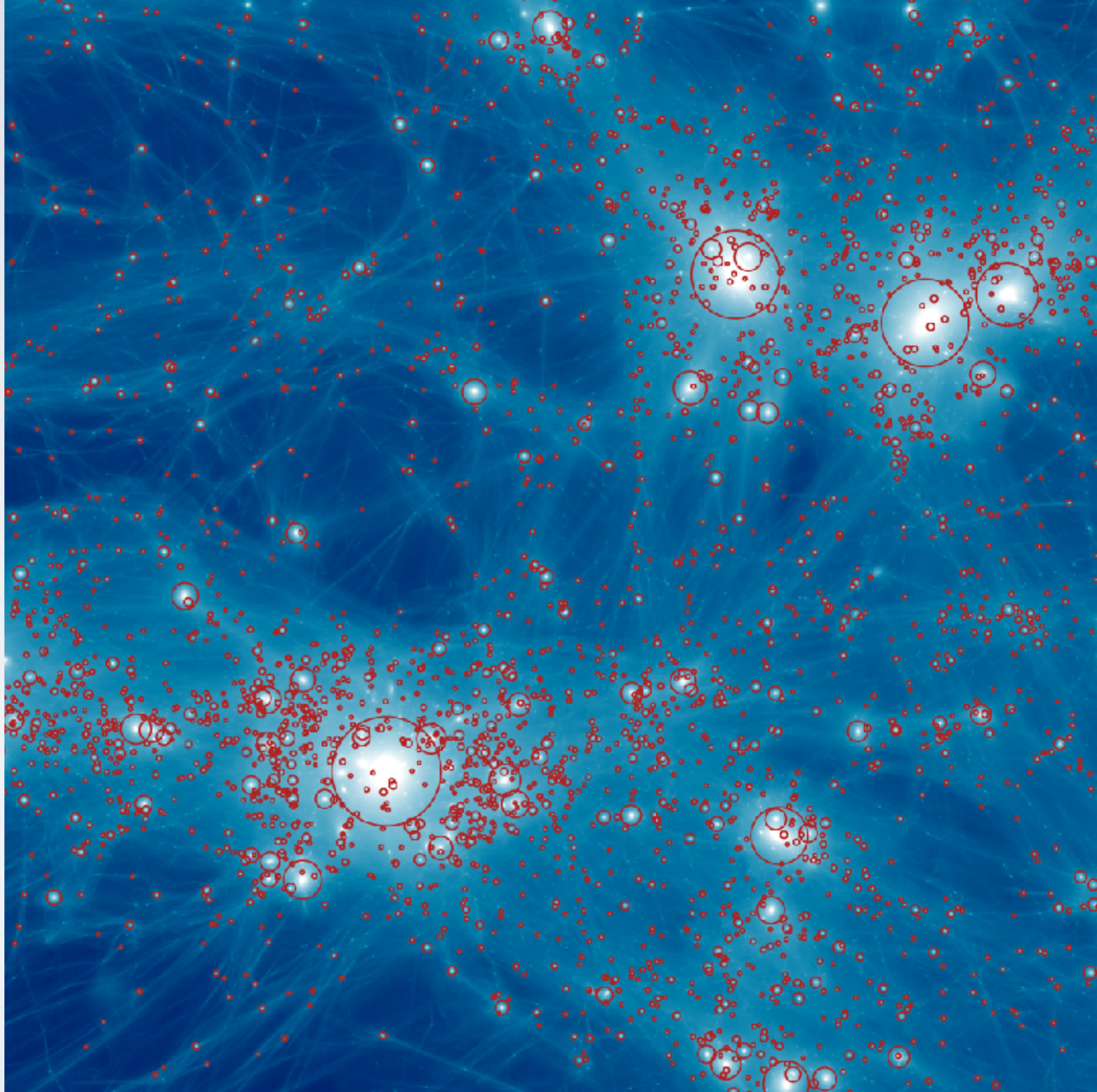








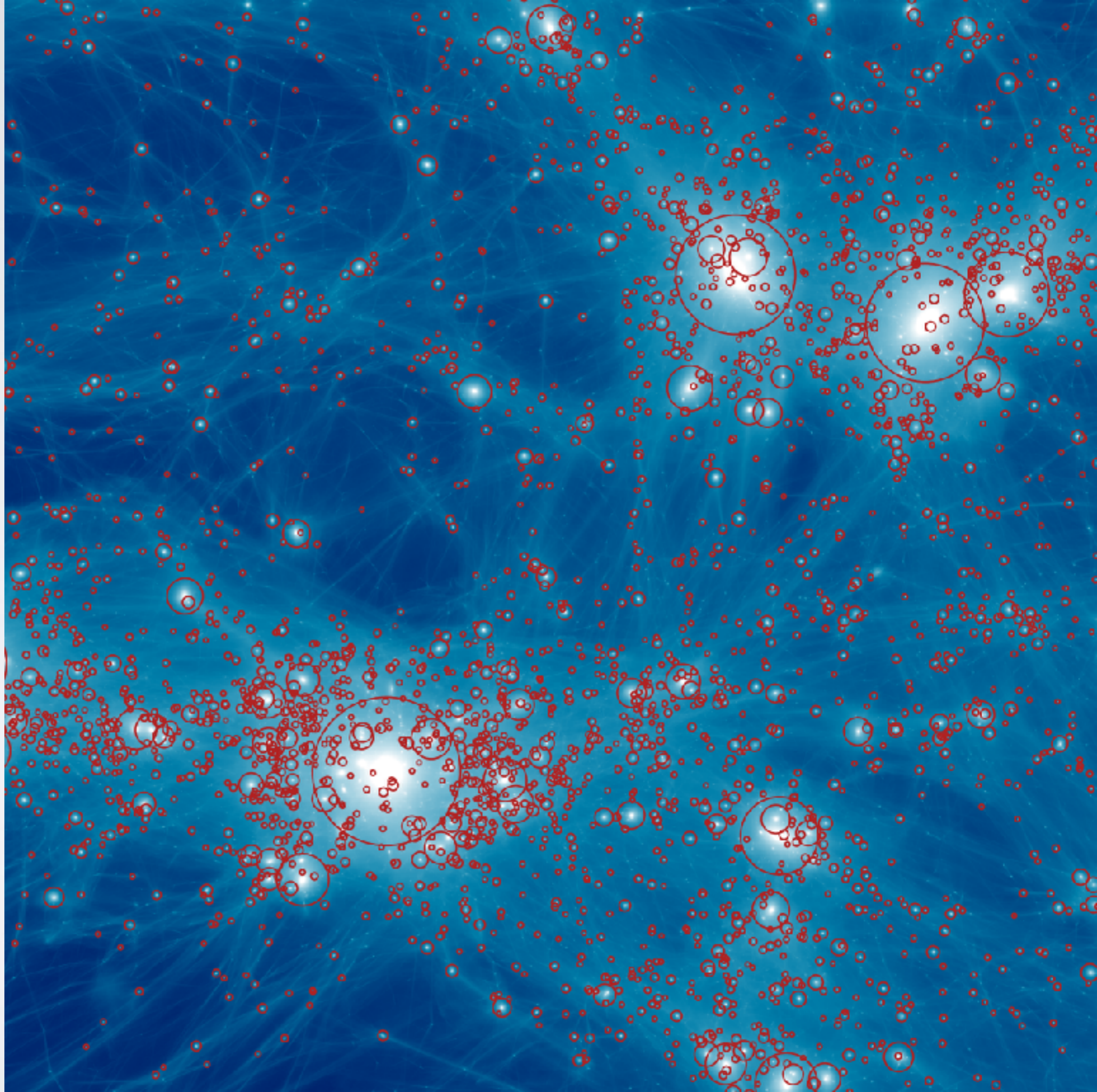
$R_{500c}$



Halo finder: Rockstar  
(Behroozi et al. 2013)



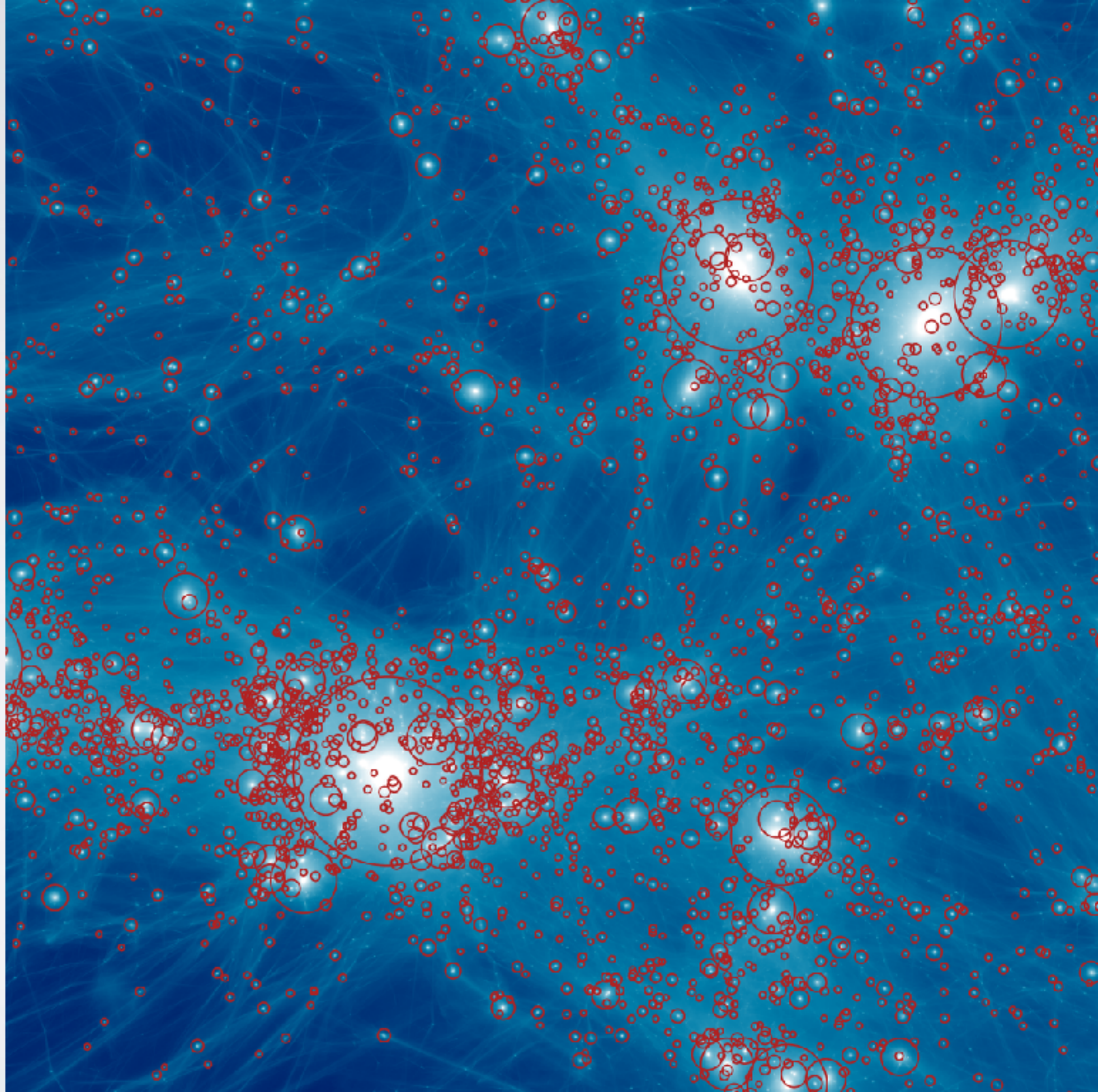
$R_{200c}$



Halo finder: Rockstar  
(Behroozi et al. 2013)



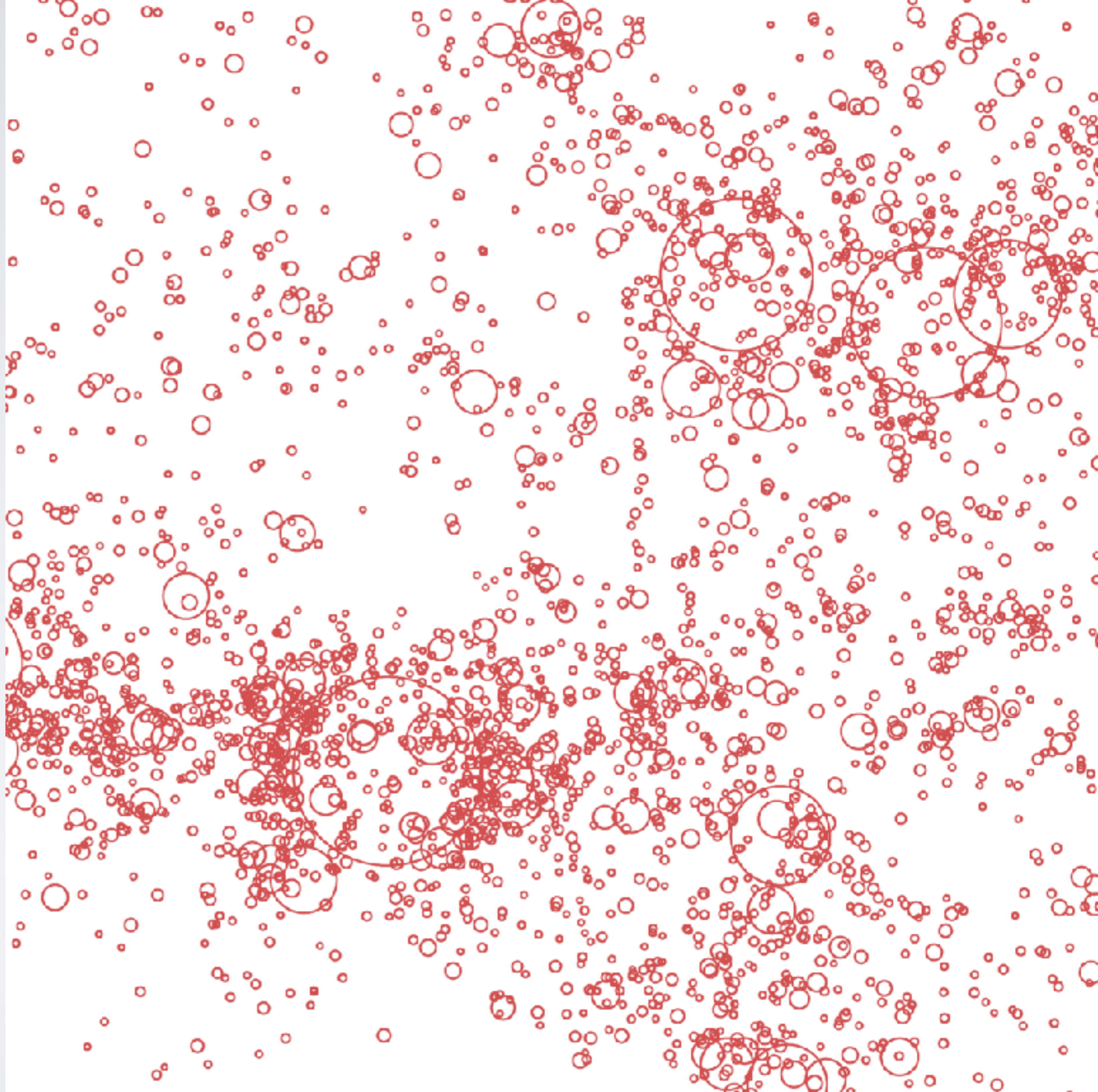
$R_{\text{vir}}$



Halo finder: Rockstar  
(Behroozi et al. 2013)



$R_{\text{vir}}$



Halo finder: Rockstar  
(Behroozi et al. 2013)



# Density profile

$\log \rho$

Scale radius:  
 $d \log(\rho) / d \log(r) = -2$

Outer radius (enclosing  
some mean overdensity)

Mass:  
 $M_{\Delta} = 4\pi/3 \Delta \rho_{\text{ref}} R_{\Delta}^3$

$r_s$

$R_{200c}$   $R_{\text{vir}}$   $R_{200m}$

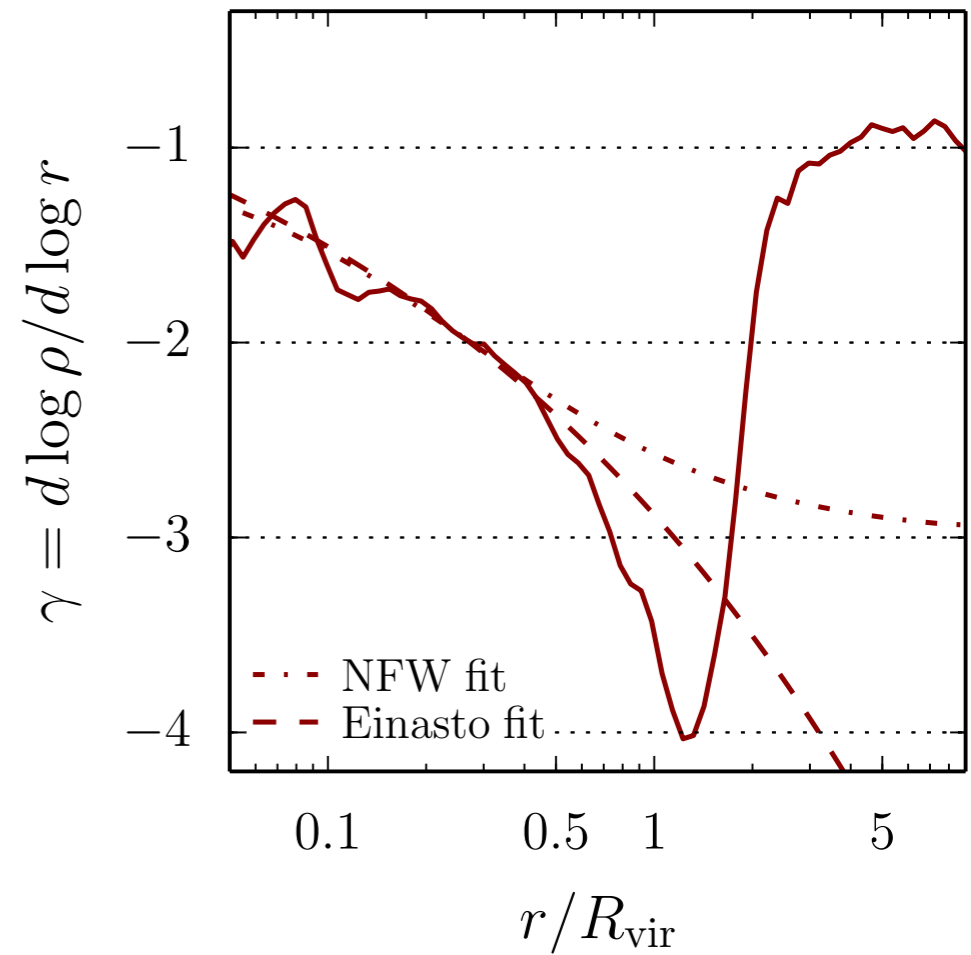
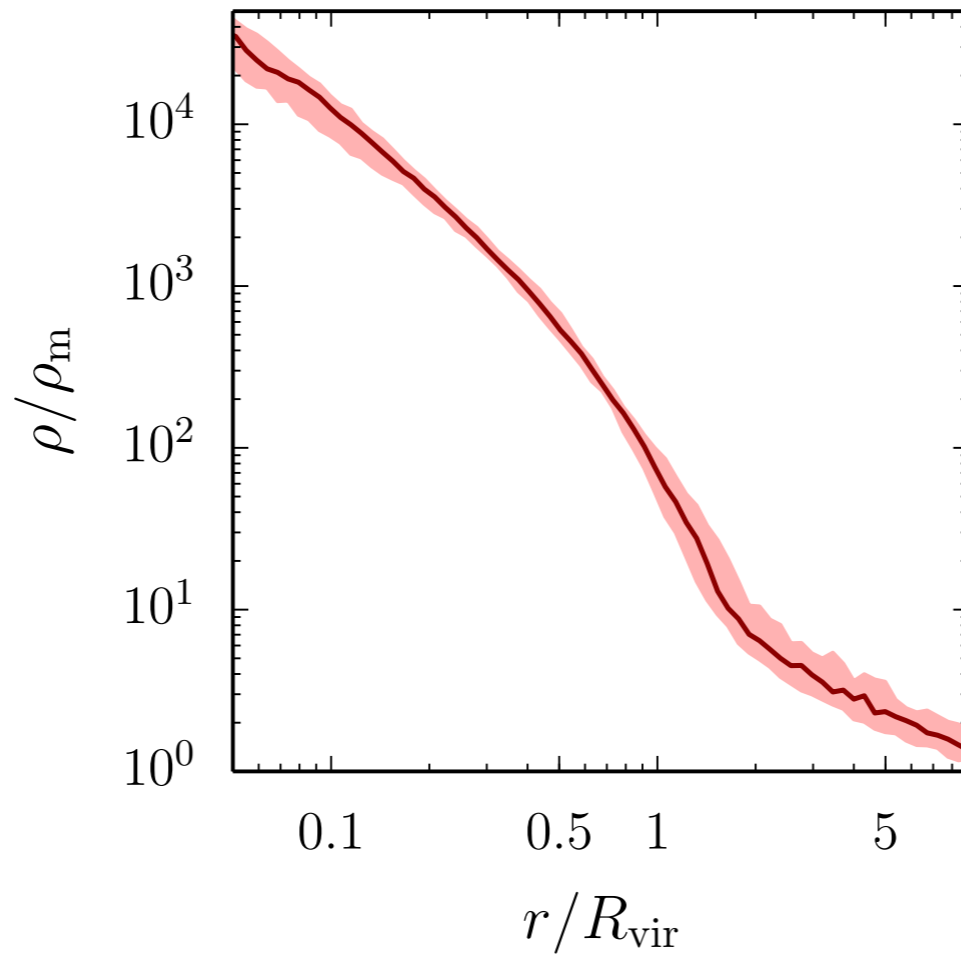
$\log r$

Einasto 1965 • Frenk et al. 1988 • Hernquist 1990 • Dubinski & Carlberg 1991

Navarro et al. 1995/1996/1997/2004



Large halos  
( $M > 10^{15}$ )



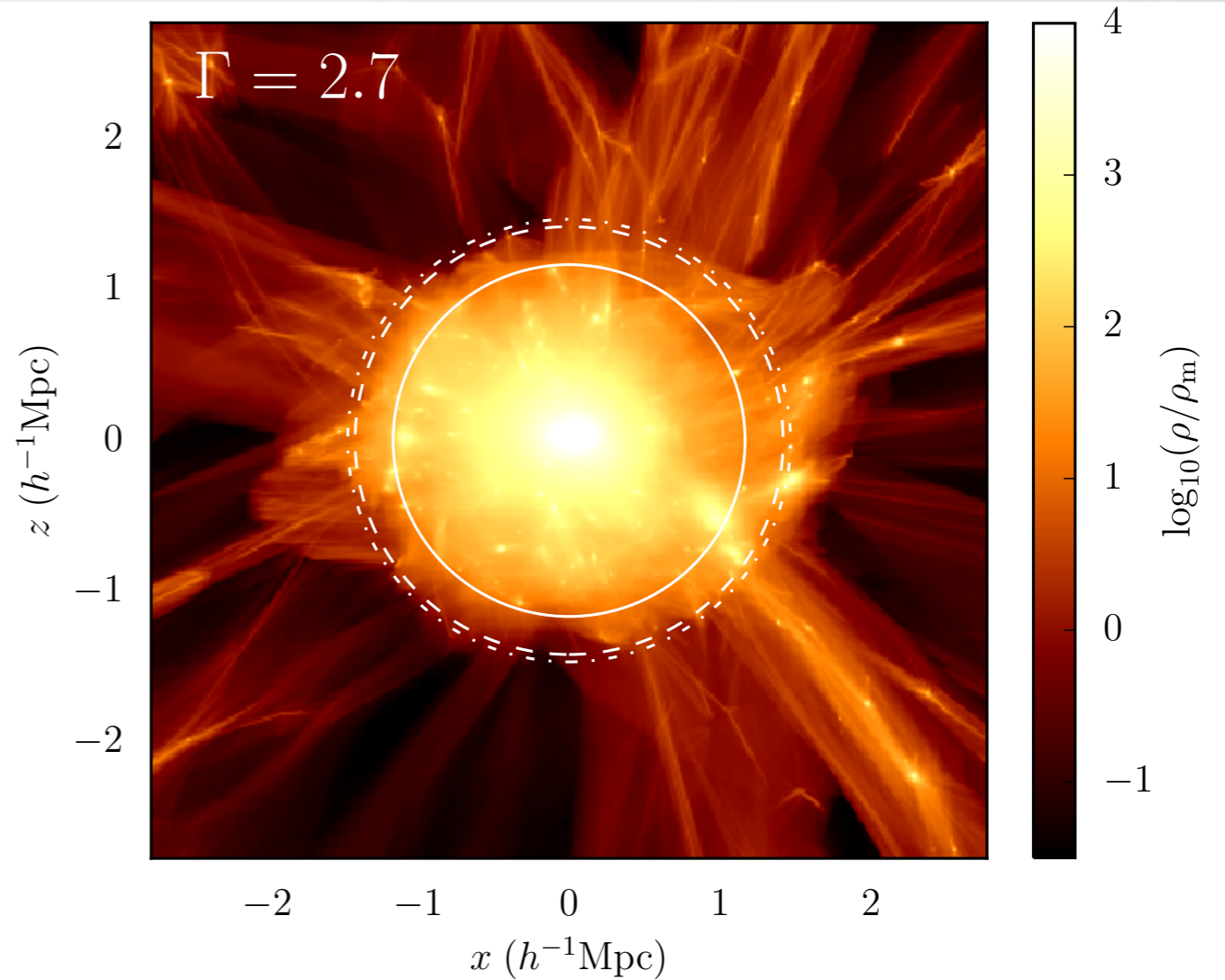
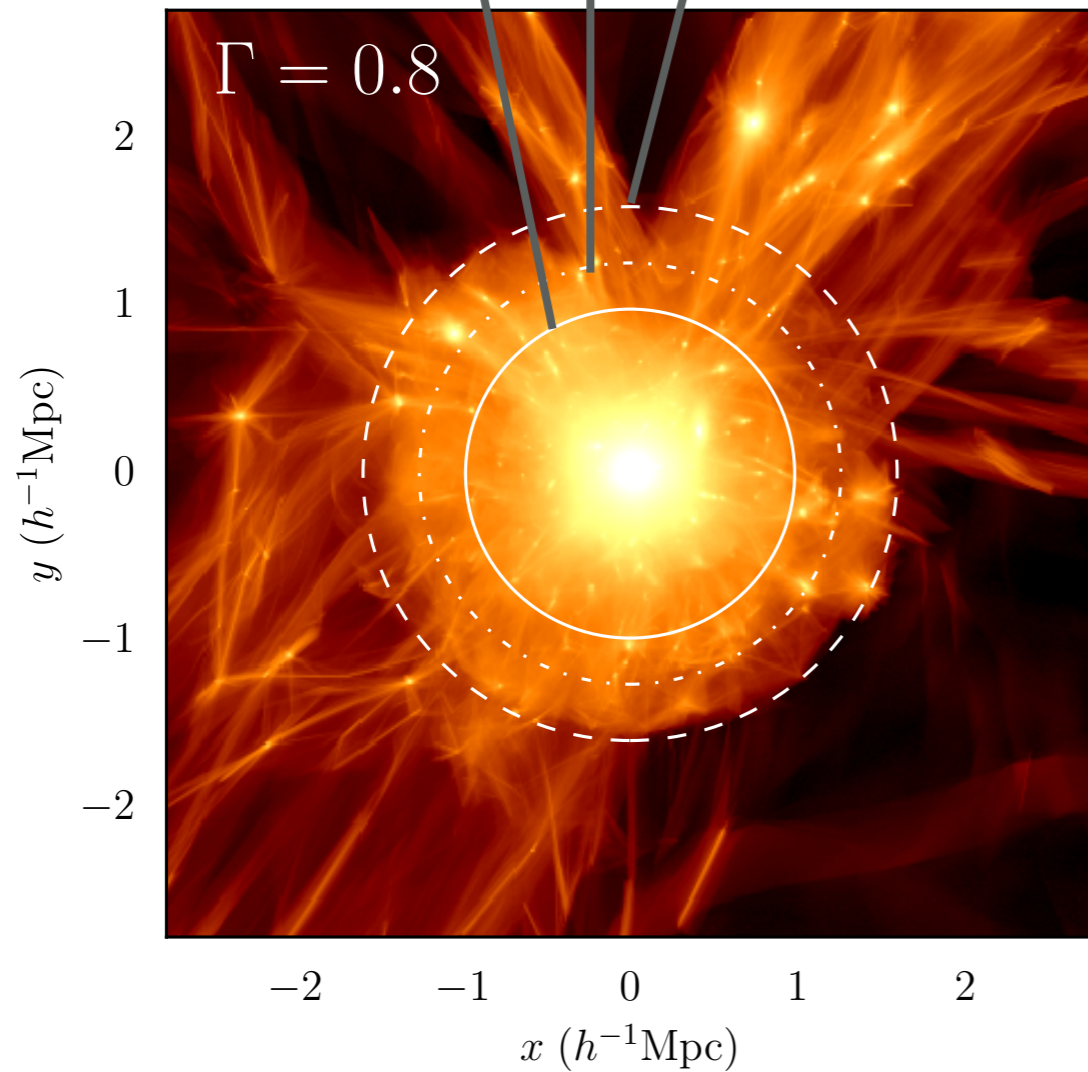


# The Splashback Radius

$R_{200m}$

$R_{vir}$

$R_{splashback}$

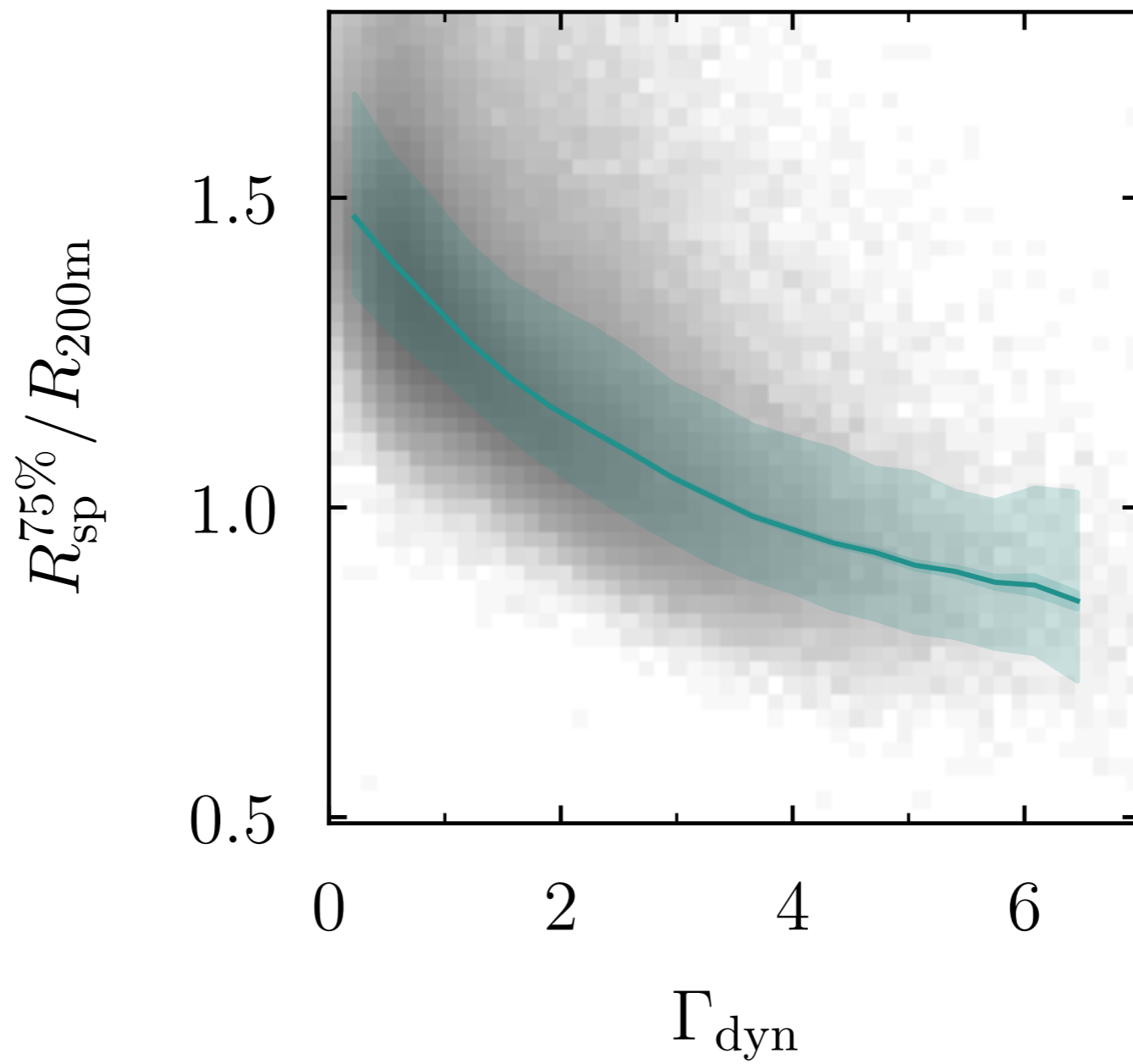


Low accretion rate

High accretion rate

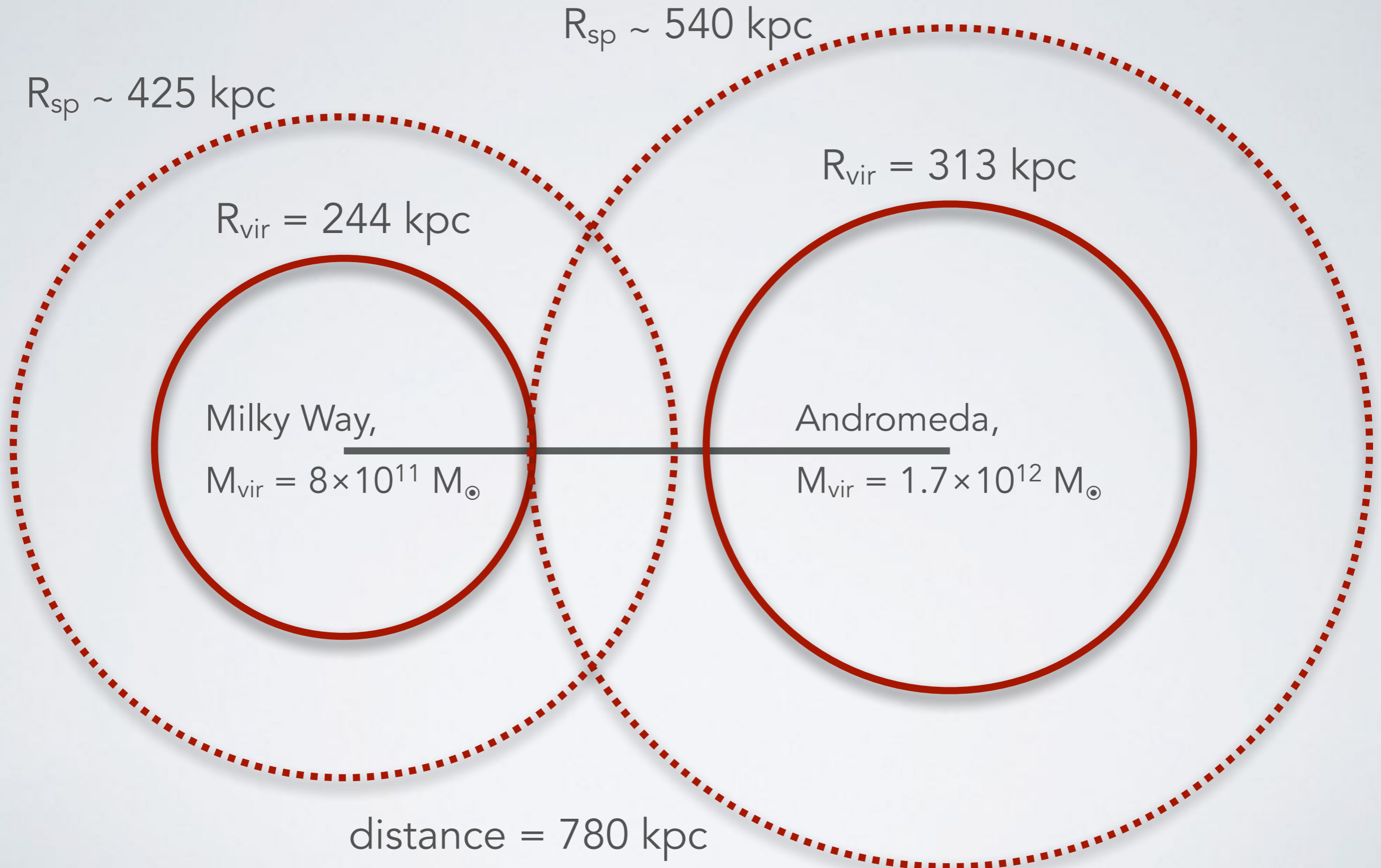


# The $\Gamma$ - $R_{sp}$ relation



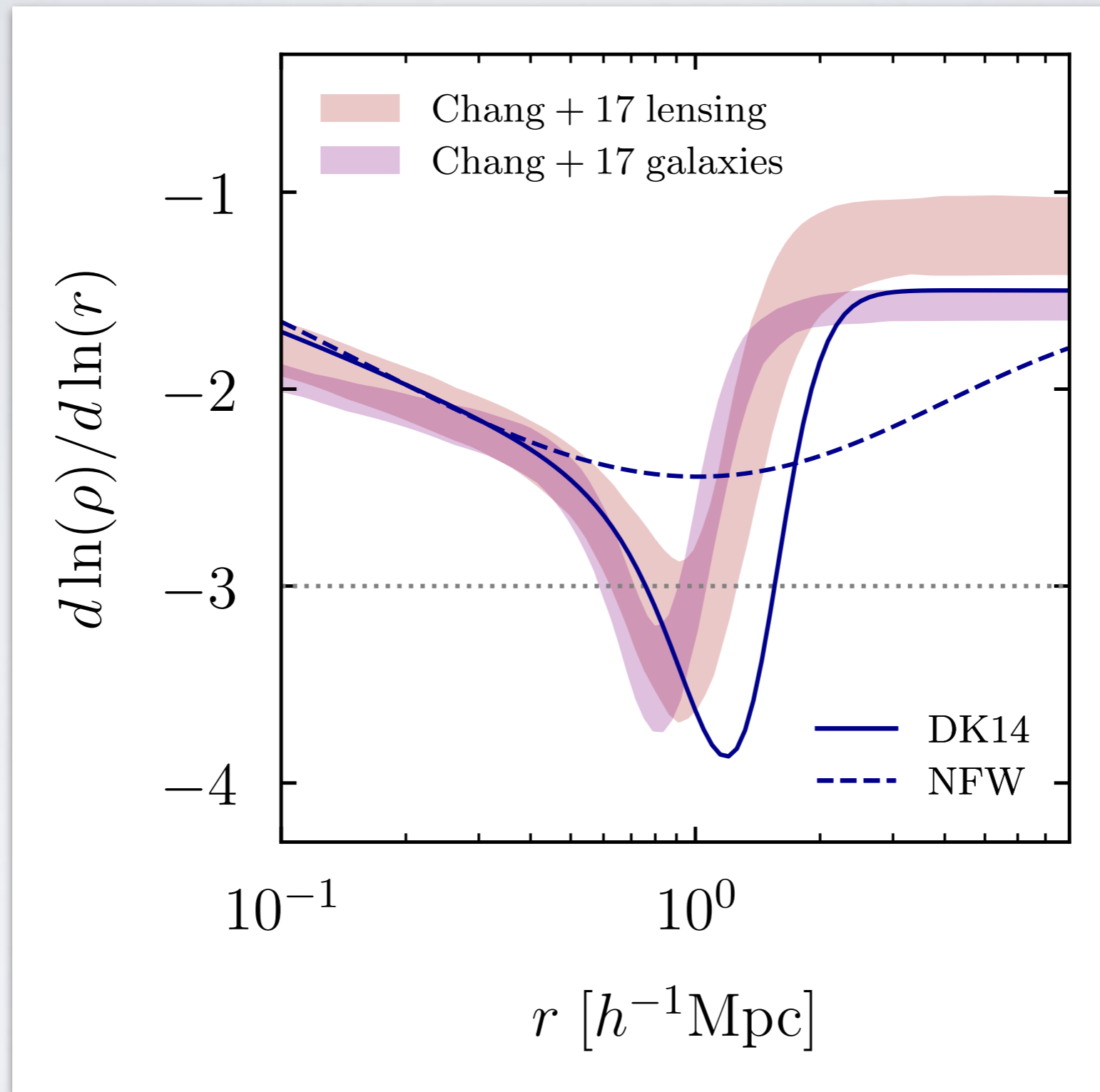


# Do the Milky Way and Andromeda halos overlap?





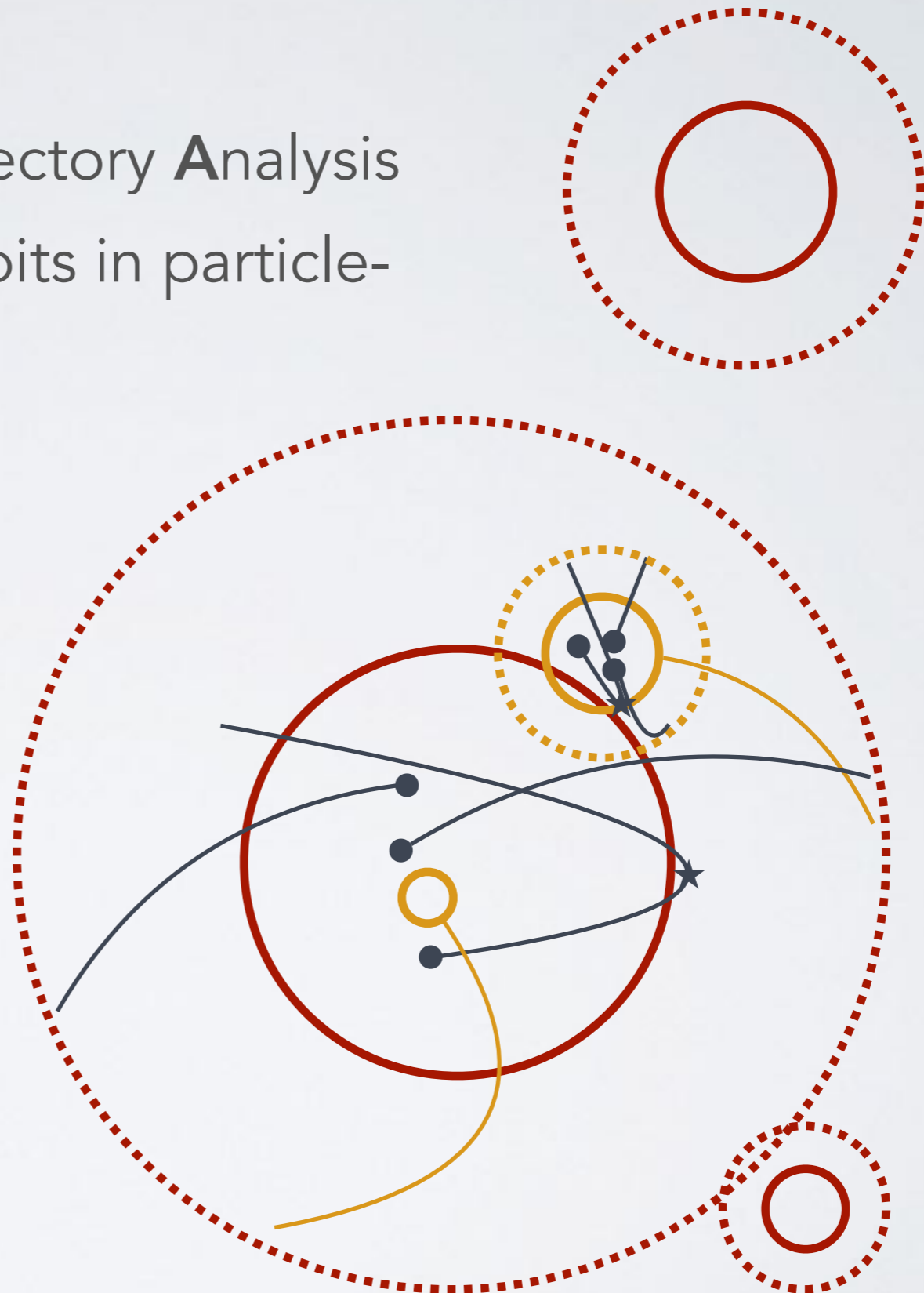
# $R_{sp}$ in cluster member profiles (DES)





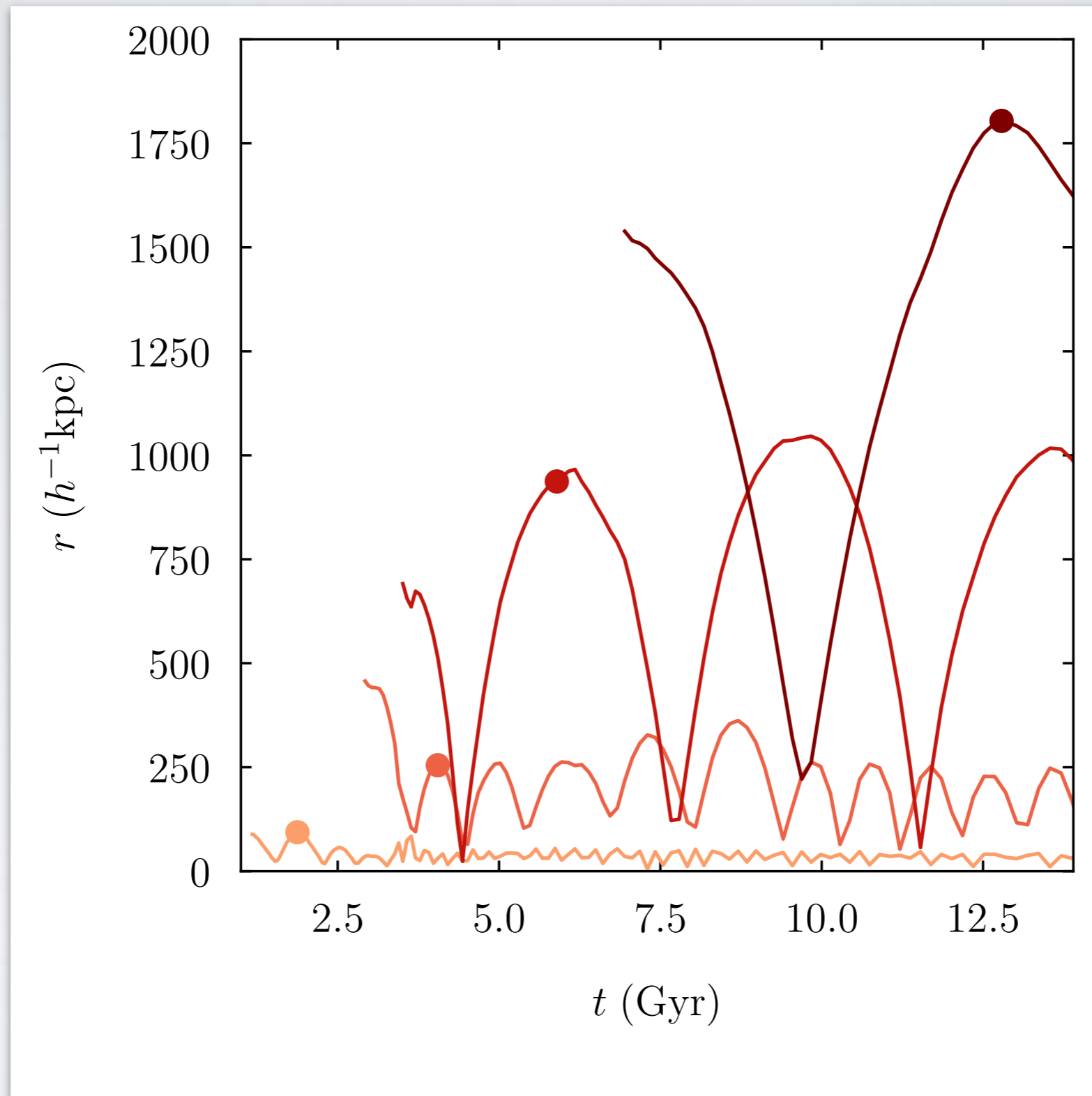
# SPARTA

- Subhalo and **PAR**ticle **T**rajectory **A**nalysis
- Framework for tracking orbits in particle-based simulations
- MPI-parallelized, pure C

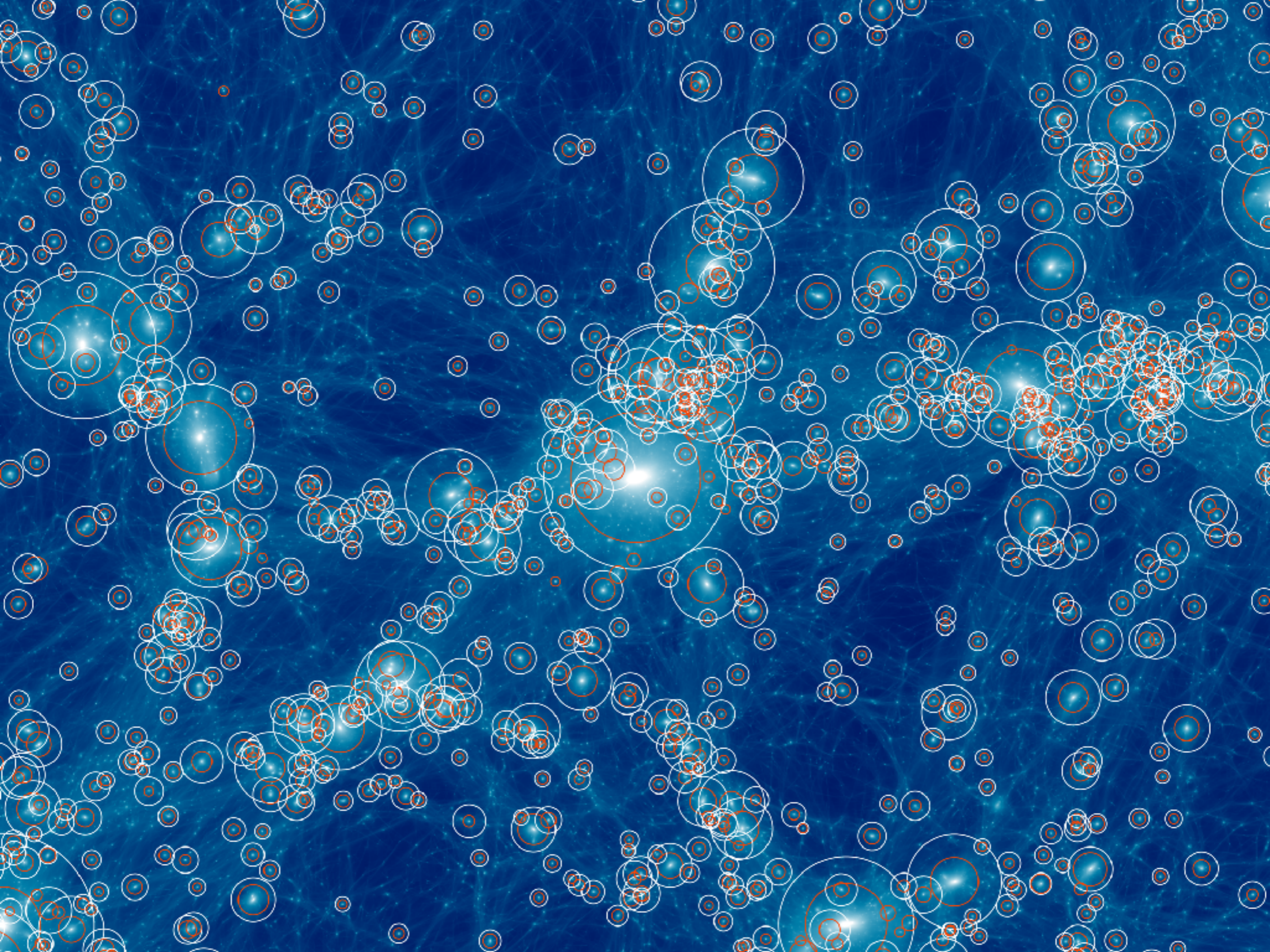




# What do the orbits look like?

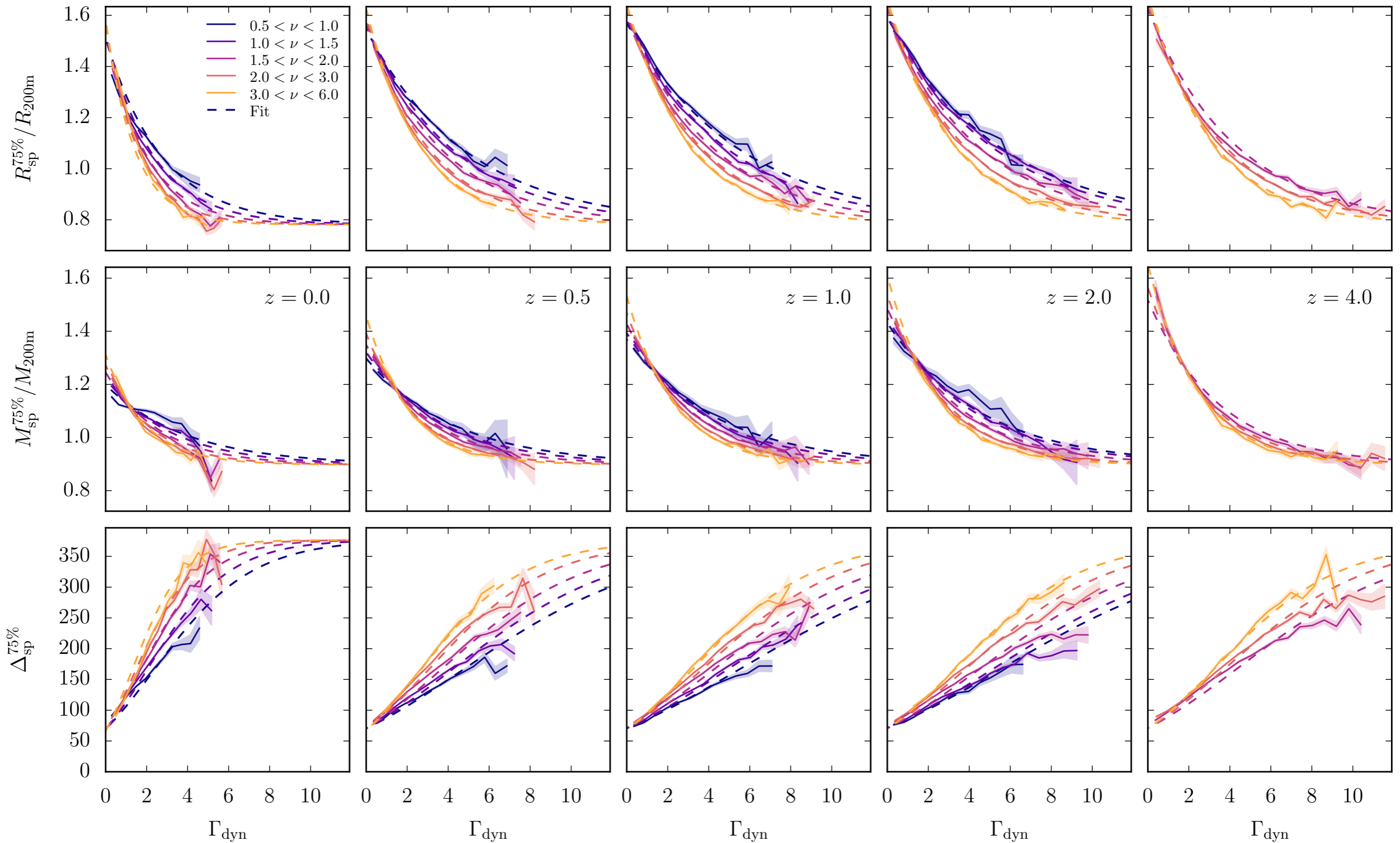








# The $\Gamma$ - $R_{\text{sp}}$ relation

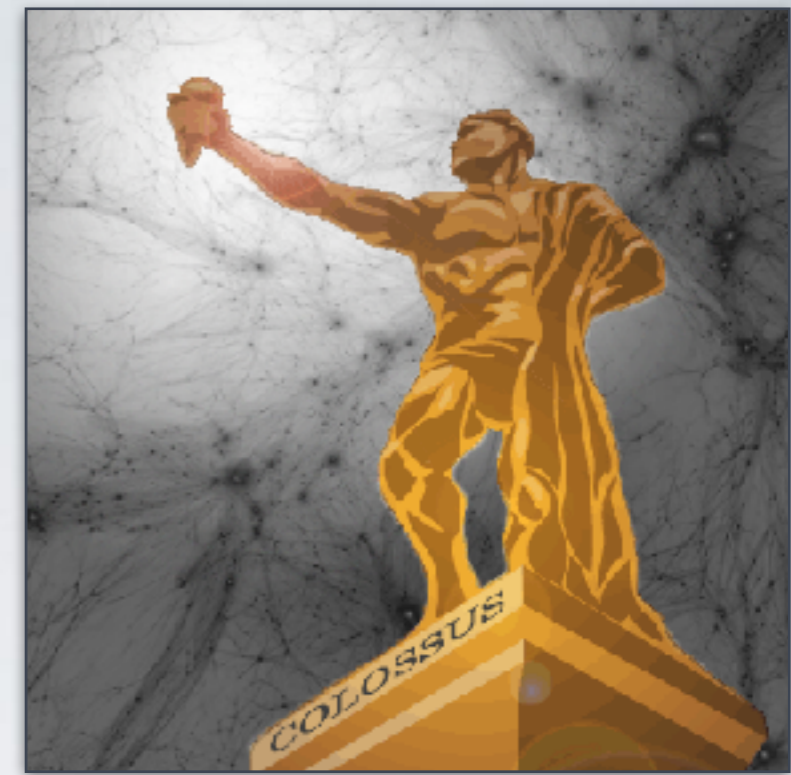




# COLOSSUS

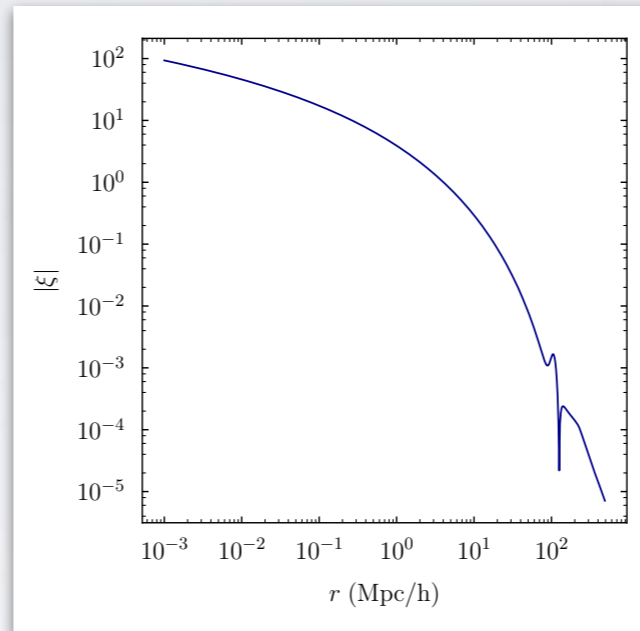
Cosmology, halos, and large-scale structure

[benediktdiemer.com/code/colossus](http://benediktdiemer.com/code/colossus)



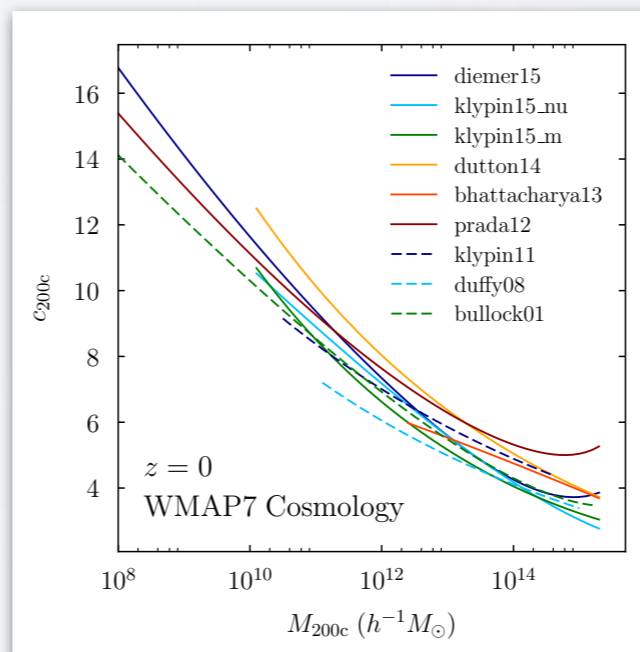
```
from colossus.cosmology import cosmology
```

```
cosmo = cosmology.setCosmology('WMAP9')  
xi = cosmo.correlationFunction(10.0)
```



```
from colossus.halo import concentration as hc
```

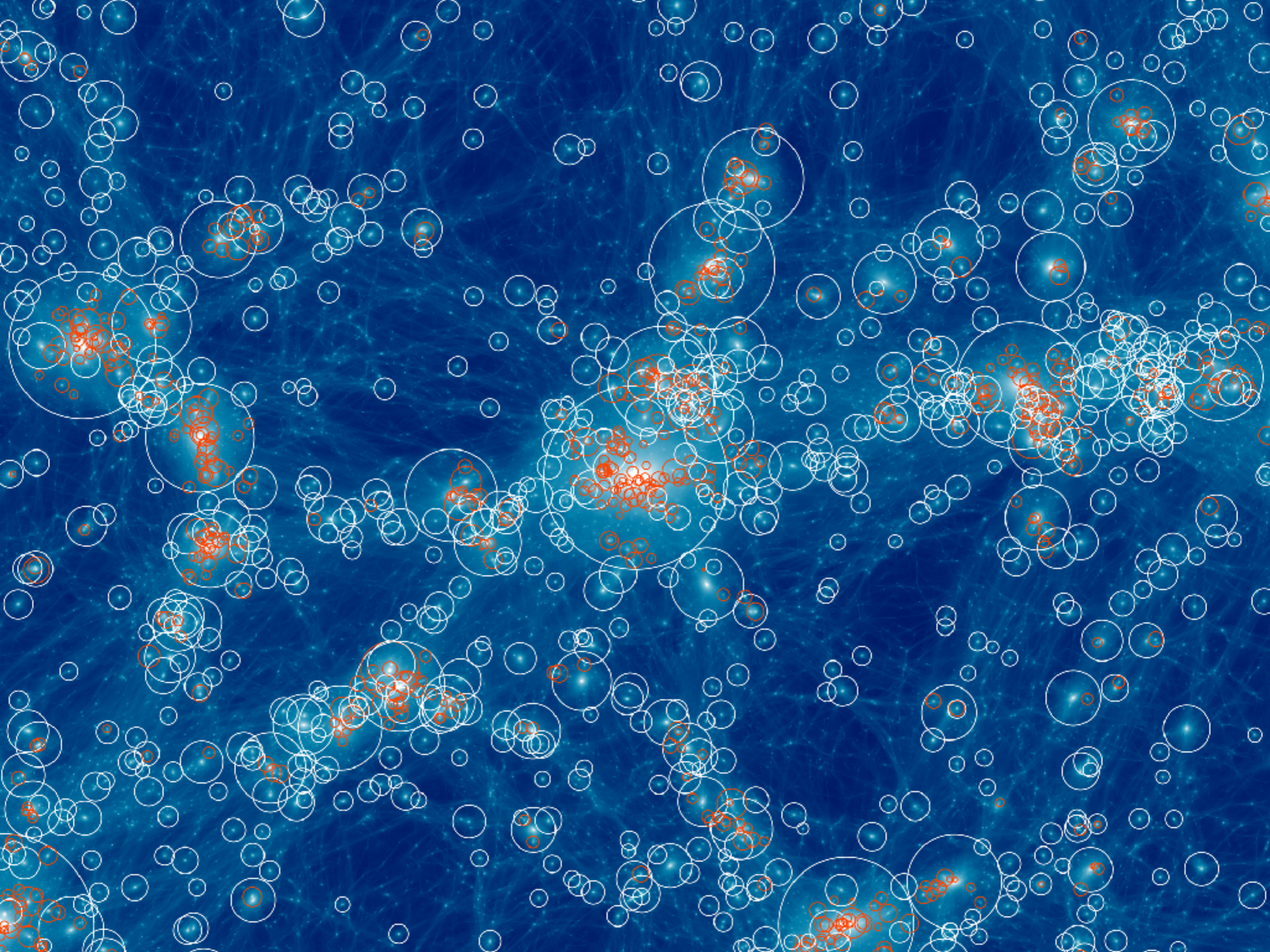
```
cosmology.setCosmology('bolshoi')  
c = hc.concentration(1E12, 'vir', 0.0,  
                    model = 'bullock01')
```



## Modules:

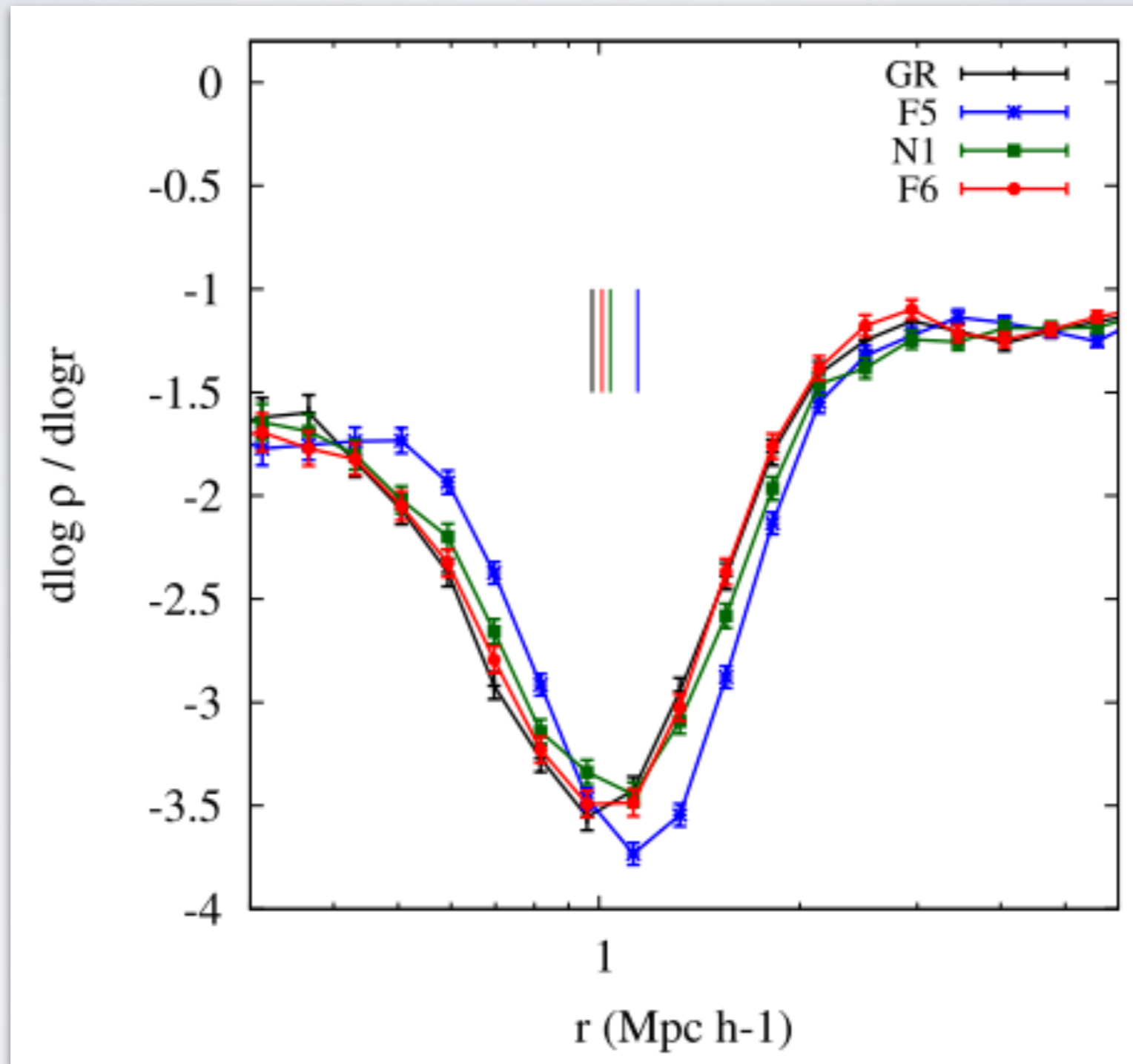
- Cosmology
- Power spectrum
- Gaussian random peaks
- Halo mass function
- Bias
- Density profiles
- Halo mass definitions
- Concentration
- Splashback
- ...



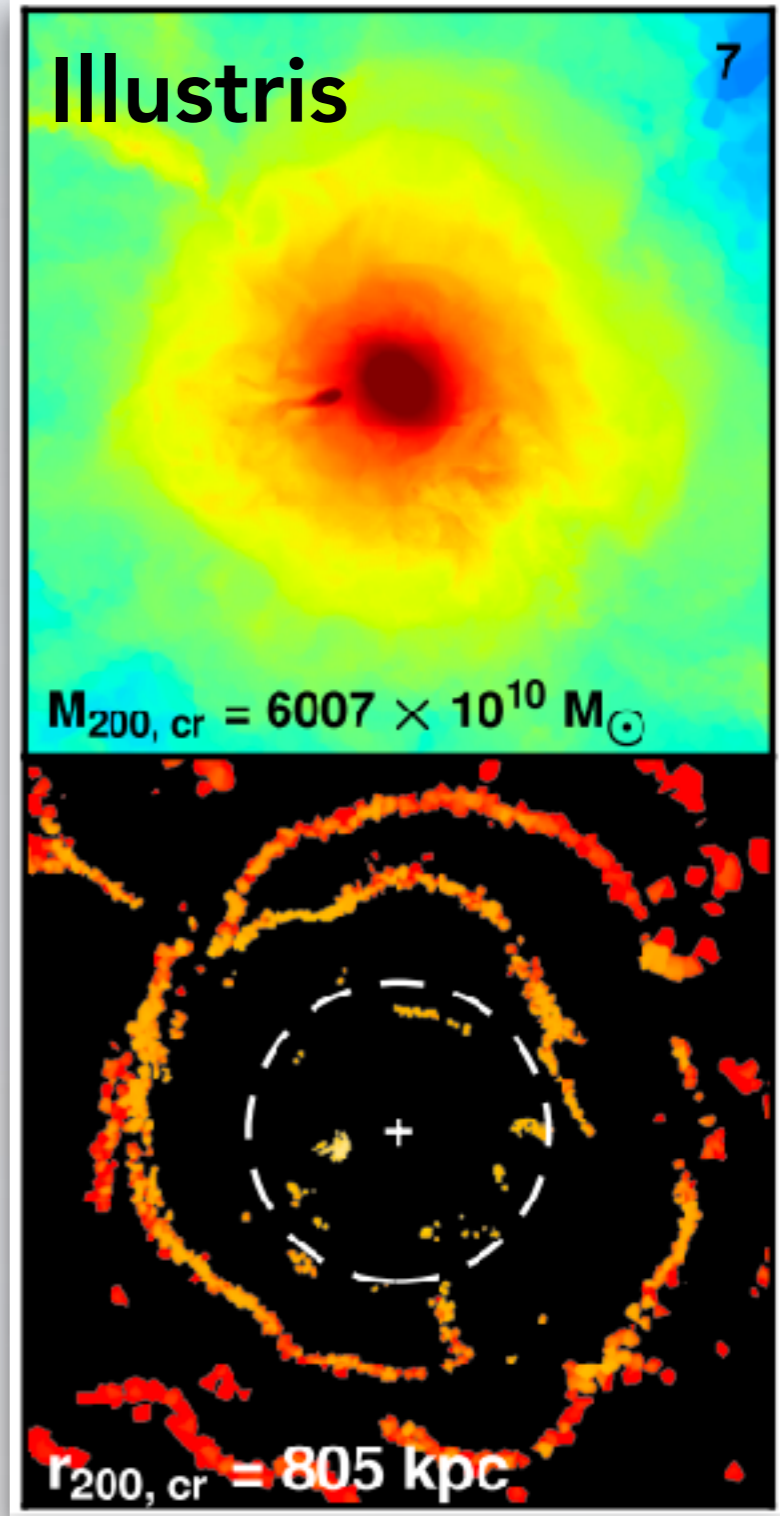
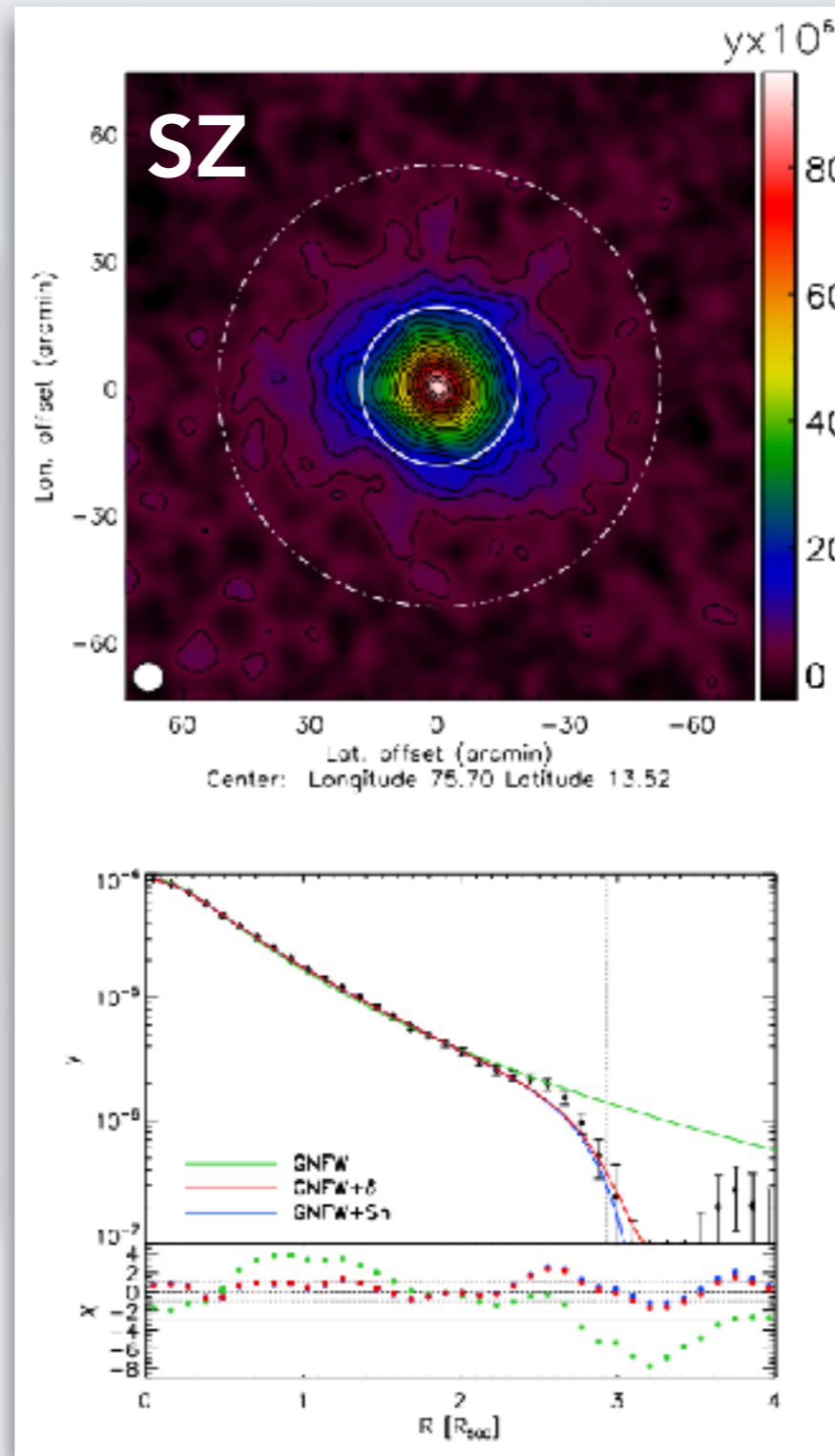
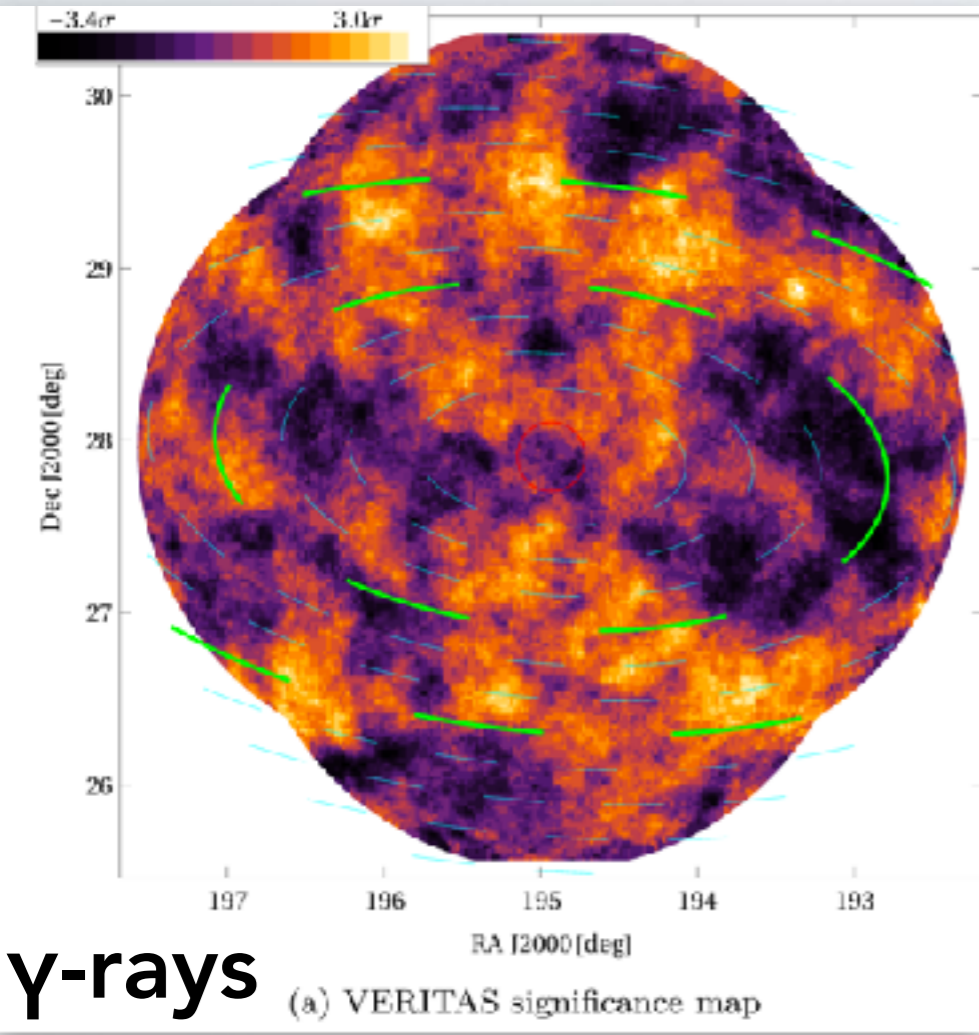




# Cosmology with the splashback radius?



# A splashback-accretion shock relation?









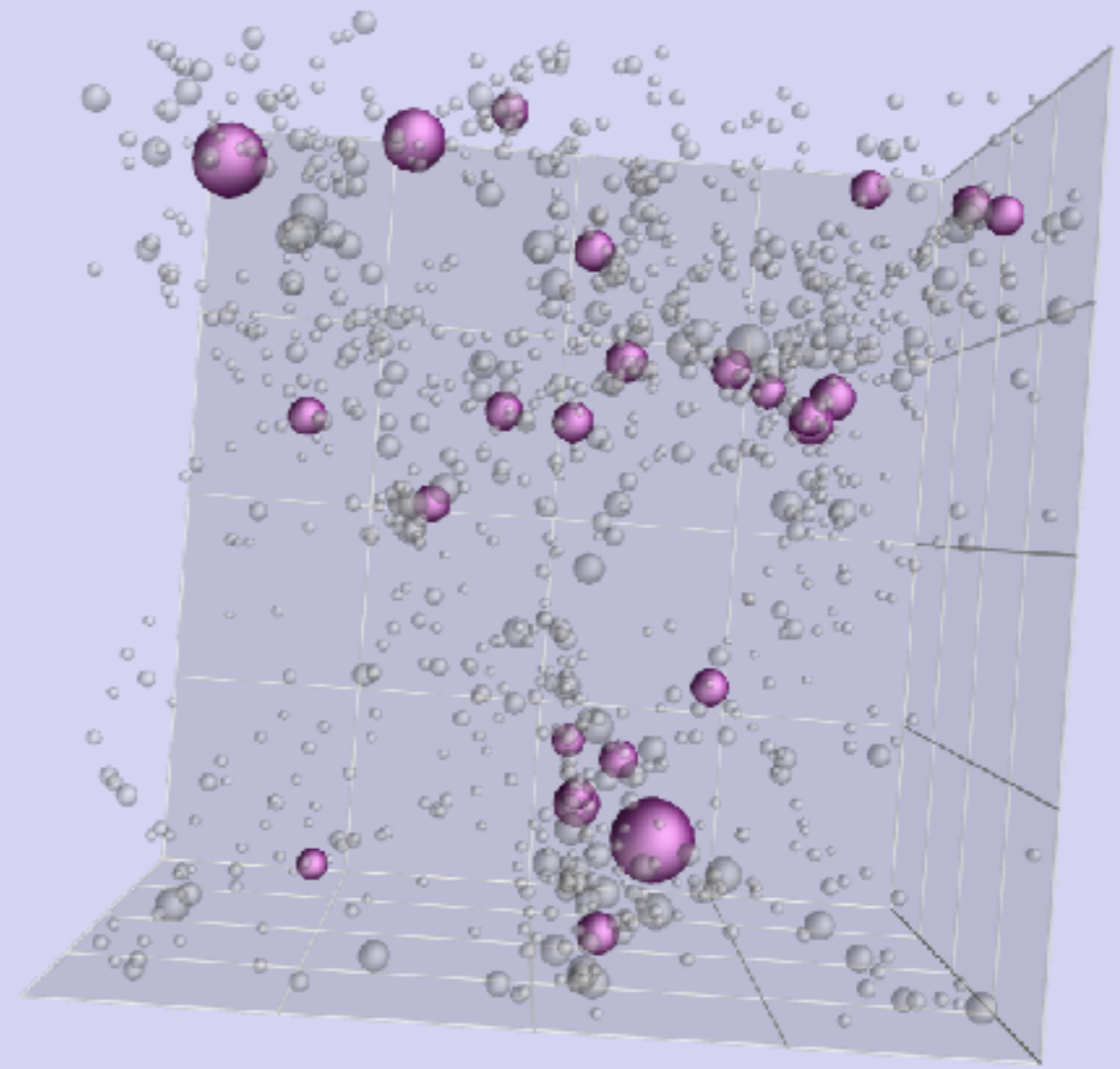






The Fabric of the Universe





The Fabric of the Universe



# Conclusions

- The **structure of CDM halos** is not a solved problem
- The **splashback radius** provides a physical halo boundary
- **SPARTA** provides entirely new ways of analyzing N-body simulations

Diemer & Kravtsov 2014 • ApJ 789, 001 • arXiv 1401.01216  
More, Diemer & Kravtsov 2015 • ApJ 810, 036 • arXiv 1504.05591  
Mansfield, Kravtsov & Diemer 2017 • ApJ 841, 034 • arXiv 1612.01531  
Diemer 2017 • ApJS 231, 05 • arXiv 1703.09712  
Diemer et al. 2017 • ApJ 843, 140 • arXiv 1703.09716