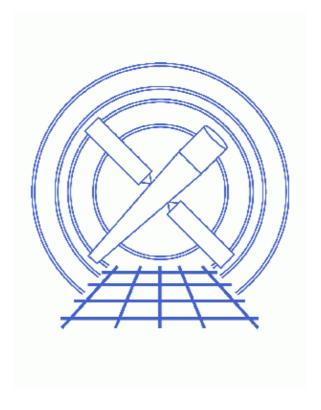
Simulation for Suzaku: Evaluate HXD systematic errors



Sherpa Threads (CIAO 3.4)

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Simulation for Suzaku: Evaluate HXD systematic

Last modified: 1 Dec 2006

Sherpa Threads

errors

Overview

Last Update: 1 Dec 2006 – updated for CIAO 3.4: removed "AO1" from thread title

Synopsis:

This thread describes how to simulate Suzaku/HXD (Hard X–ray Detector) observations and evaluate the HXD background using *Sherpa*. It is based on the <u>Cookbook for HXD Simulations</u> described by the HXD Team, but uses a somewhat different approach that leads to similar but not identical results. It is presented here to show how HXD simulations can be done with *Sherpa*, and to show how systematic errors can be treated within *Sherpa*.

Related Links:

• Related thread: <u>Simulating 1–D Data: the Sherpa FAKEIT Command</u>

Proceed to the <u>HTML</u> or hardcopy (PDF: <u>A4 | letter</u>) version of the thread.

Getting Started

The sensitivity of the HXD is dominated by the systematic error of the instrumental background estimates. Only objects brighter than the systematic error of the background estimate at the energy band in question are detectable. The HXD team strongly suggests that users check the effect of background systematic error (BGD-sys-err) in the simulation.

The PHA and response files used in this thread may be obtained from the <u>Cookbook for HXD Simulations</u>. The following files are needed:

```
ae_hxd_gso_20051019.rsp
ae_hxd_gsobkg_20051105.pha
ae_hxd_pinbkg_20051105.pha
ae_hxd_pinhxnom_20051104.rsp
```

Note that we use the HXD nominal postion from the file ae_hxd_pinhxnom_20051104.rsp in these simulations. If you choose to use the default XIS nominal position instead, you will need to replace that file with ae hxd pinxinom 20051104.rsp in your analysis.

Simulate the Model Spectra

In order to create a simulated dataset, it is necessary to define a source model and a grid over which to evaluate the model stack. We will use the datasets ae_hxd_pinbkg_20051105.pha and ae_hxd_gsobkg_20051105.pha to create the grid; the input datasets will be overwritten by the simulated data created by <u>FAKEIT</u>. Note that we read each file twice, first as a data set (with the "data" command) and then as a background ("back" command). After reading the data and the background files, we input the response files and set the instrument models for the two data sets.

```
sherpa> paramprompt off
Model parameter prompting is off
sherpa> sherpa.dataplot.x_log = 1
sherpa> sherpa.dataplot.y_log = 1
sherpa> data 1 ae_hxd_pinbkg_20051105.pha
The inferred file type is PHA. If this is not what you want, please
specify the type explicitly in the data command.
sherpa> <u>back</u> 1 ae_hxd_pinbkg_20051105.pha
The inferred file type is PHA. If this is not what you want, please
specify the type explicitly in the data command.
sherpa> data 2 ae_hxd_gsobkg_20051105.pha
The inferred file type is PHA. If this is not what you want, please
specify the type explicitly in the data command.
sherpa> back 2 ae_hxd_gsobkg_20051105.pha
The inferred file type is PHA. If this is not what you want, please
specify the type explicitly in the data command.
sherpa> <u>rsp</u>[pin]
sherpa> rsp[gso]
sherpa> pin.rmf = ae_hxd_pinhxnom_20051104.rsp
sherpa> gso.rmf = ae_hxd_gso_20051019.rsp
sherpa> <u>instrument</u> 1 = pin
sherpa> instrument 2 = gsc
```

Next, define the source model and set the inital value of various model parameters. Here we use a power law model multiplied by a constant:

```
sherpa> source 1,2 = constld[c1]*powlawld[p1]
sherpa> c1 integrate off
sherpa> c1.c0 = 0.1
sherpa> p1.gamma = 2.1
sherpa> p1.ref = 1.0
sherpa> p1.ampl.min = 0.0
sherpa> p1.ampl.max = 100.0
sherpa> p1.ampl = 9.7
sherpa> freeze p1.ampl
```

The simulation time is set to 1e5 seconds for each of the datasets before running fakeit to create the spectra:

```
sherpa> fakeit 1 time = 1e5
sherpa> fakeit 1
FAKEIT: The current background data have been added to the faked spectrum.
sherpa> fakeit 2 time = 1e5
sherpa> fakeit 2
FAKEIT: The current background data have been added to the faked spectrum.
```

Finally, restrict the energy range of each file to the default energy band of the <u>Hard X-ray Detector</u> (HXD):

```
sherpa> ignore 1 energy 60.0:
sherpa> ignore 2 energy :30.0,600.0:
```

Examine the HXD Background Level

Now we can examine the HXD background levels and plot the data and the background. First we subtract the background from the simulated PHA files, then overplot the data and background for both instruments. The color of the curves is changed to distinguish between them easily.

```
sherpa> subtract 1,2
sherpa> oplot data 1 data 2 back 1 back 2
Warning: negative and zero values ignored in log scale
sherpa> limits x 10 600
sherpa> limits y 0.001 1
sherpa> c 2 blue
sherpa> c 3 green
sherpa> c 4 red
sherpa> redraw
```

Figure 1 shows the resulting plot. Another filter is used to ignore the undetectable band:

```
sherpa> ignore 2 energy 250.0:
```

Effect of Background Systematic Error

We can evaluate the background systematic error using the backscale in *Sherpa*. First we unsubtract the data and set the backscale parameter to 1 for both simulated data sets and the background files. Then we fit the data to get the best fit parameters for the photon index and the constant, and obtain the confidence range for these parameters (projection). Note that the power law amplitude remains frozen at the simulated value.

```
sherpa> unsubtract 1,2
sherpa> setdata 1 backscale = 1.0
sherpa> setdata 2 backscale = 1.0
sherpa> setback 1 backscale = 1.0
sherpa> setback 2 backscale = 1.0
sherpa> subtract 1,2
sherpa> fit
LVMQT: V2.0
LVMQT: initial statistic value = 76.9445
LVMQT: final statistic value = 76.5725 at iteration 9
           c1.c0 0.0997023
           pl.gamma 2.09921
sherpa> projection
Projection complete for parameter: c1.c0
Projection complete for parameter: pl.gamma
Computed for sherpa.proj.sigma = 1
        Parameter Name Best-Fit Lower Bound Upper Bound
          c1.c0 0.0997023 -0.000915563 +0.000925169
p1.gamma 2.09921 -0.00321998 +0.00322589
sherpa> oplot data 1 data 2 back 1 back 2
Warning: negative and zero values ignored in log scale
sherpa> limits x 10 600
sherpa> limits y 0.001 1
```

```
sherpa> c 2 blue
sherpa> c 3 green
sherpa> c 4 red
sherpa> redraw
```

Figure 2 shows the results with all backscales set to 1.0. The simulated values of photon index and constant parameters are recovered with a 1 sigma lower and upper bound as indicated by projection. This range needs to be compared to the systematic background error, which is 10%.

Next the background is overestimated by setting the background backscales to 0.9:

```
sherpa> unsubtract 1,2
sherpa> setback 1 backscale = 0.9
sherpa> setback 2 backscale = 0.9
sherpa> subtract 1,2
sherpa> fit
LVMQT: V2.0
LVMQT: initial statistic value = 6215.34
LVMQT: final statistic value = 4071.34 at iteration 9
          c1.c0 0.144673
          pl.gamma 2.24062
sherpa> projection
Projection complete for parameter: c1.c0
Projection complete for parameter: pl.gamma
Computed for sherpa.proj.sigma = 1
       Parameter Name
                        Best-Fit Lower Bound Upper Bound
          sherpa> oplot data 1 data 2 back 1 back 2
Warning: negative and zero values ignored in log scale
sherpa> limits x 10 600
sherpa> limits y 0.001 1
sherpa> c 2 blue
sherpa> c 3 green
sherpa> c 4 red
sherpa> redraw
```

Figure 3 shows the results with background backscales set to 0.9.

Finally we underestimate the background by setting the background backscales to 1.1:

```
sherpa> unsubtract 1,2
sherpa> setback 1 backscale = 1.1
sherpa> setback 2 backscale = 1.1
sherpa> subtract 1,2
sherpa> fit
LVMOT: V2.0
LVMQT: initial statistic value = 11579.9
LVMQT: final statistic value = 2763.36 at iteration 12
           c1.c0 0.0679399
           pl.gamma 1.95657
sherpa> projection
Projection complete for parameter: c1.c0
Projection complete for parameter: pl.gamma
Computed for sherpa.proj.sigma = 1
                          Best-Fit Lower Bound
       Parameter Name
                                                   Upper Bound
```

```
c1.c0 0.0679399 -0.000576908 +0.000579968 p1.gamma 1.95657 -0.00290537 +0.00289905

sherpa> oplot data 1 data 2 back 1 back 2

Warning: negative and zero values ignored in log scale sherpa> limits x 10 600 sherpa> limits y 0.001 1 sherpa> c 2 blue sherpa> c 3 green sherpa> c 4 red sherpa> redraw
```

Figure 4 shows the results with background backscales set to 1.1.

The three p1.gamma values obtained in this thread for different background backscale indicate a range of the systematic error in the simulation. The systematic error is larger than the statistical error (found by projection in each case). We assumed p1.gamma=2.1 and c1.c0=0.1 in the simulation. The systematic error in this simulation indicates p1.gamma values between 1.95–2.24 and c1.c0 values between 0.06–0.14.

Save the Results

First we write out the simulated datasets. The <u>save</u> and <u>write mdl</u> commands are then used to record the rest of the session information.

```
sherpa> write data fakel.pha PHA
sherpa> write data 2 fake2.pha PHA
sherpa> save all suzaku.shp
sherpa> write mdl "suzaku_mdl.fits"
```

suzaku.shp is text file with all the information needed to restore this session. suzaku_mdl.fits (the MDL file) stores the data, model, and other information in FITS format.

Commands Used in this thread

The text file <u>suzaku commands.txt</u> contains all the commands issued in this thread:

```
paramprompt off
sherpa.dataplot.x_log = 1
sherpa.dataplot.y_log = 1
data 1 ae_hxd_pinbkg_20051105.pha
back 1 ae_hxd_pinbkg_20051105.pha
data 2 ae_hxd_gsobkg_20051105.pha
back 2 ae_hxd_gsobkg_20051105.pha
rsp[pin]
rsp[gso]
pin.rmf = ae_hxd_pinhxnom_20051104.rsp
gso.rmf = ae_hxd_gso_20051019.rsp
instrument 1 = pin
instrument 2 = gso
source 1,2 = const1d[c1]*powlaw1d[p1]
cl integrate off
c1.c0 = 0.1
```

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```
p1.gamma = 2.1
p1.ref = 1.0
p1.ampl.min = 0.0
p1.ampl.max = 100.0
p1.ampl = 9.7
freeze pl.ampl
fakeit 1 time = 1e5
fakeit 1
fakeit 2 \text{ time} = 1e5
fakeit 2
ignore 1 energy 60.0:
ignore 2 energy :30.0,600.0:
subtract 1,2
oplot data 1 data 2 back 1 back 2
limits x 10 600
limits y 0.001 1
c 2 blue
c 3 green
c 4 red
redraw
ignore 2 energy 250.0:
unsubtract 1,2
setdata 1 backscale = 1.0
setdata 2 backscale = 1.0
setback 1 backscale = 1.0
setback 2 backscale = 1.0
subtract 1,2
fit
projection
oplot data 1 data 2 back 1 back 2
limits x 10 600
limits y 0.001 1
c 2 blue
c 3 green
c 4 red
redraw
unsubtract 1,2
setback 1 backscale = 0.9
setback 2 backscale = 0.9
subtract 1,2
fit
projection
oplot data 1 data 2 back 1 back 2
limits x 10 600
limits y 0.001 1
c 2 blue
c 3 green
c 4 red
redraw
unsubtract 1,2
setback 1 backscale = 1.1
setback 2 backscale = 1.1
subtract 1,2
fit
projection
oplot data 1 data 2 back 1 back 2
limits x 10 600
limits y 0.001 1
c 2 blue
```

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```
c 3 green
c 4 red
redraw
write data fakel.pha PHA
write data 2 fake2.pha PHA
save all suzaku.shp
write mdl "suzaku_mdl.fits"
```

History

14 Jan 2005 reviewed for CIAO 3.2: no changes

14 Dec 2005 updated for CIAO 3.3: Astro–E2 is now named Suzaku; new HXD data files, fits and plots updated accordingly

Last modified: 1 Dec 2006

01 Dec 2006 updated for CIAO 3.4: removed "AO1" from thread title

URL: http://cxc.harvard.edu/sherpa/threads/astro e2/

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Image 1: Background-subtracted data and backgrounds

This plot shows the two simulated datasets (with the background subtracted) and the background from the original PHA files.

- white: dataset 1blue: dataset 2
- green: background 1 (ae_hxd_pinbkg_20051105.pha)
- red: background 2 (ae_hxd_gsobkg_20051105.pha)

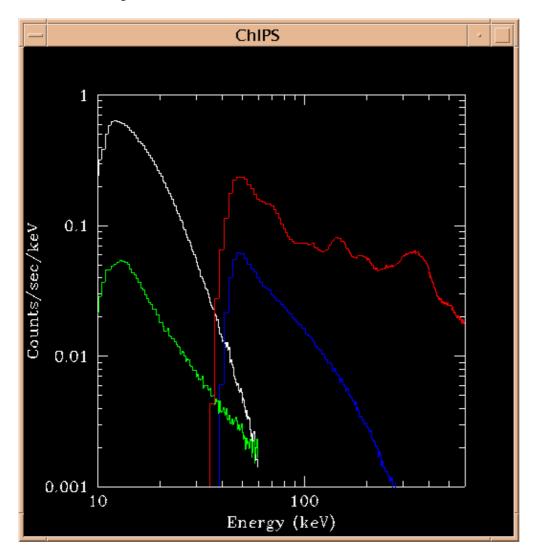


Image 2: Source and background backscales = 1.0

white: dataset 1 blue: dataset 2

• green: background 1 (ae_hxd_pinbkg_20051105.pha)

• red: background 2 (ae_hxd_gsobkg_20051105.pha)

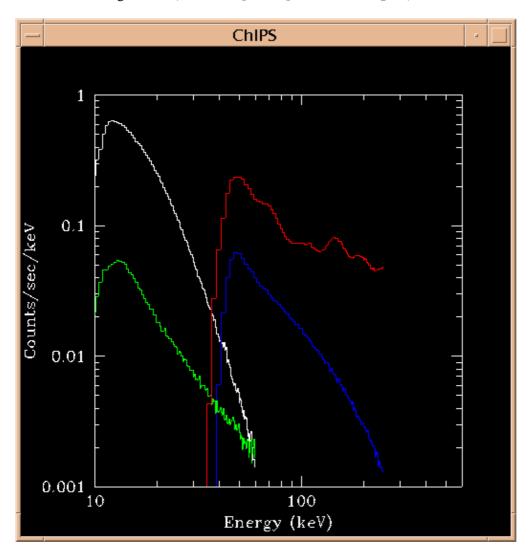


Image 3: Overestimate background: backscales = 0.9

- white: dataset 1blue: dataset 2
- green: background 1 (ae_hxd_pinbkg_20051105.pha)
- red: background 2 (ae_hxd_gsobkg_20051105.pha)

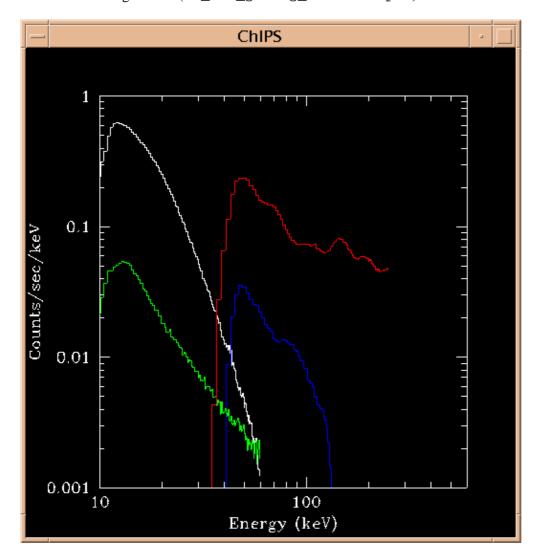


Image 4: Underestimate background: backscales = 1.1

- white: dataset 1 blue: dataset 2
- green: background 1 (ae_hxd_pinbkg_20051105.pha)
- red: background 2 (ae_hxd_gsobkg_20051105.pha)

